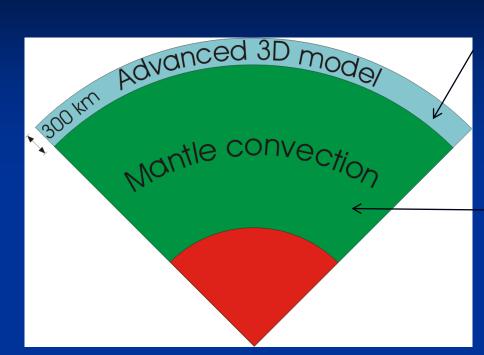
Remind from previous lectures

Coupling mantle convection and lithospheric deformation

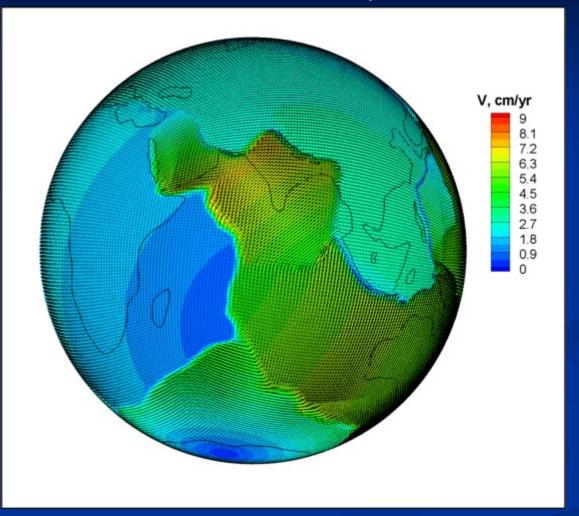
Lithospheric code (Finite Elements)



Mantle code (spectral or FEM)

Mantle and lithospheric codes are coupled through continuity of velocities and tractions at 300 km.

Friction at boundaries 0.02 (Smax= 30 MPa)



about right magnitudes of velocities

How to make friction so low?

$$\tau = c + \mu \cdot (\sigma_n - P_f)$$

 $\tau = c + \mu_{eff} \cdot \sigma_n$ $\mu_{eff} = \mu \cdot (1 - P_f / \sigma_n)$

Assume $\mu = 0.6$, $P_f = 0.95\sigma_n$ then $\mu_{eff} = 0.03$

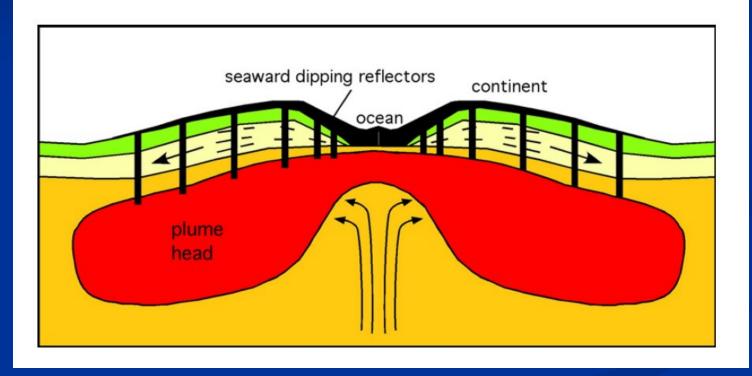
How to make friction so low?



$\frac{\text{then } \mu_{eff} = 0.03}{\text{Subducting slabs are aquaplaning deep}}$

Problems of classical plume model

I.H. Campbell / Chemical Geology 241 (2007) 153-176

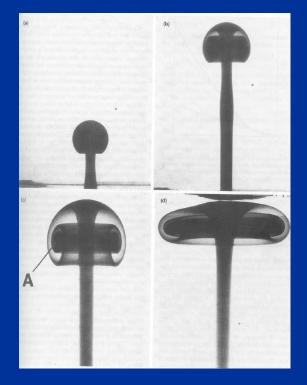


Prediction: Surface uplift = 0.7-1.0 km/100°, i.e. 1.4-3 km for DT= 200-300°

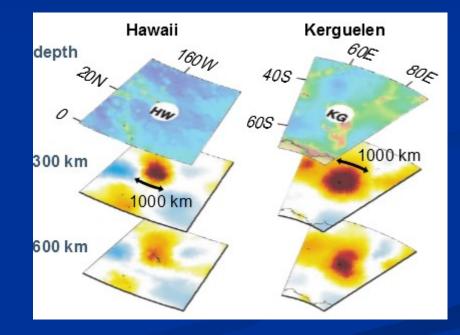
Observation: Often less than 1 km or even not detectable surface uplift

Problems of classical plume model

Prediction: narrow (R=100km) plume conduits (tails)



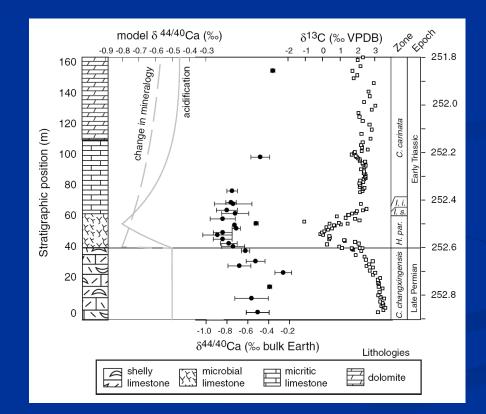
Seismic observations: wide (R=500km) plumes



From Montelli et al., 2006

Problems of classical plume model

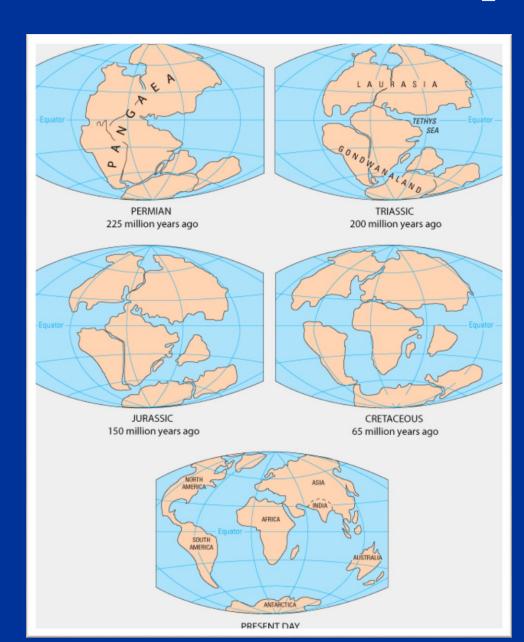
Volumes and isotopic composition of gases expected from eruptions above plume heads are not sufficient to explain observations for mass extinctions



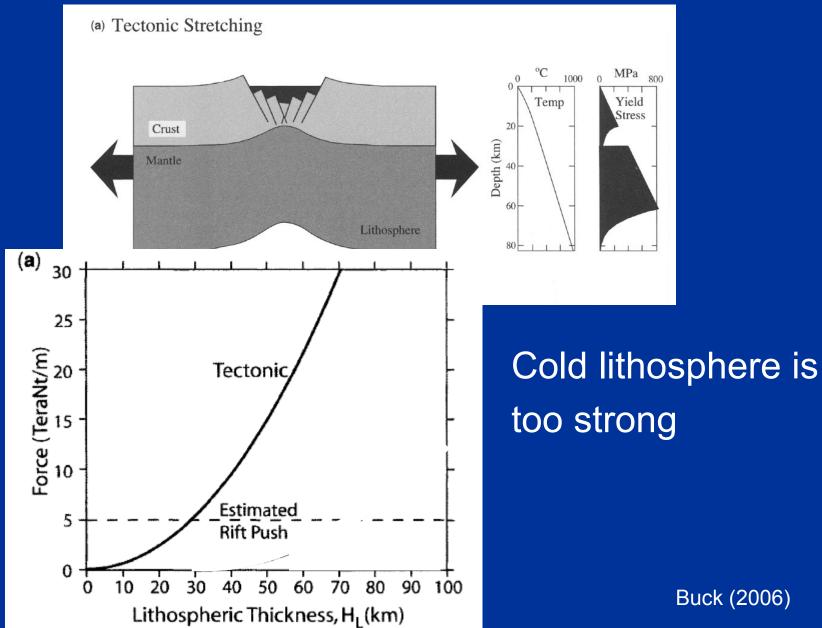
Lecture 5. Rifting, Continental break-up, Transform faults

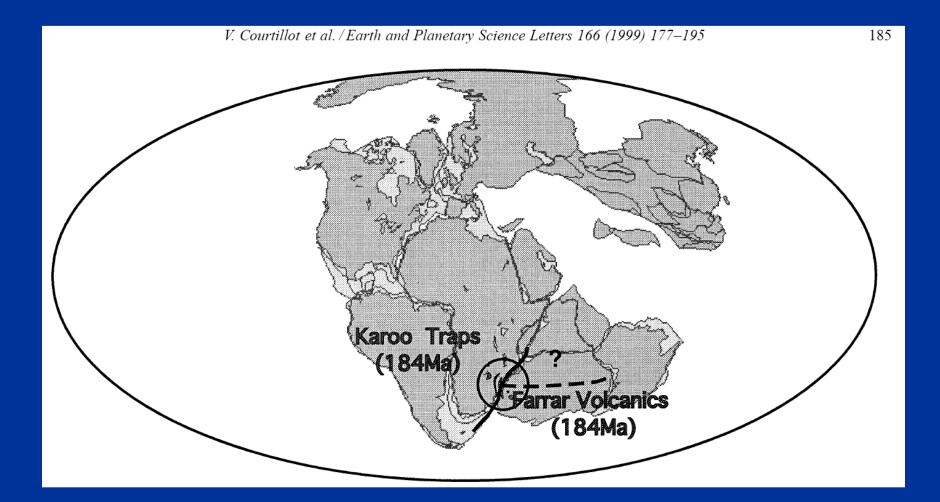
How to break a continent?

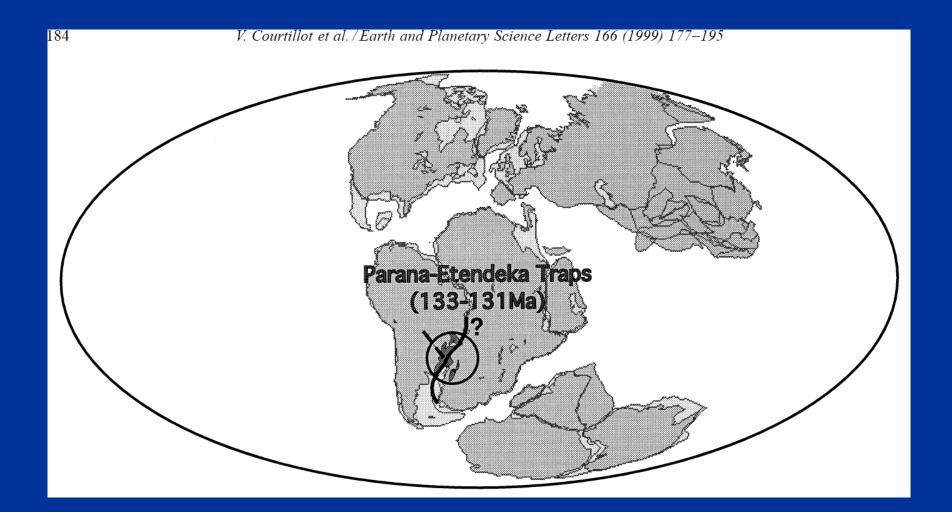
- Effect of magmas and Large Igneous Provinces
- Effect of oblique rifting
- Continental transform faults
 - > What caused Dead Sea transform?
 - San Andreas Fault System

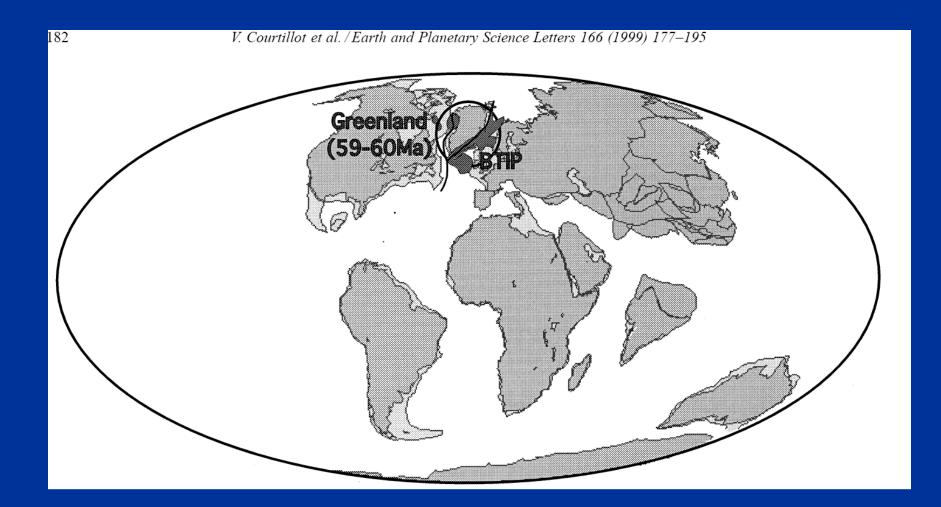


How to break continent?

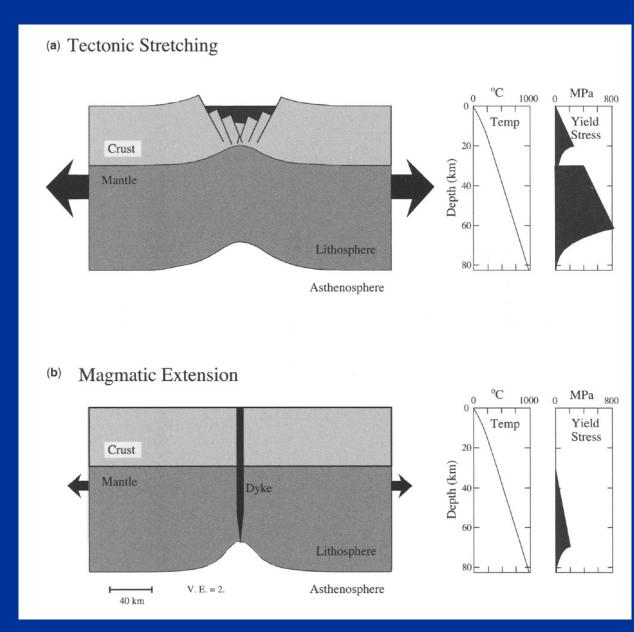






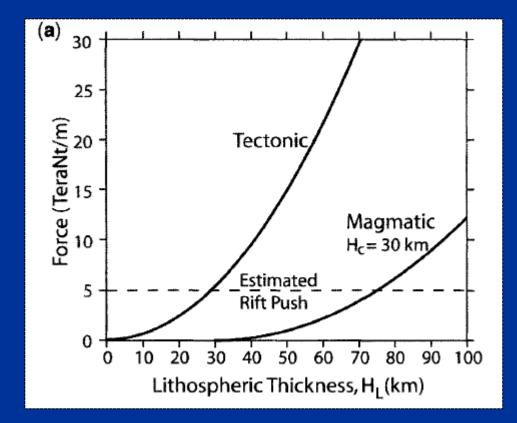


Effect of magma-filled dikes



Buck (2006)

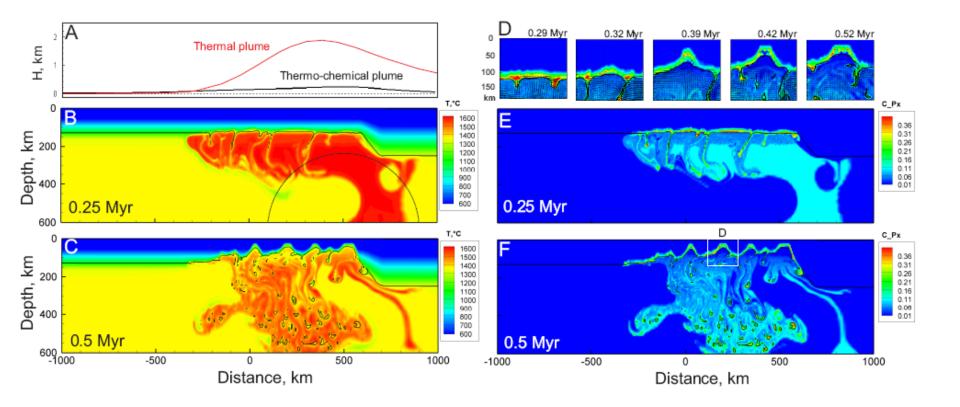
Effect of magma-filled dikes



It works if lithosphere is first thinned to about 75 km

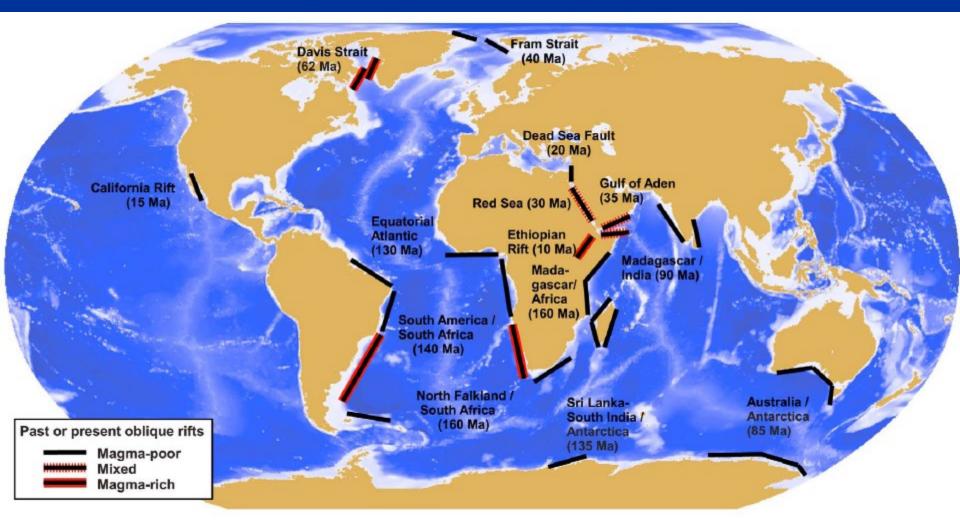
Buck (2006)

Lithospheric thinning above mantle plume



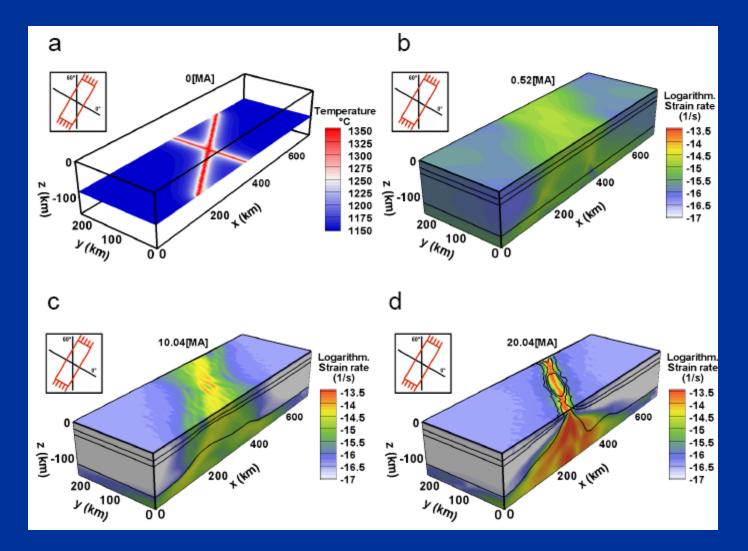
Sobolev et al. Nature 2011

Effect of oblique rifting



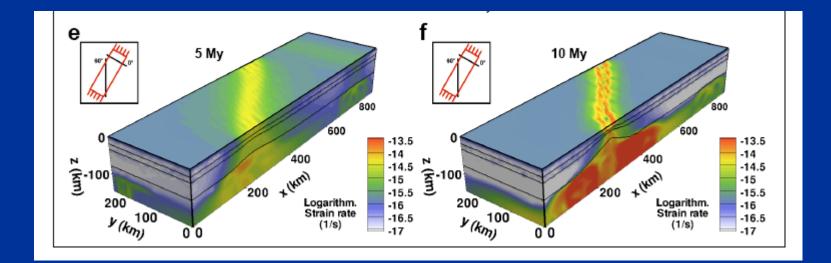
Brune, Popov, Sobolev JGR 2012

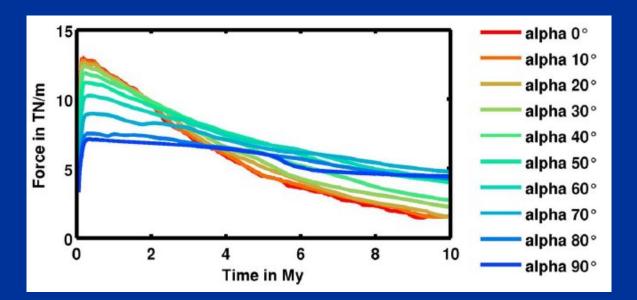
Effect of oblique rifting



Brune, Popov, Sobolev JGR 2012

Effect of oblique rifting

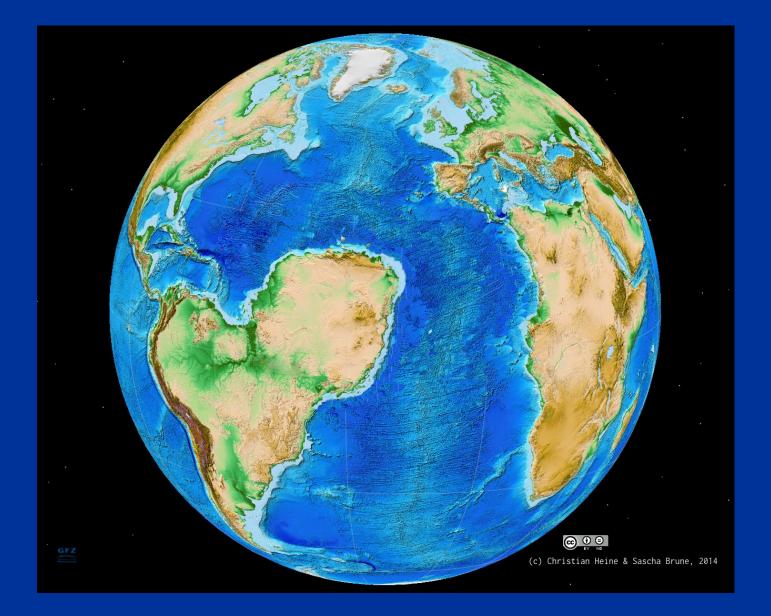




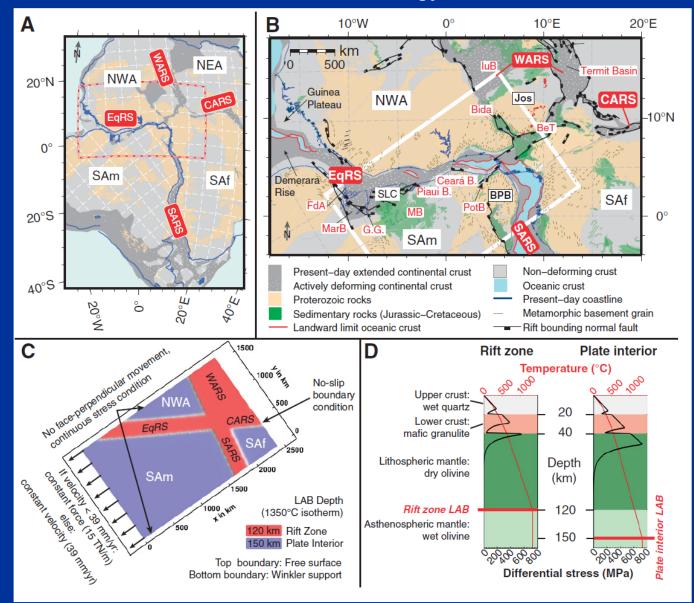
$$F_{strike-slip} = \tau_{yield} L_z$$

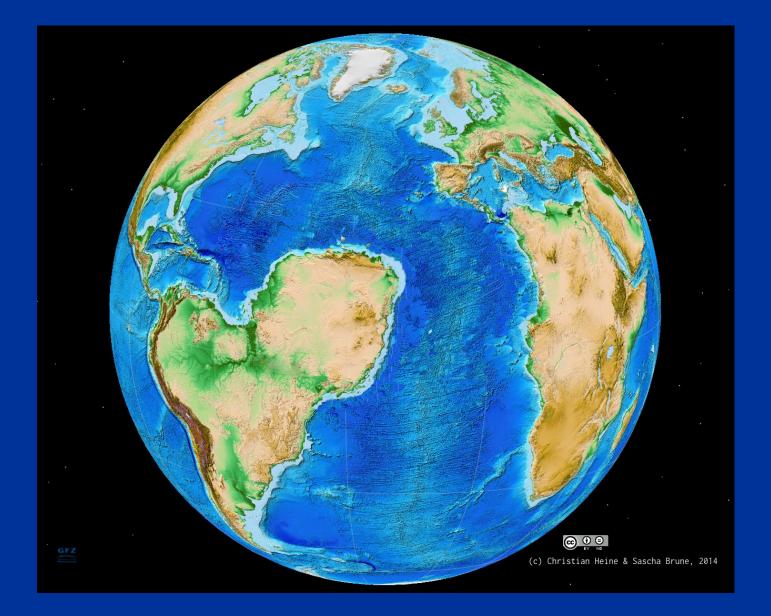
$$F_{extension} = \frac{\tau_{yield}L_z}{\sqrt{\frac{1}{3}(\nu^2 - \nu + 1)}}.$$

 $F_{extension} = 2\tau_{yield}L_z.$

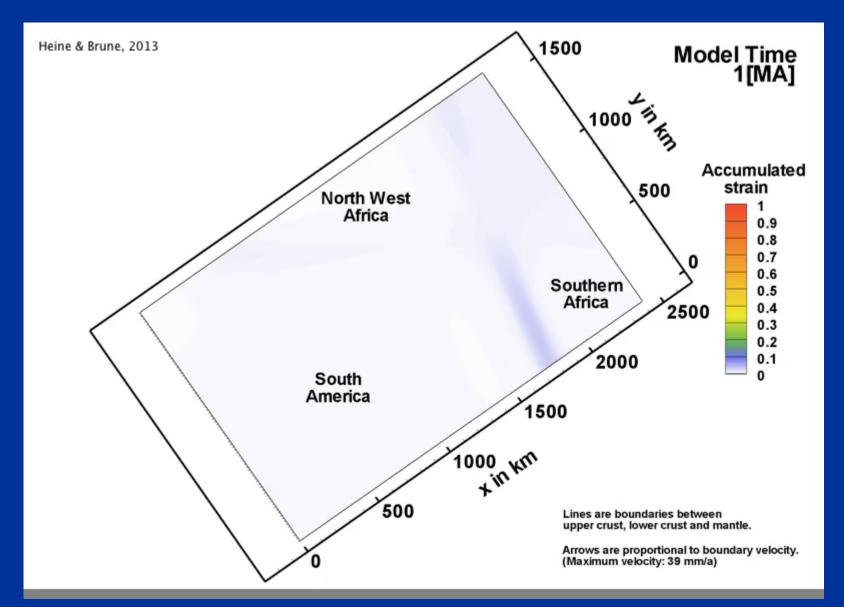


Heine and Brune, Geology, 2014

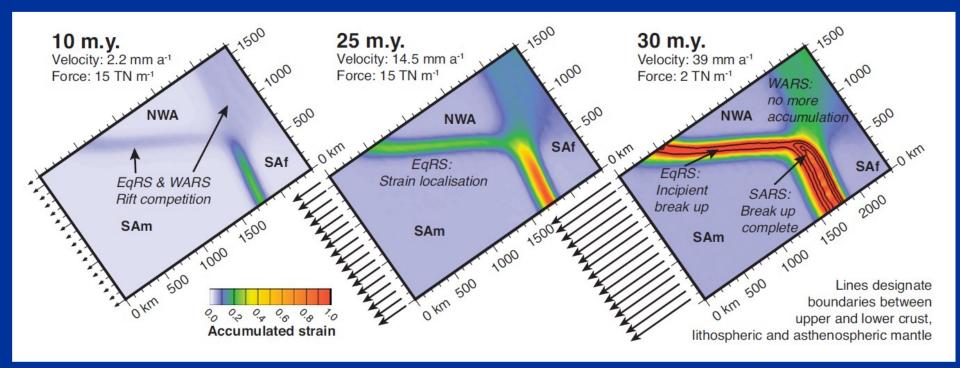




Heine and Brune, Geology, 2014



Heine and Brune, Geology, 2014



Conclusion

To break a continent are required:

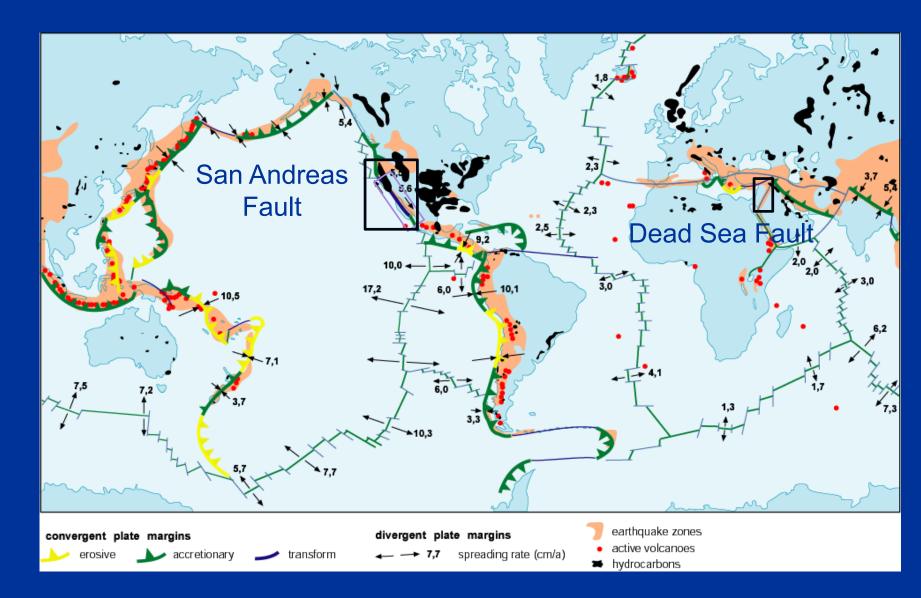
(1) extensional deviatoric stresses (internal, from ridge push or subduction zones roll-back) and (2) lithospheric weakening

Large Igneous Provinces are optimal for lithospheric weakening, as they may both thin lithosphere and generate magma-filled dikes.

Intensive strike-slip deformation is also helpful

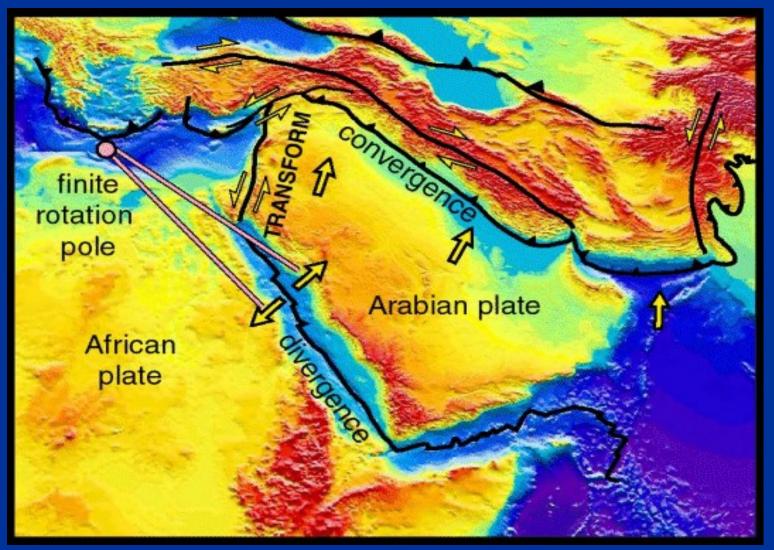
Continental transform faults (case Dead Sea Transform)

Continental Transform Faults



Regional setting

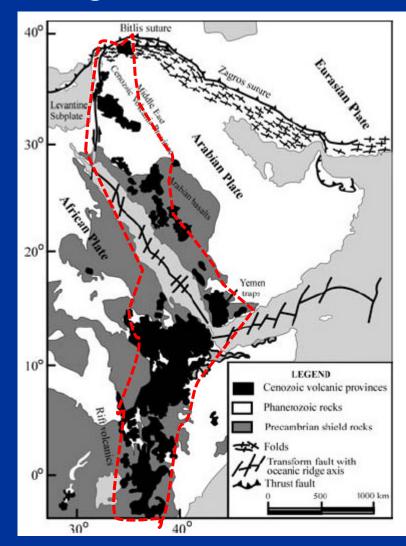
With the surface heat flow of 50-60 mW/m2, the DST is the coldest continental transform boundary



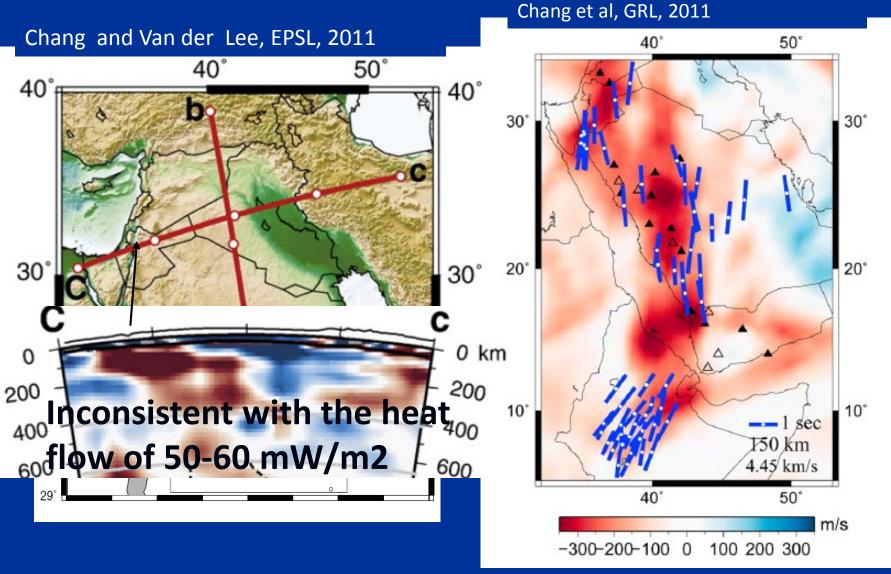
Lithospheric thickness and magmatism

30°N 25° RS AP 20° LAB Depth (km) 15° 80 120 160

Magmatism at 30-0 Ma



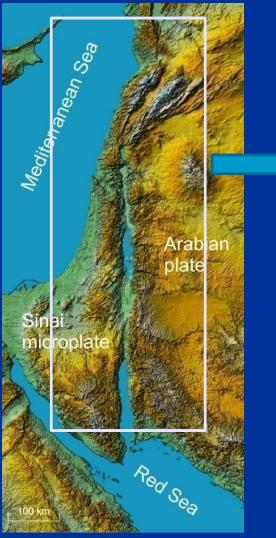
Lithosphere-asthenosphere boundary (LAB) from seismic data

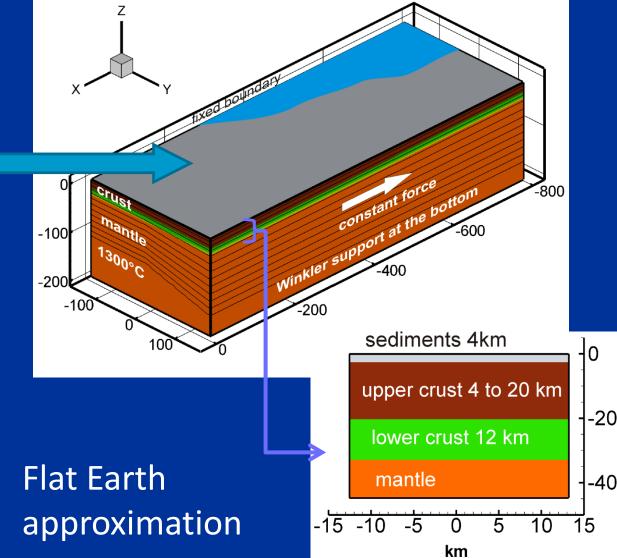


Conclusion

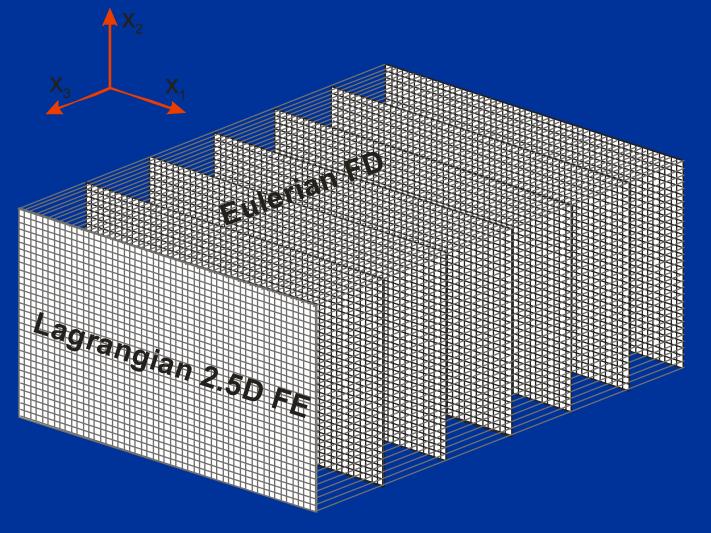
Lithosphere around DST was thinned in the past and related high heat flow had not enough time to reach the surface

Model setup

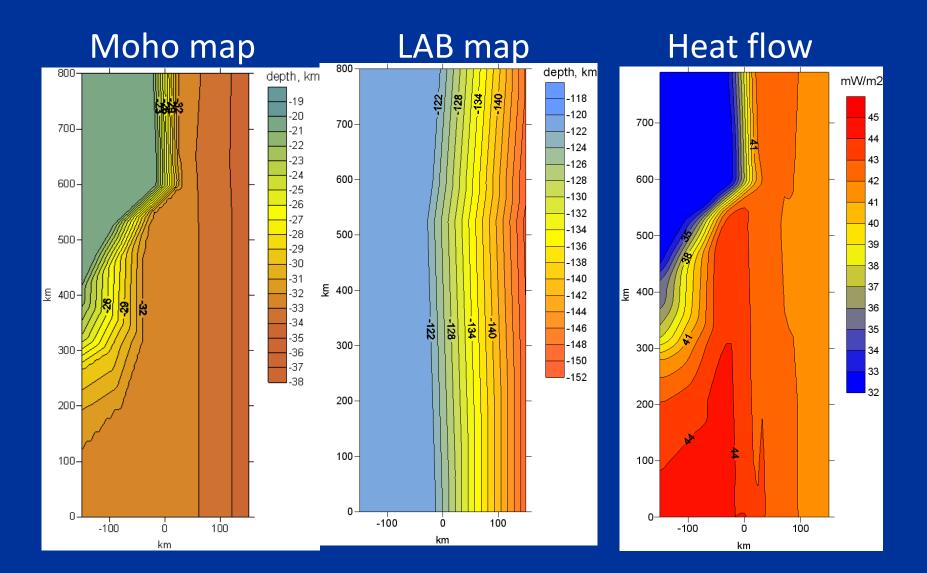


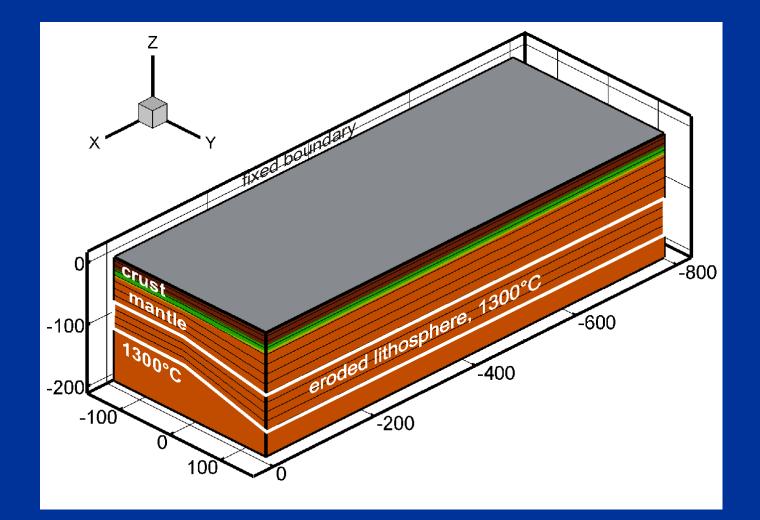


Modeling technique LAPEX 3D combining FE and FD (Petrunin and Sobolev, Geology, 2006, PEPI, 2008)

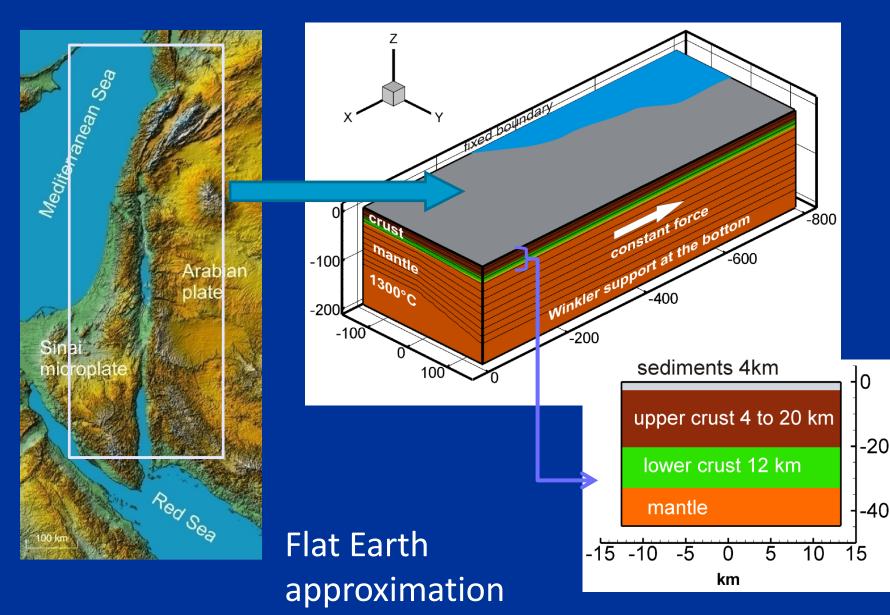


Initial lithospheric structure:

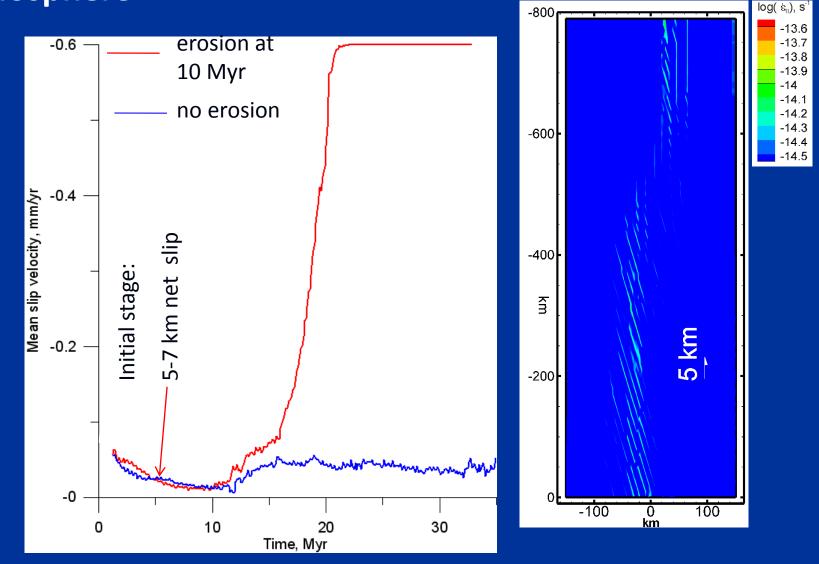


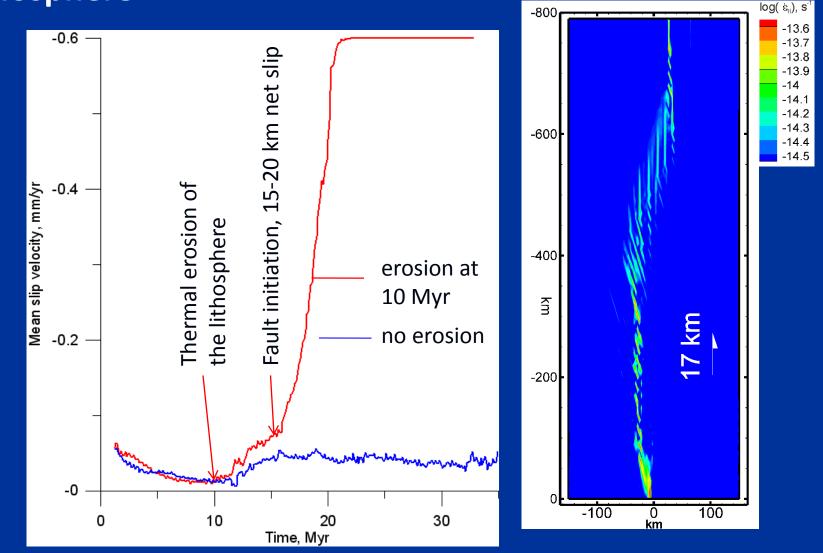


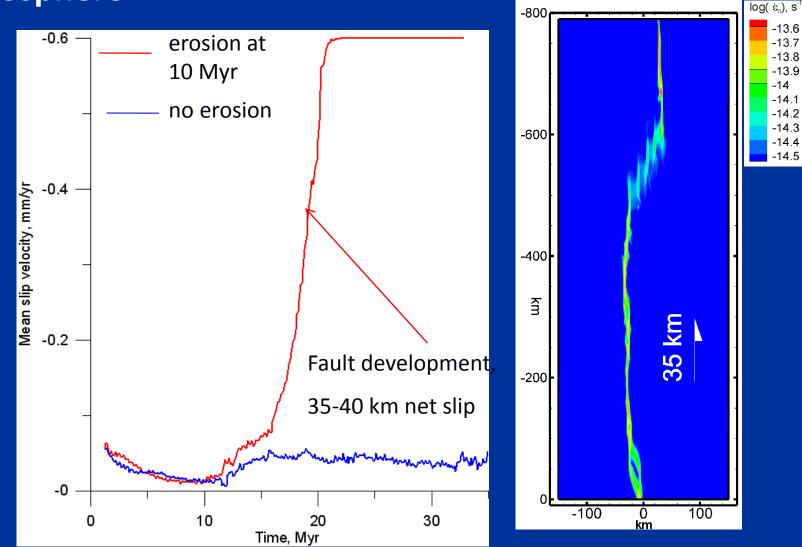
Model setup

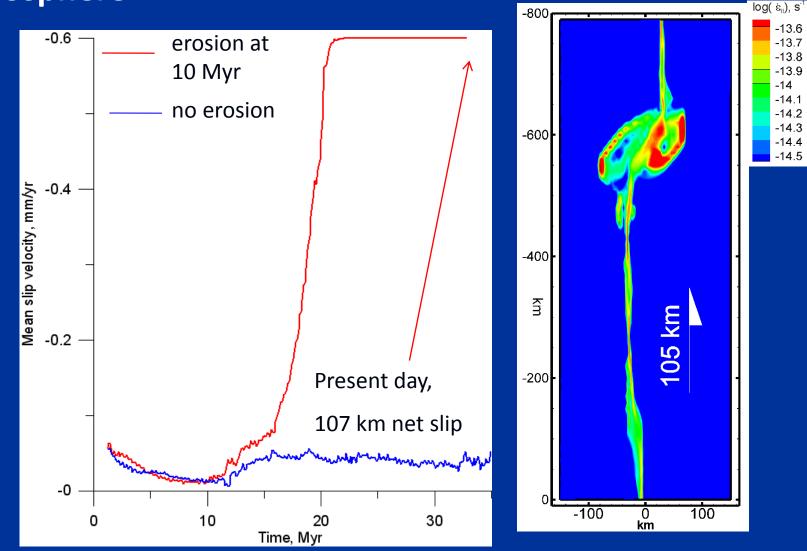


Modeling results: role of the thermal erosion of the Applied force is 1.6e13N/m

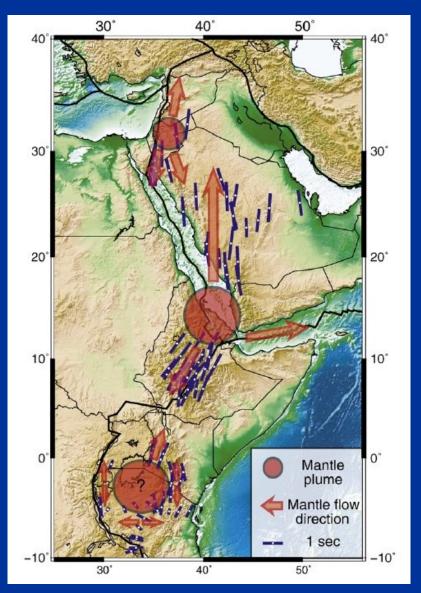








Possible scenario



Plumes at 25-35 Ma

Lithospheric erosion 20-30 Ma

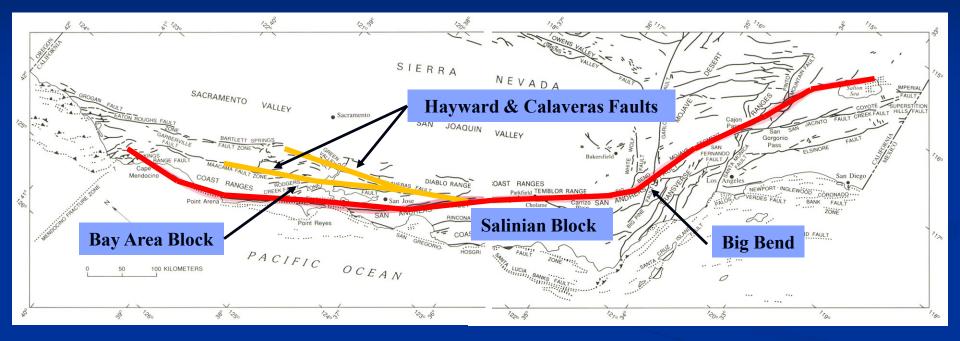
Localization of the DST 15-17 Ma

Lithospheric erosion has triggered the DST

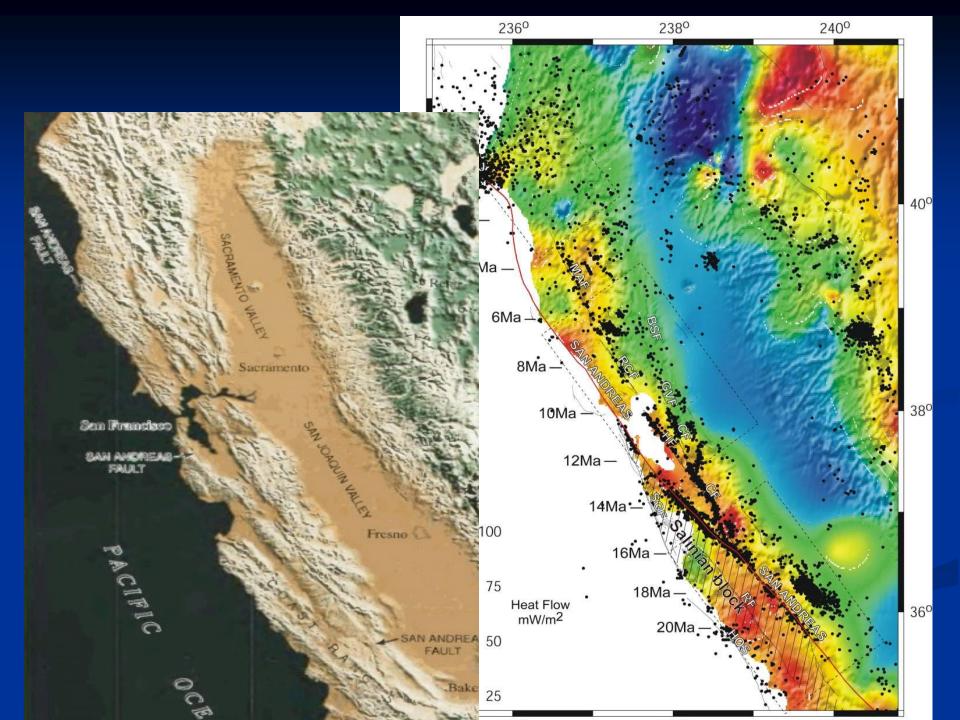
Chang and Van der Lee, EPSL, 2011

San Andreas Fault System

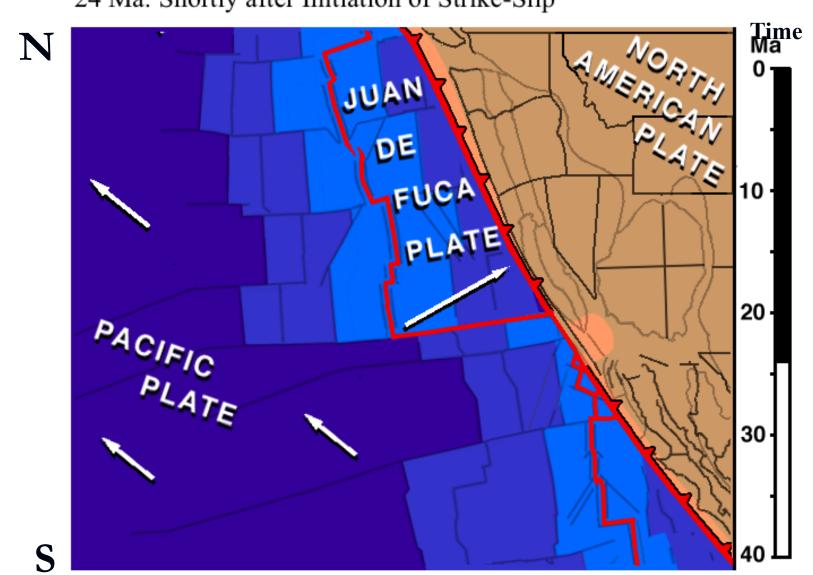
San Andreas Fault System



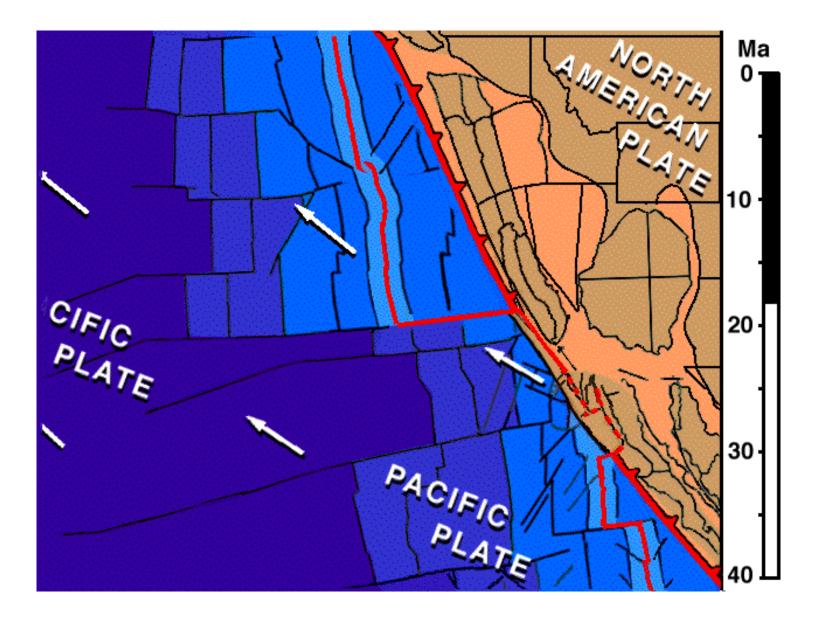
USGS Professional Paper 1515

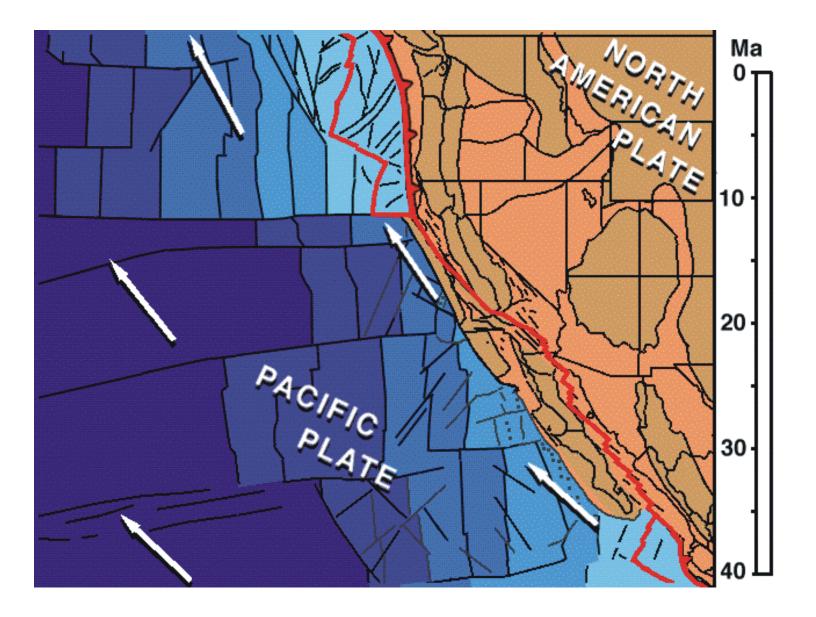


24 Ma: Shortly after Initiation of Strike-Slip



(animation by T. Atwater)



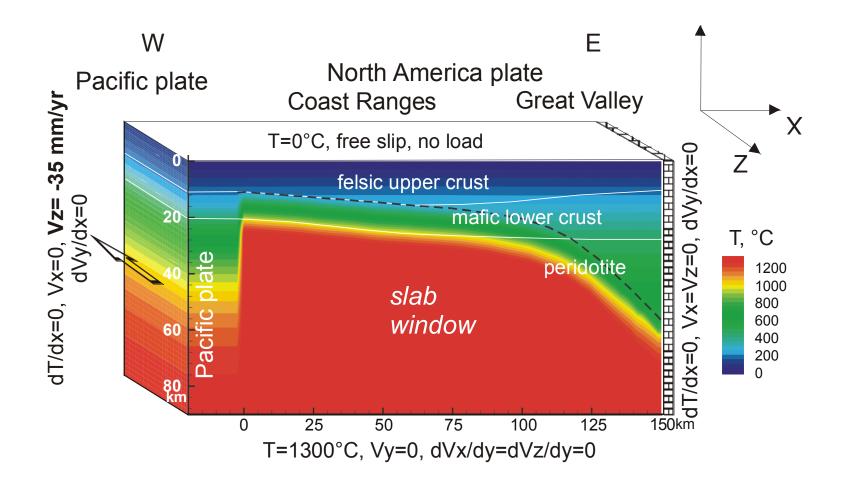


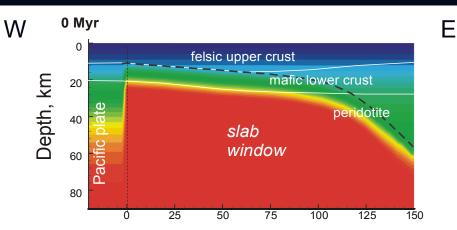
Questions addressed

Why the locus of deformation in SAFS migrates landwards with time?

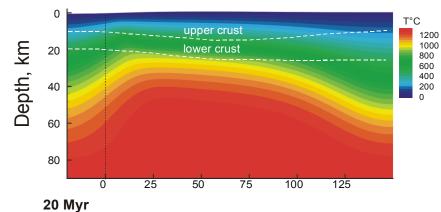
How differently would evolve SAFS with "strong" and "weak" major faults? Why the locus of deformation in SAFS migrates landwards with time?

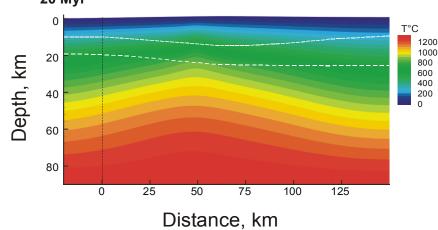
Extended 2D Model Setup (South view)

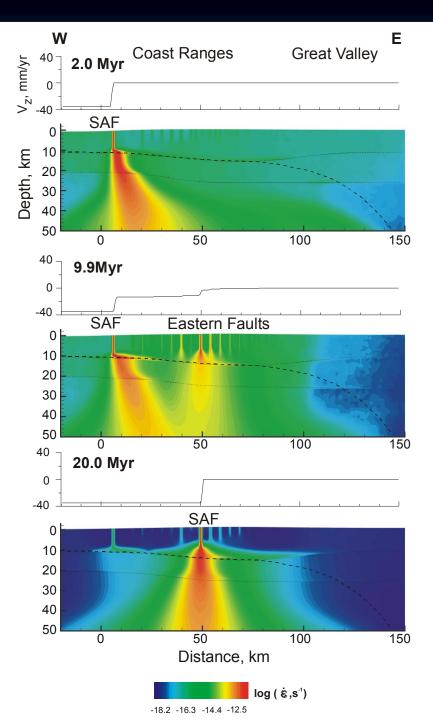




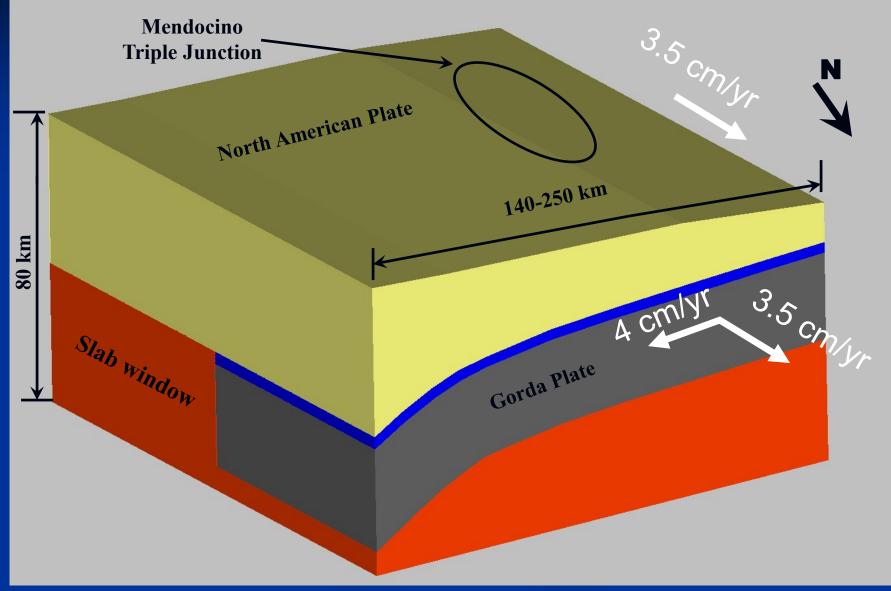








Model Setup (view from the North)



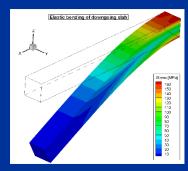
Popov, Sobolev, Zoback, G3 2012

Physical background

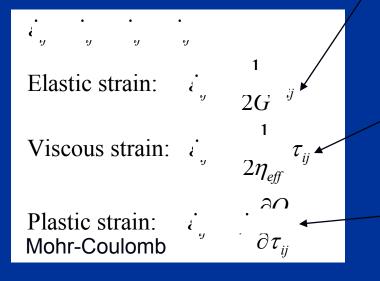
Balance equations

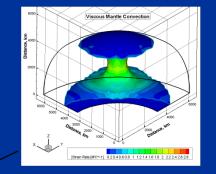
Momentum:
$$\frac{\partial \sigma_{ij}}{\partial x_j} + \Delta \rho g z_i = 0$$

Energy: $\frac{DU}{Dt} = -\frac{\partial q_i}{\partial x_i} + r$



Deformation mechanisms





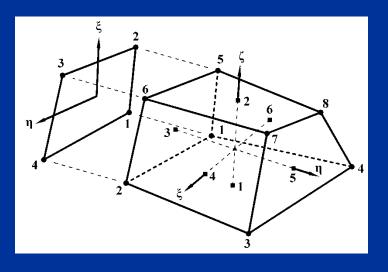
Plastic strain localization (thrust fault) φ = 30° ψ = 0° Arthur's angle: 0 = 45° - (φ + ψ)4 = 37.5° 38° Plastic Strain: 0.050 0.150 0.250 0.350 0.450 0.550

Popov and Sobolev (2008)

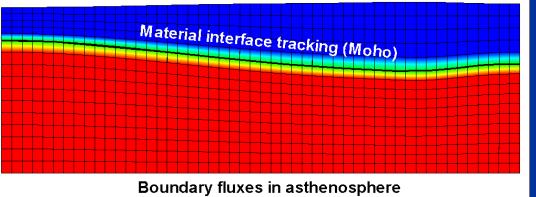
Numerical background

Discretization by Finite Element Method

Arbitrary Lagrangian-Eulerian kinematical formulation



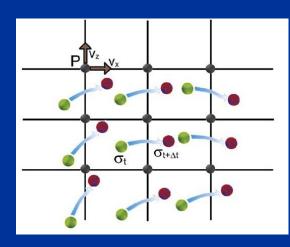
Free surface effects (erosion, sedimentation)



Fast implicit time stepping + Newton-Raphson solver

 $u_{k+1} = u_k - K_k^{-1} r_k$ r - Residual Vector $K = \frac{\partial r}{\partial \Delta u} - Tangent Matrix$

Remapping of entire fields by Particle-In-Cell technique



Popov and Sobolev (2008)

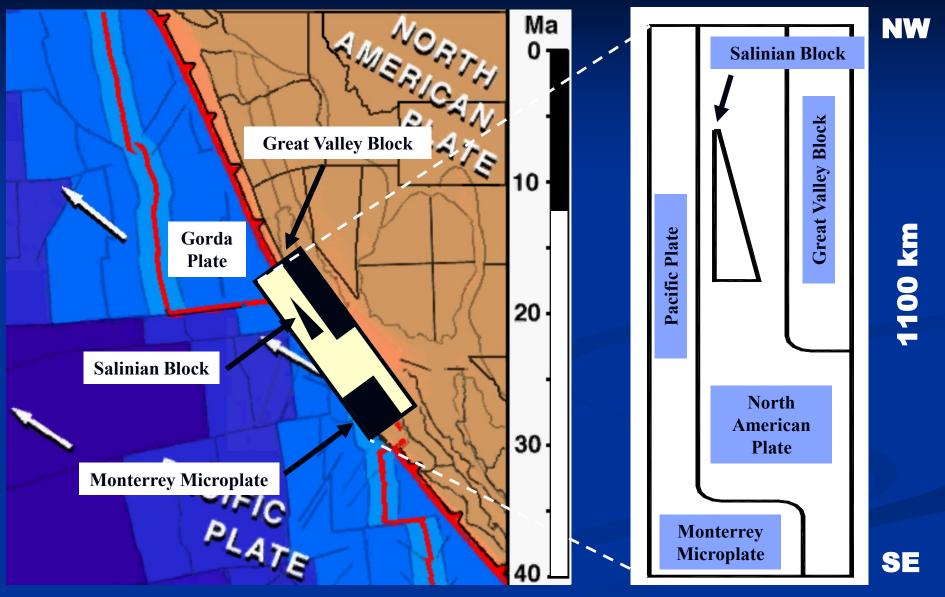
"Strong" and "weak" faults models

"Strong faults" model: the friction coefficient decreases only slightly (from 0.6 to 0.3) with increasing plastic strain

"Weak faults" model: the friction coefficient decreases drastically (from 0.6 to 0.07) with increasing plastic strain

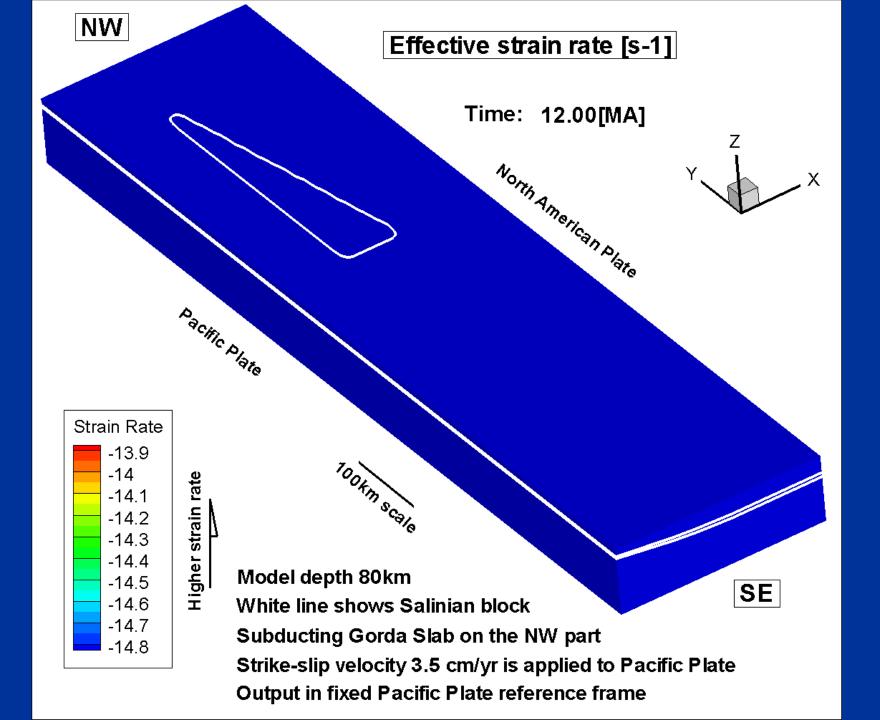
Popov, Sobolev, Zoback, G3 2012

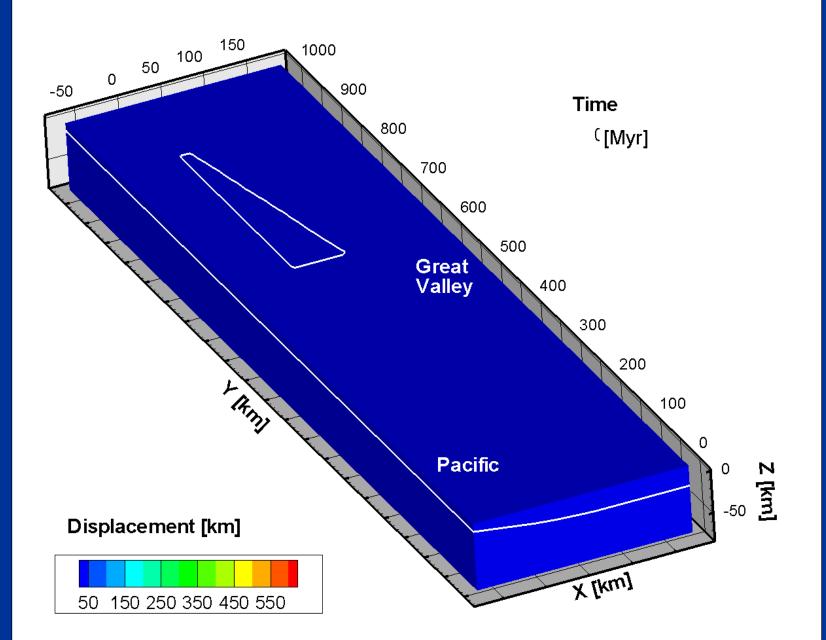
3D Model Setup (12-15 MA)



Popov, Sobolev, Zoback, G3 2012

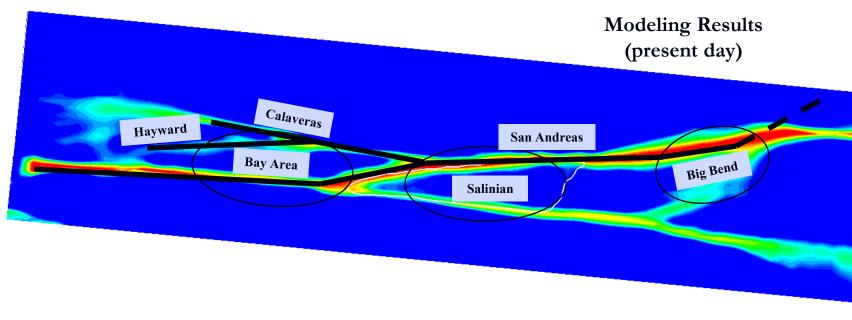
250km

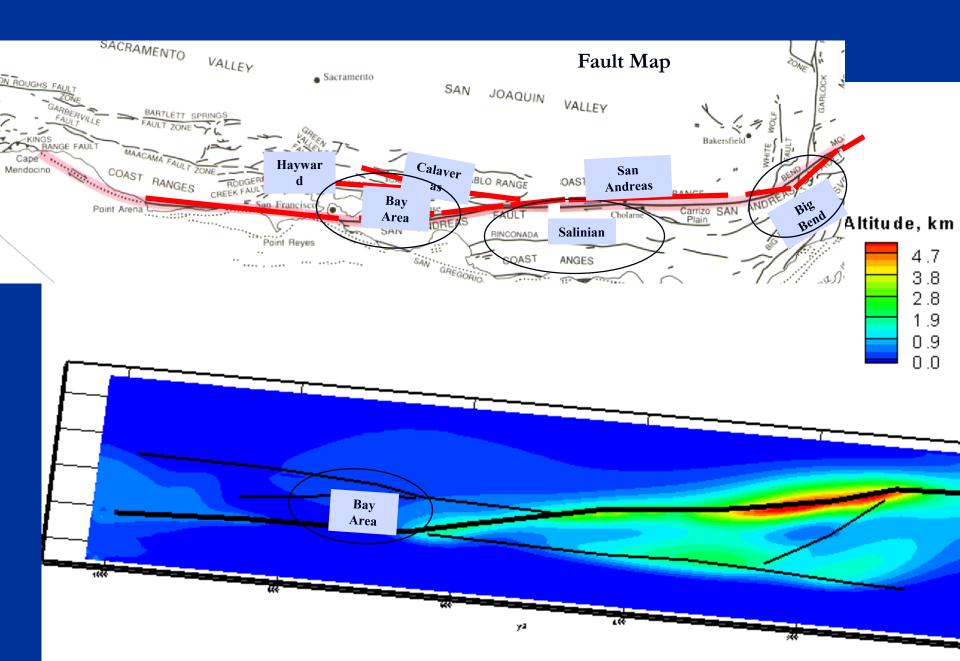




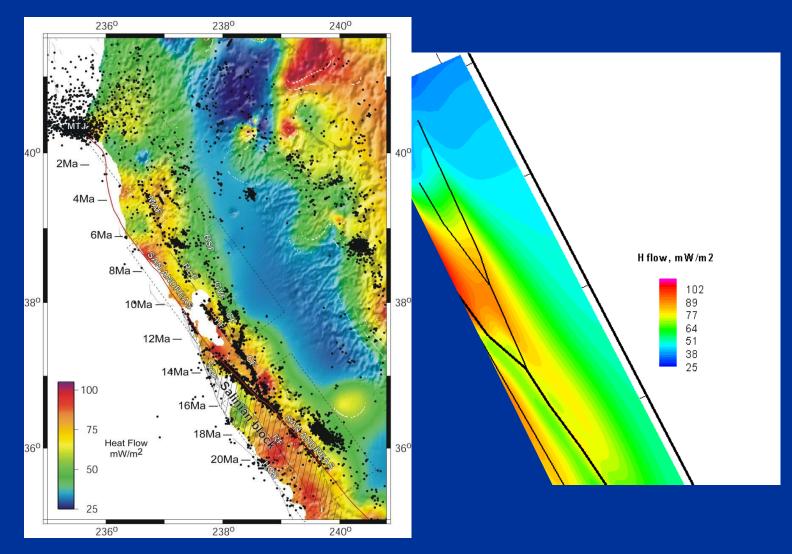
Qualitative comparison of basic fault features







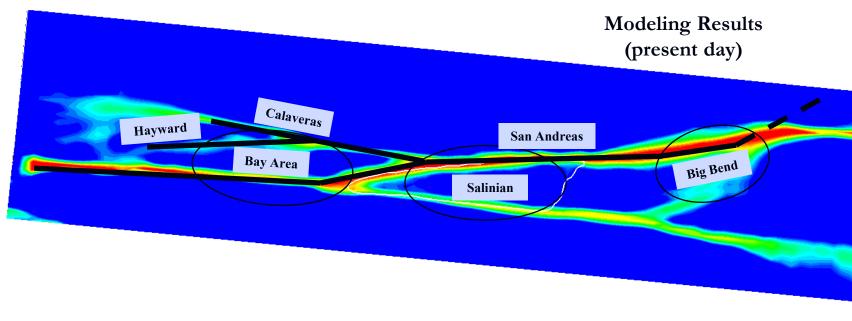
Modeled surface heat flow



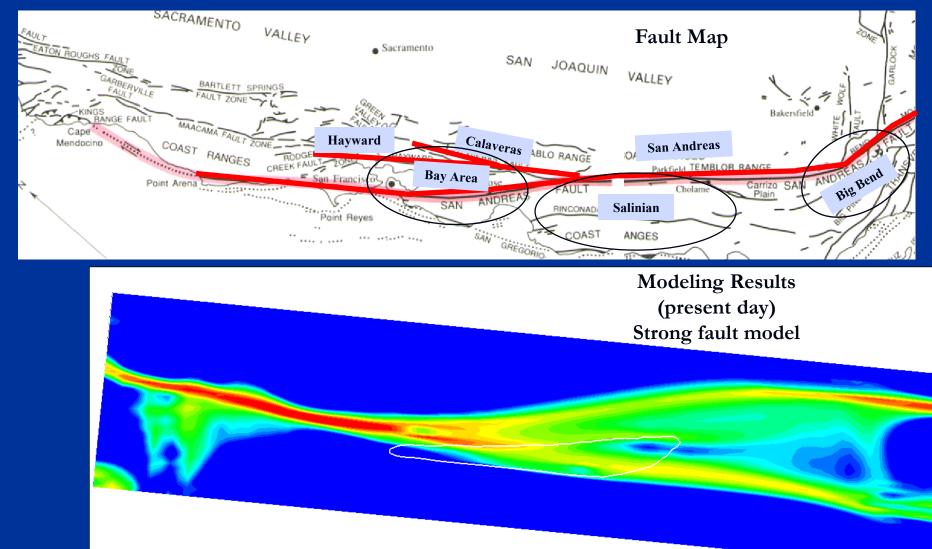
"Weak faults" versus "strong faults" model

Weak-faults model





Strong-faults model



Major faults in SAFS must be weak!

Conclusions for SAFS

Present day structure and landward motion of SAFS is controlled by kinematic boundary conditions and lithospheric heterogeneity, including captured Monterray microplate

Major faults at SAFS must be "week"