

Lecture 5. Rifting, Continental break-up, Transform faults

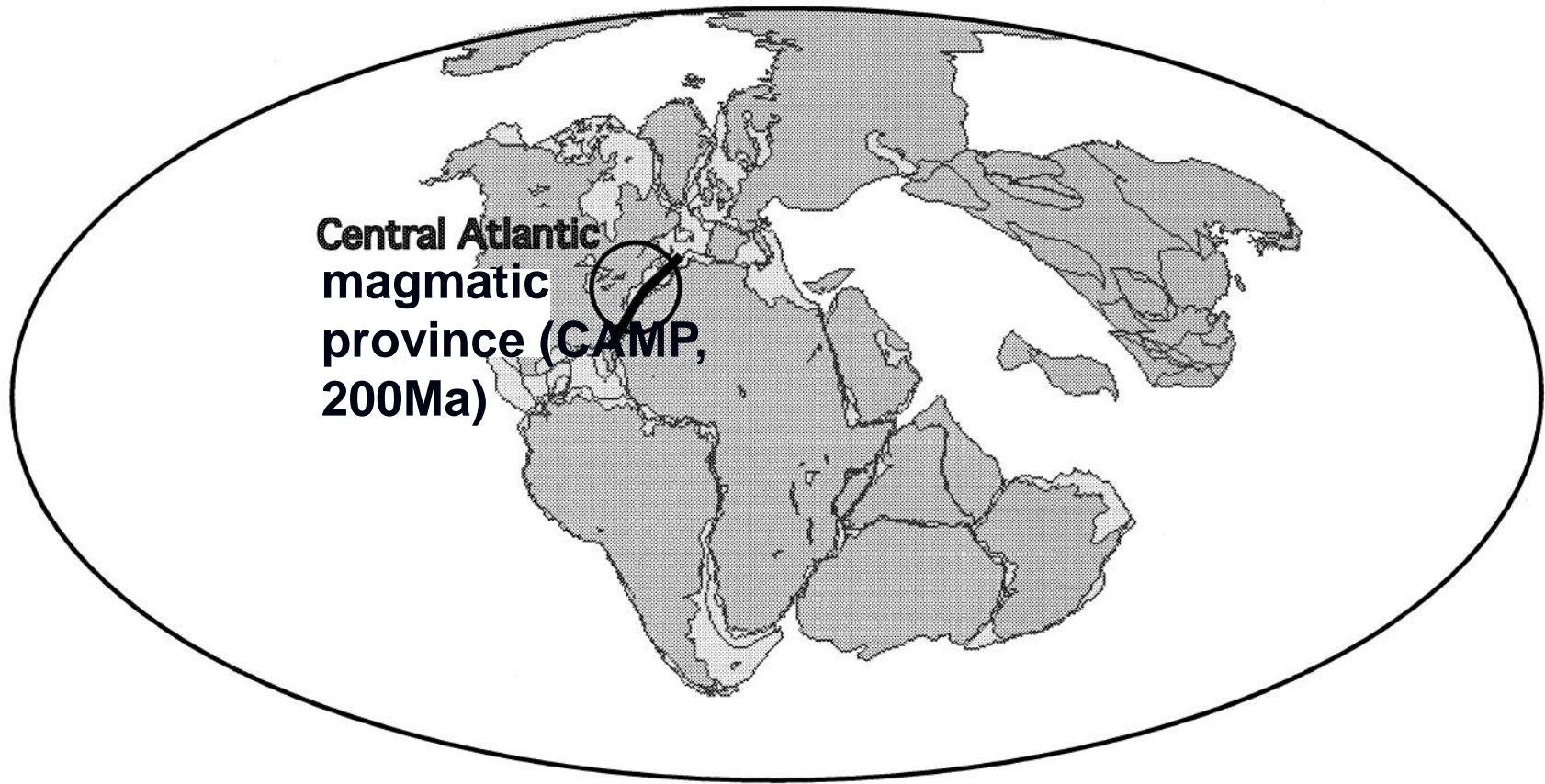
- How to break a continent?
 - Effect of magmas and Large Igneous Provinces
 - Effect of oblique rifting
- Continental transform faults
 - What caused Dead Sea transform?
 - San Andreas Fault System

Continental break-up

Continental break-up

186

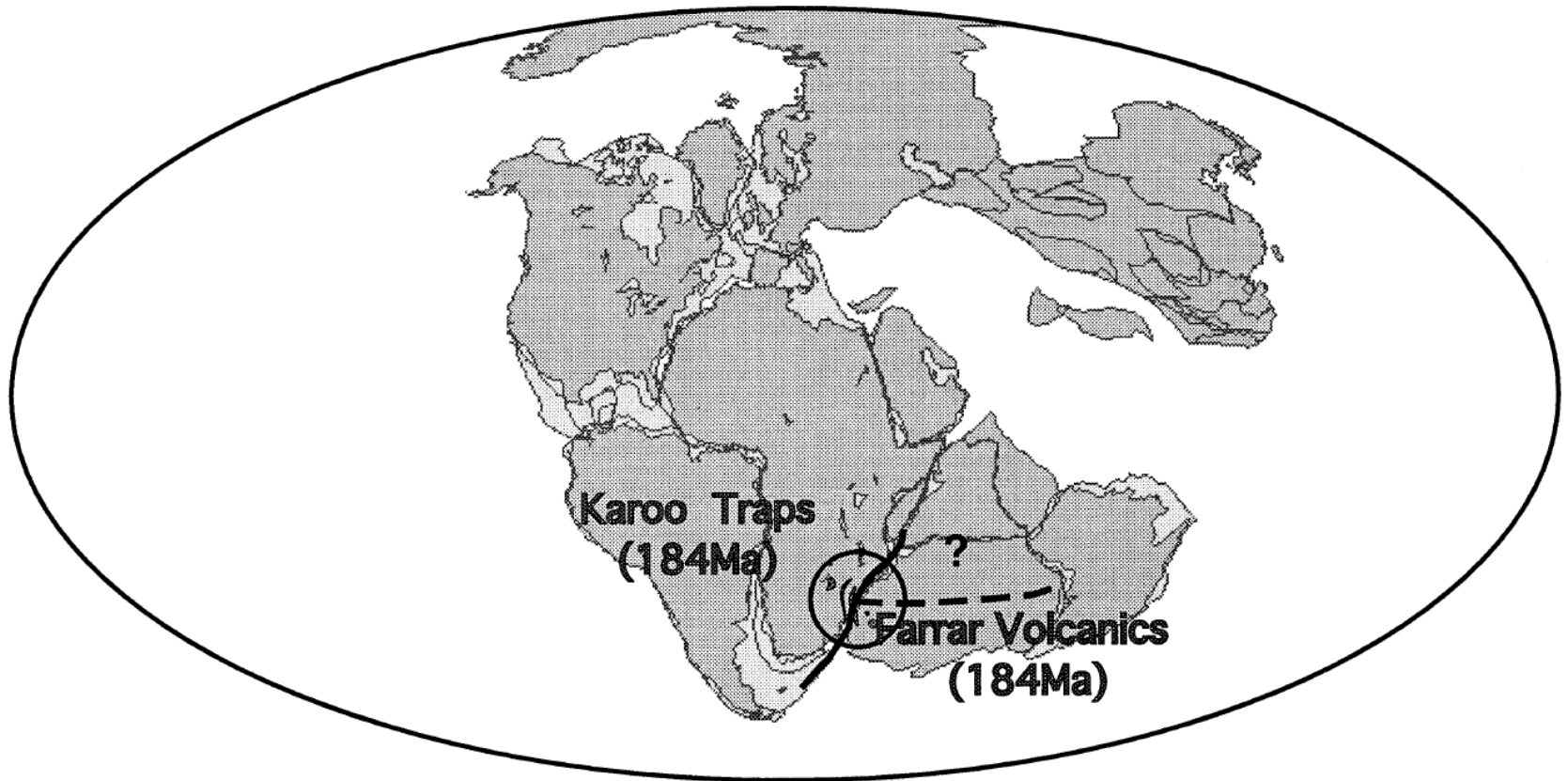
V. Courtillot et al. / Earth and Planetary Science Letters 166 (1999) 177–195



Continental break-up

V. Courtillot et al. / Earth and Planetary Science Letters 166 (1999) 177–195

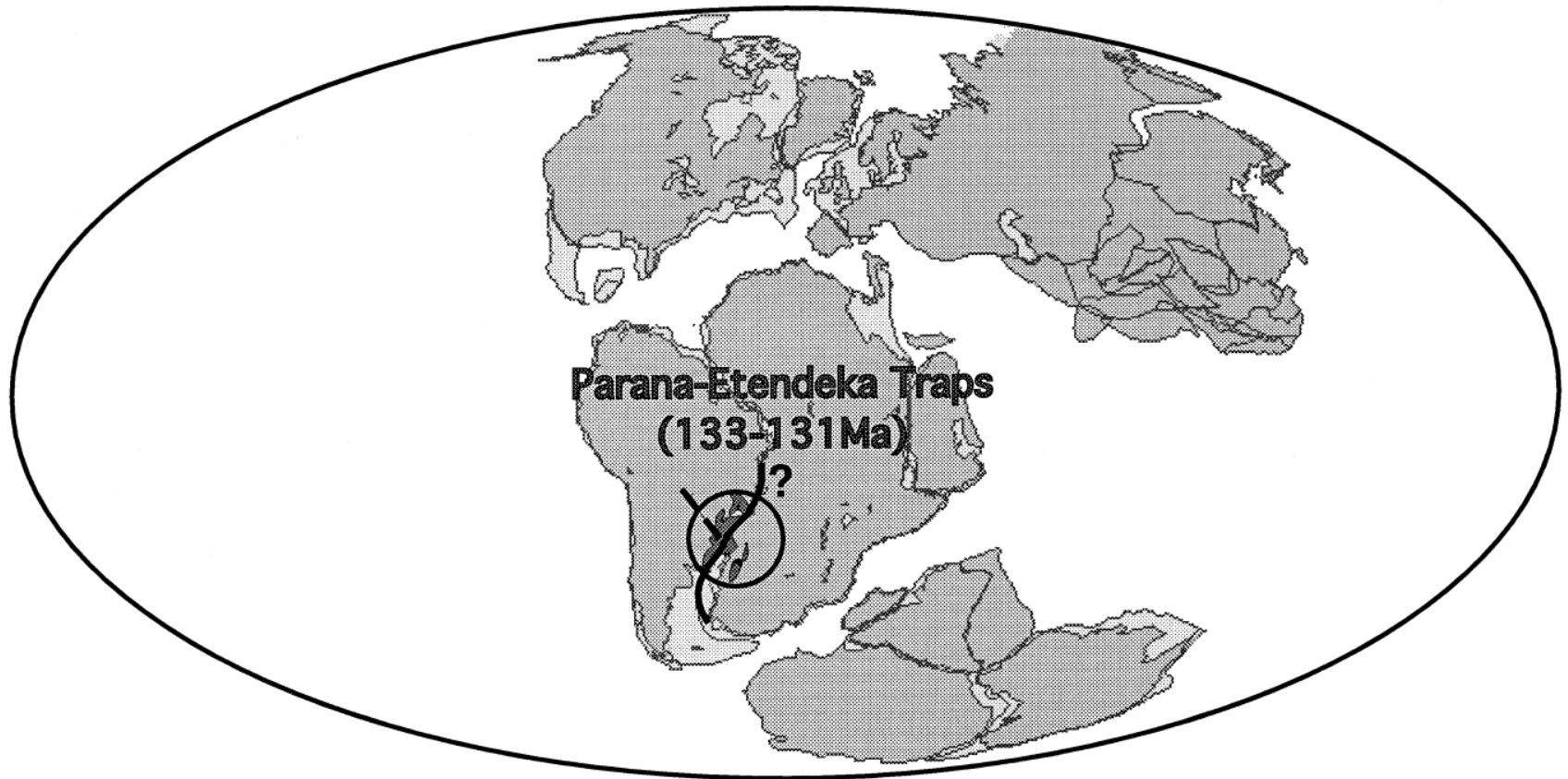
185



Continental break-up

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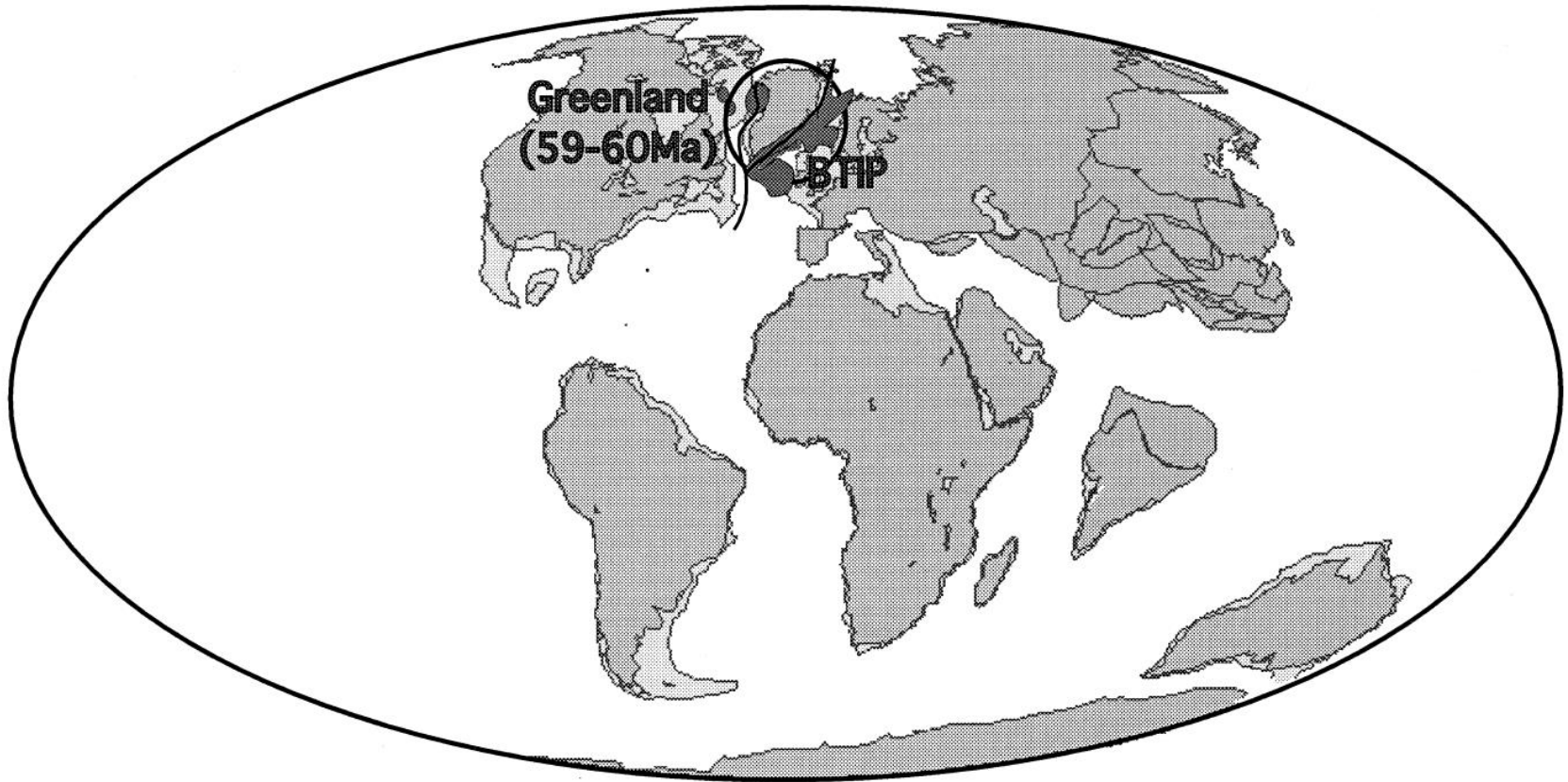
V. Courtillot et al. / Earth and Planetary Science Letters 166 (1999) 177–195



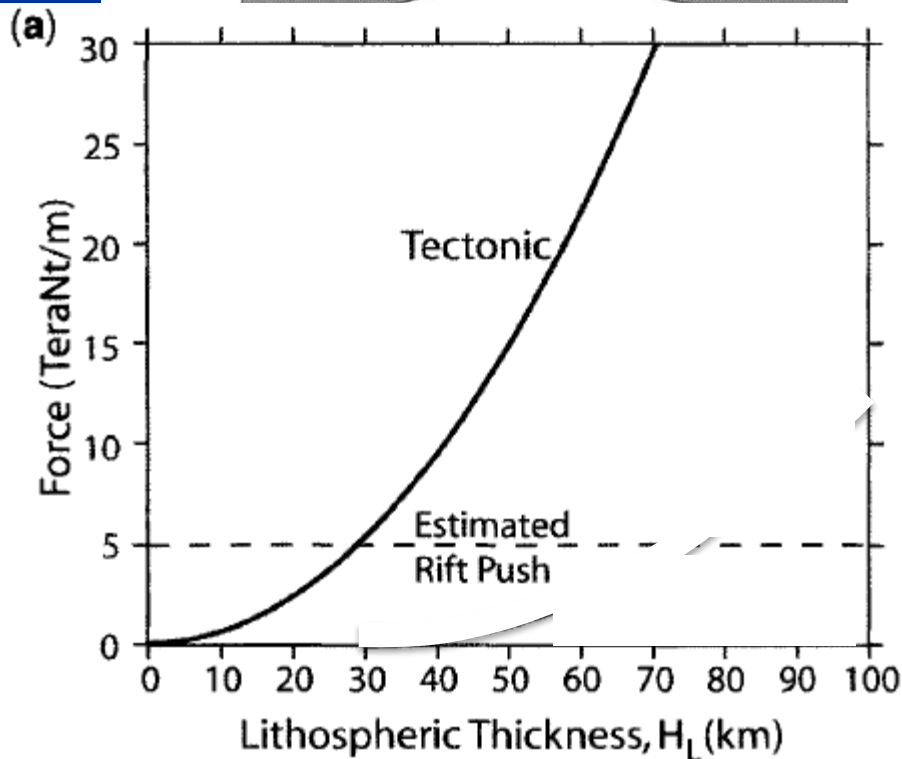
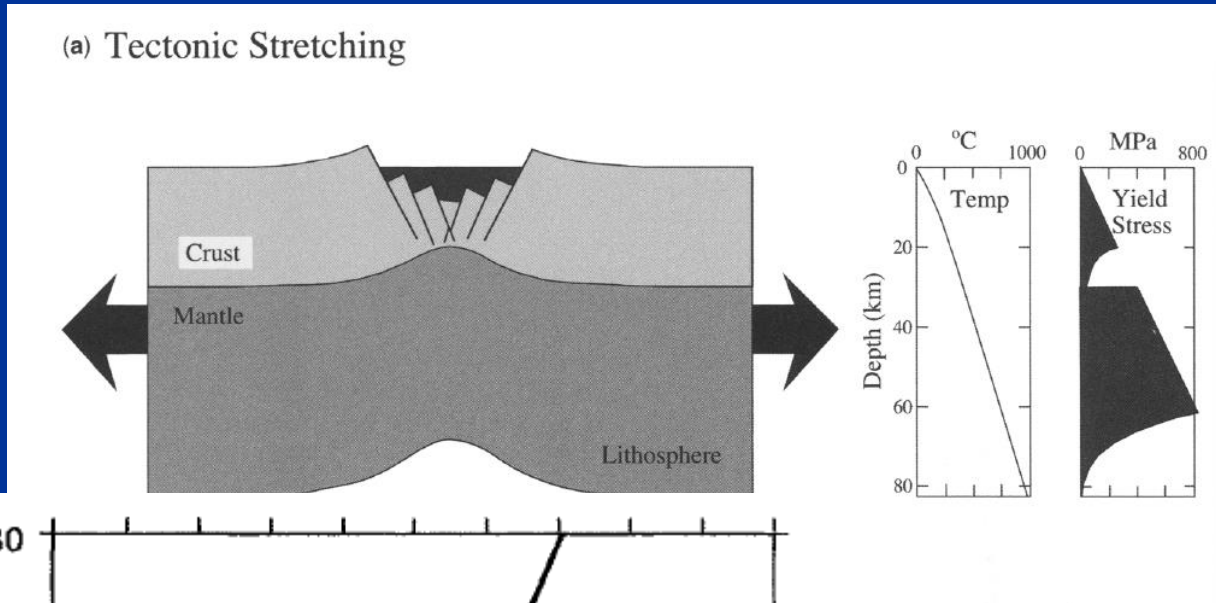
Continental break-up

182

V. Courtillot et al. / Earth and Planetary Science Letters 166 (1999) 177–195



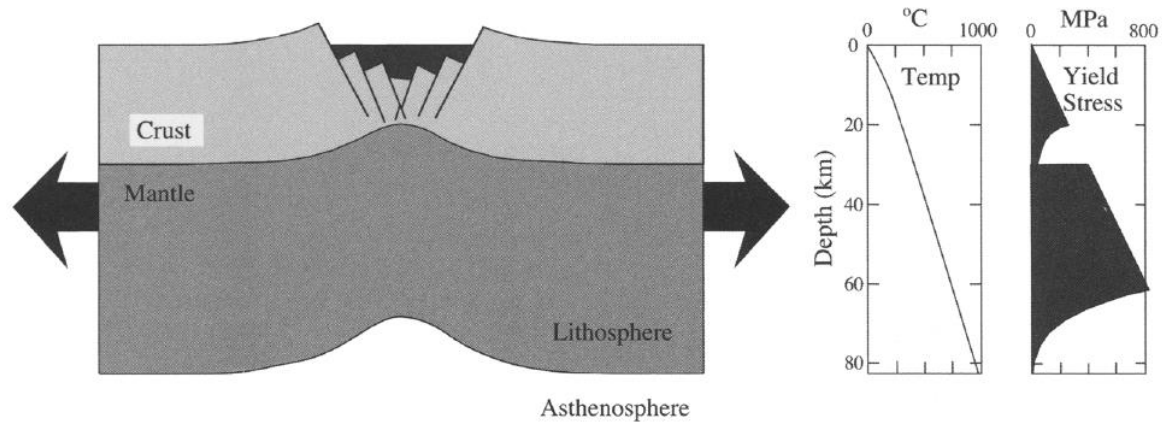
How to break continent?



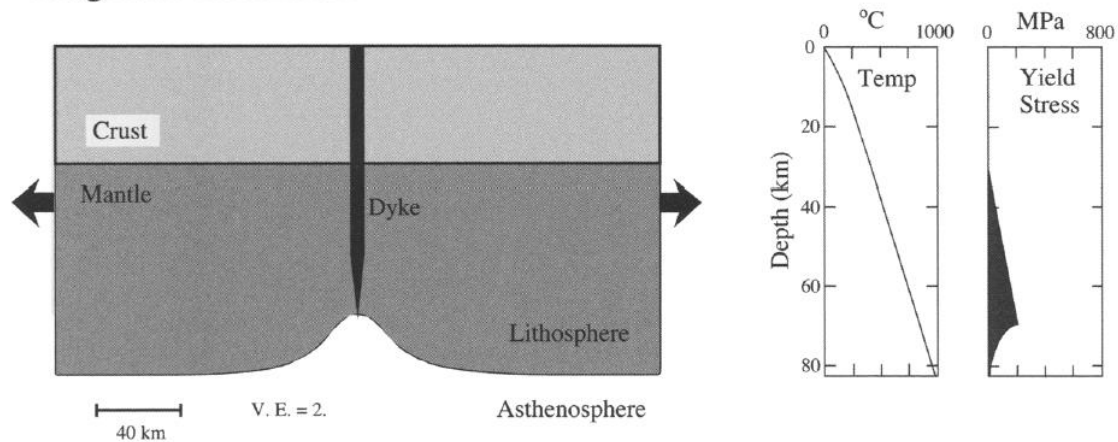
Cold lithosphere is too strong

Effect of magma-filled dikes

(a) Tectonic Stretching

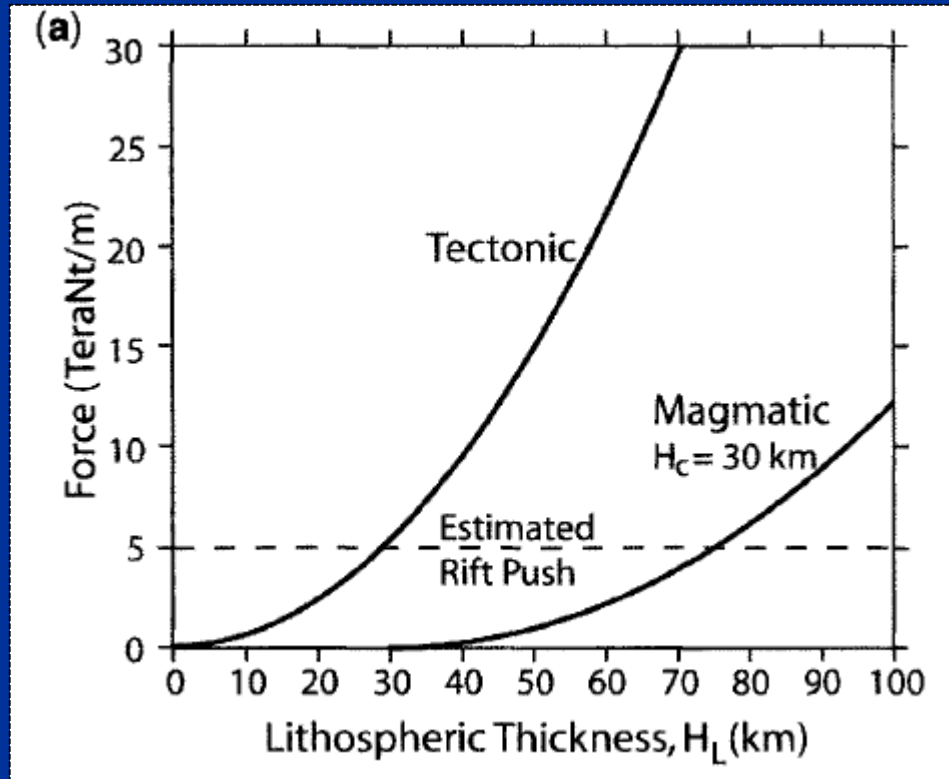


(b) Magmatic Extension



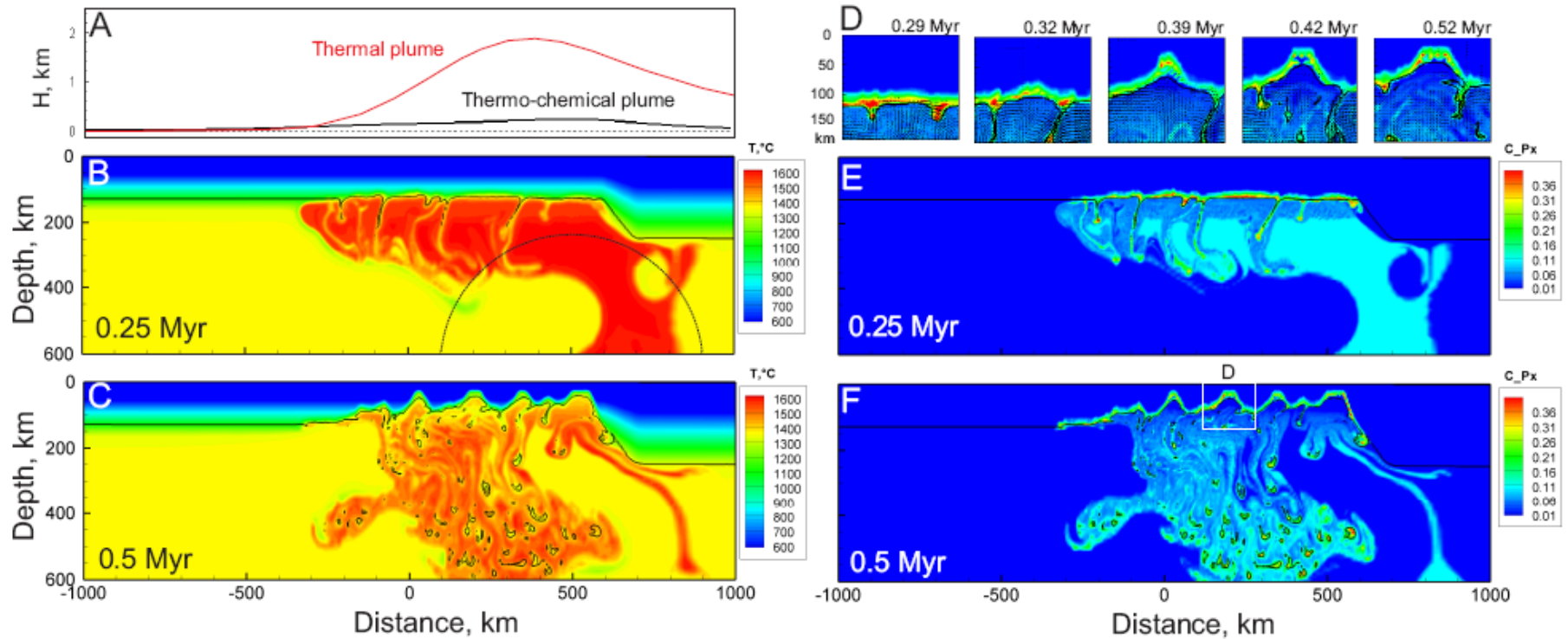
Buck (2006)

Effect of magma-filled dikes

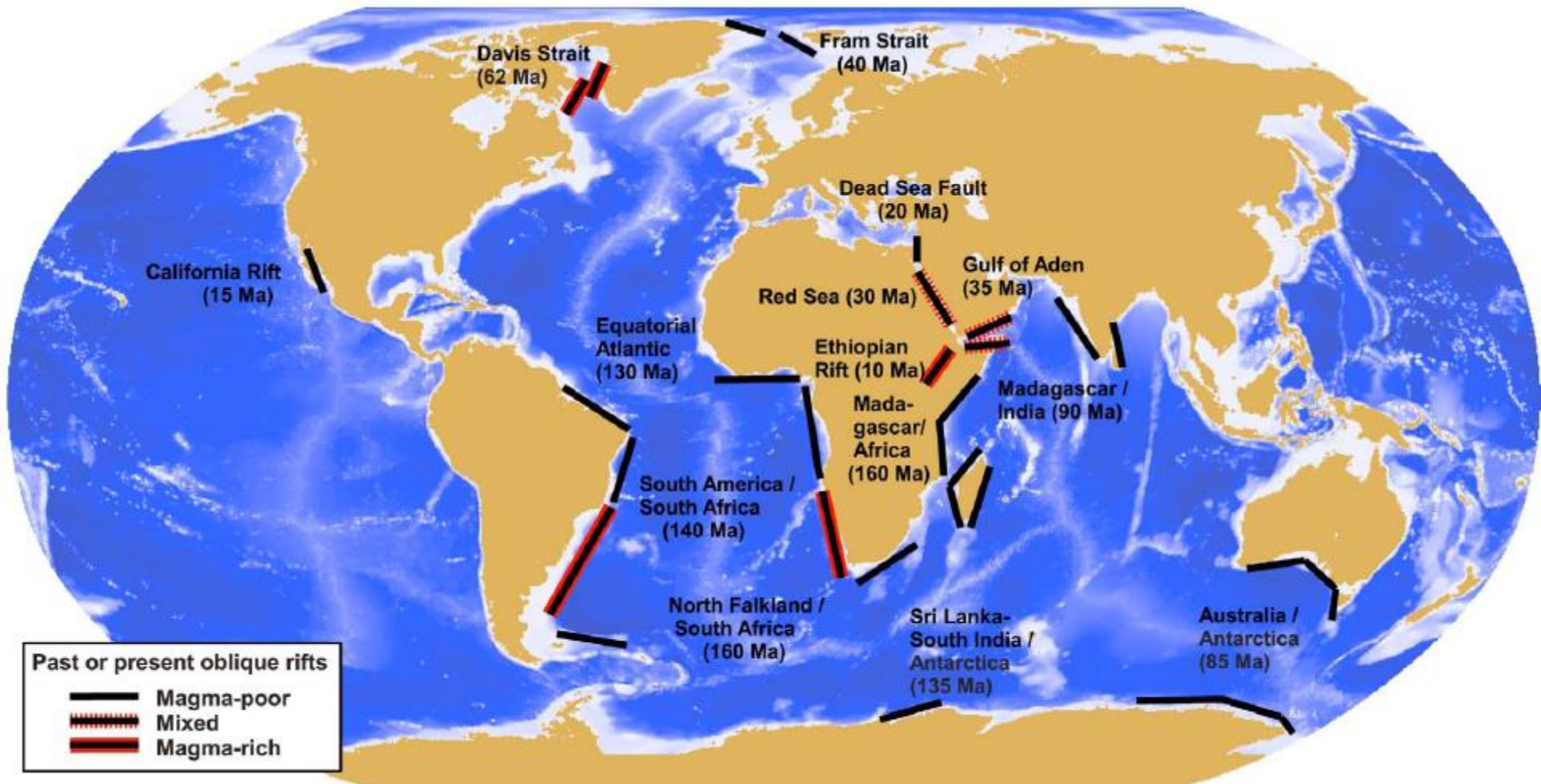


It works if lithosphere is first thinned to about 75 km

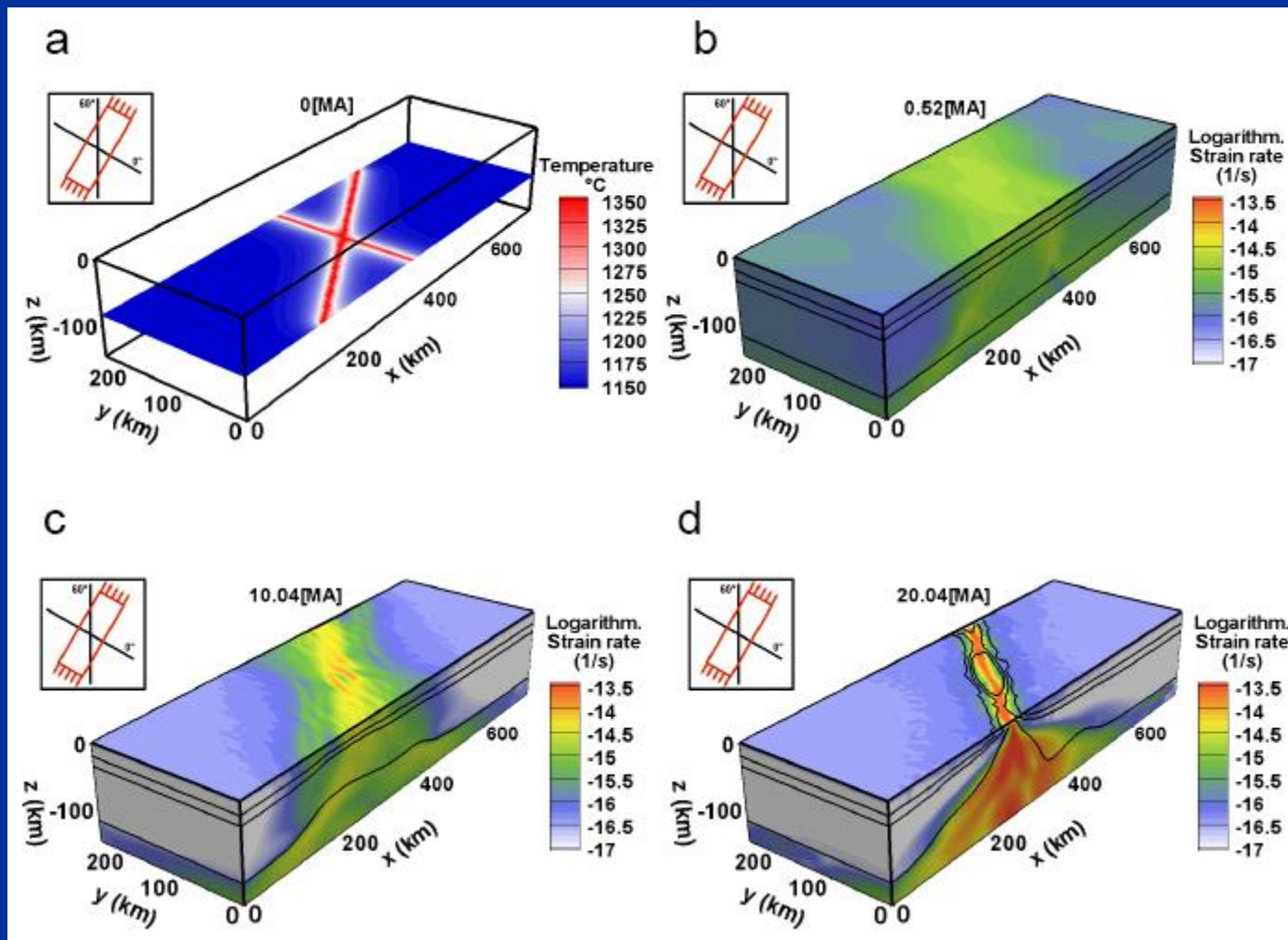
Lithospheric thinning above mantle plume



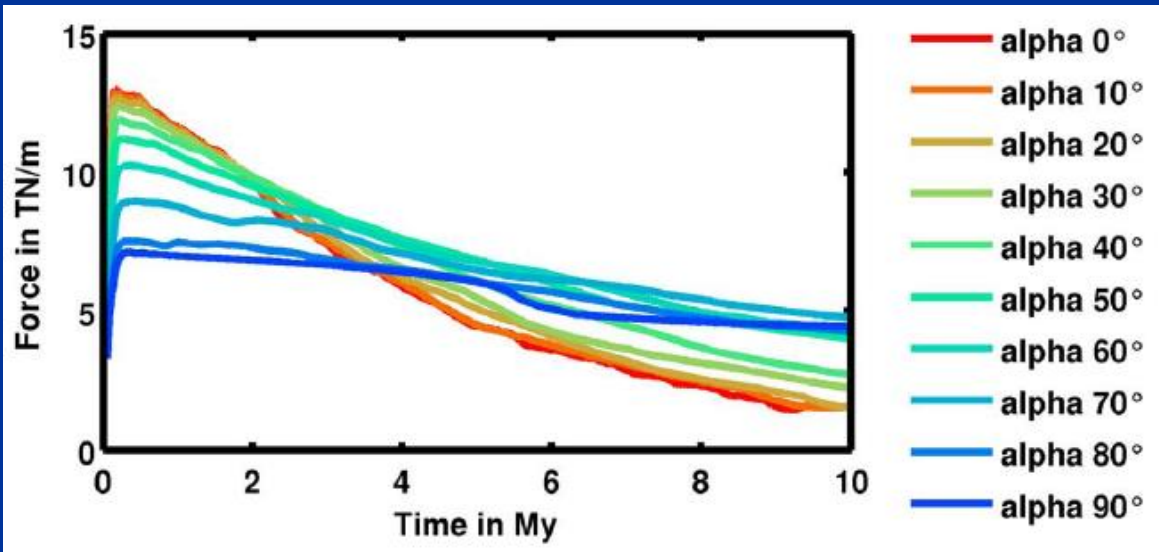
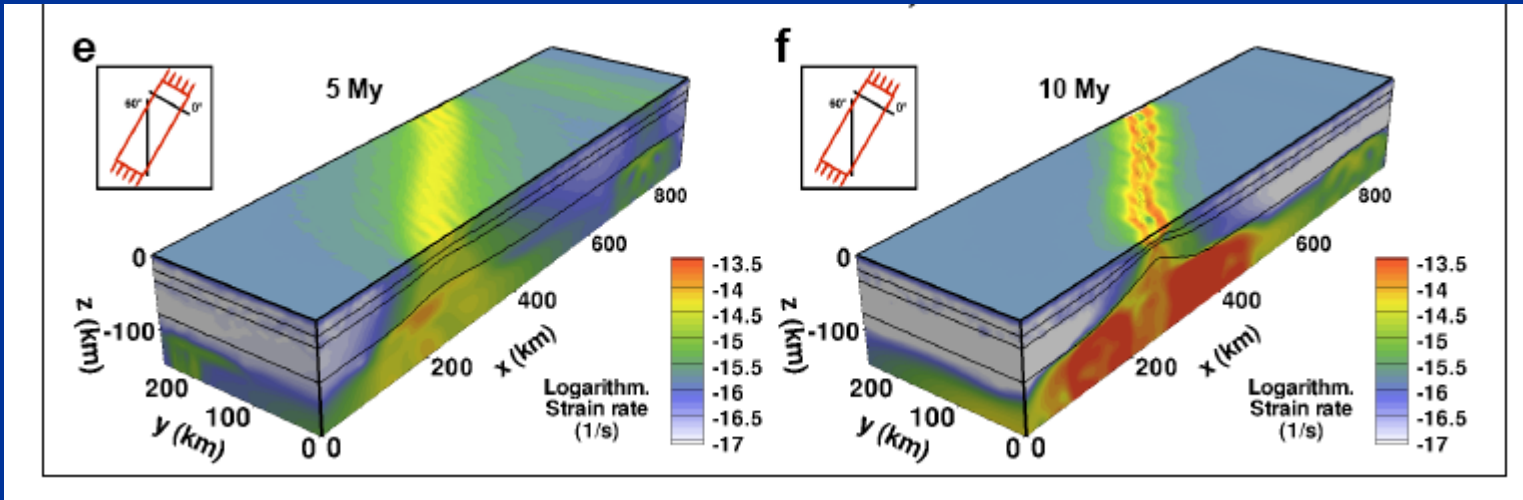
Effect of oblique rifting



Effect of oblique rifting



Effect of oblique rifting

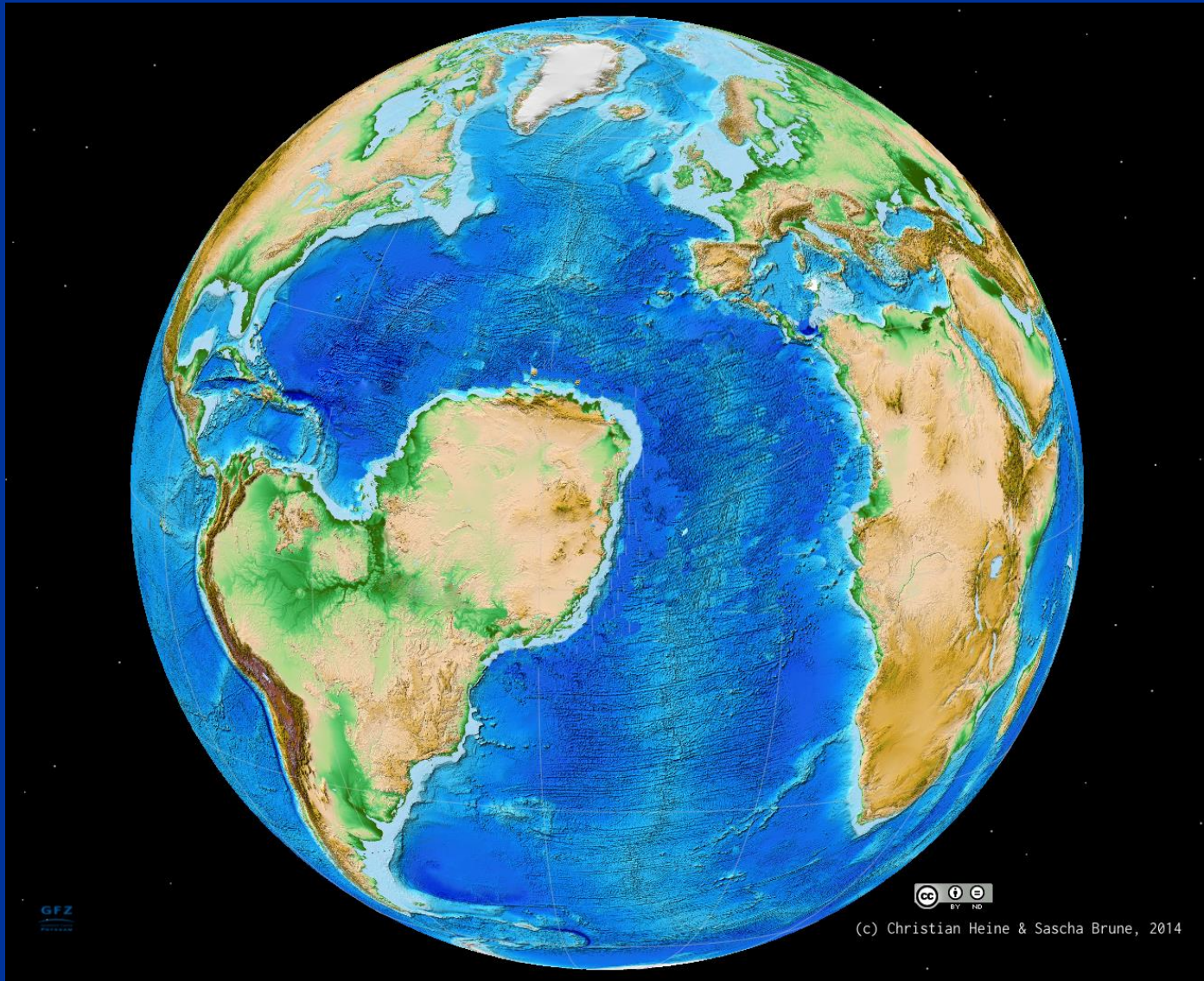


$$F_{strike-slip} = \tau_{yield} L_z$$

$$F_{extension} = \frac{\tau_{yield} L_z}{\sqrt{\frac{1}{3}(\nu^2 - \nu + 1)}}$$

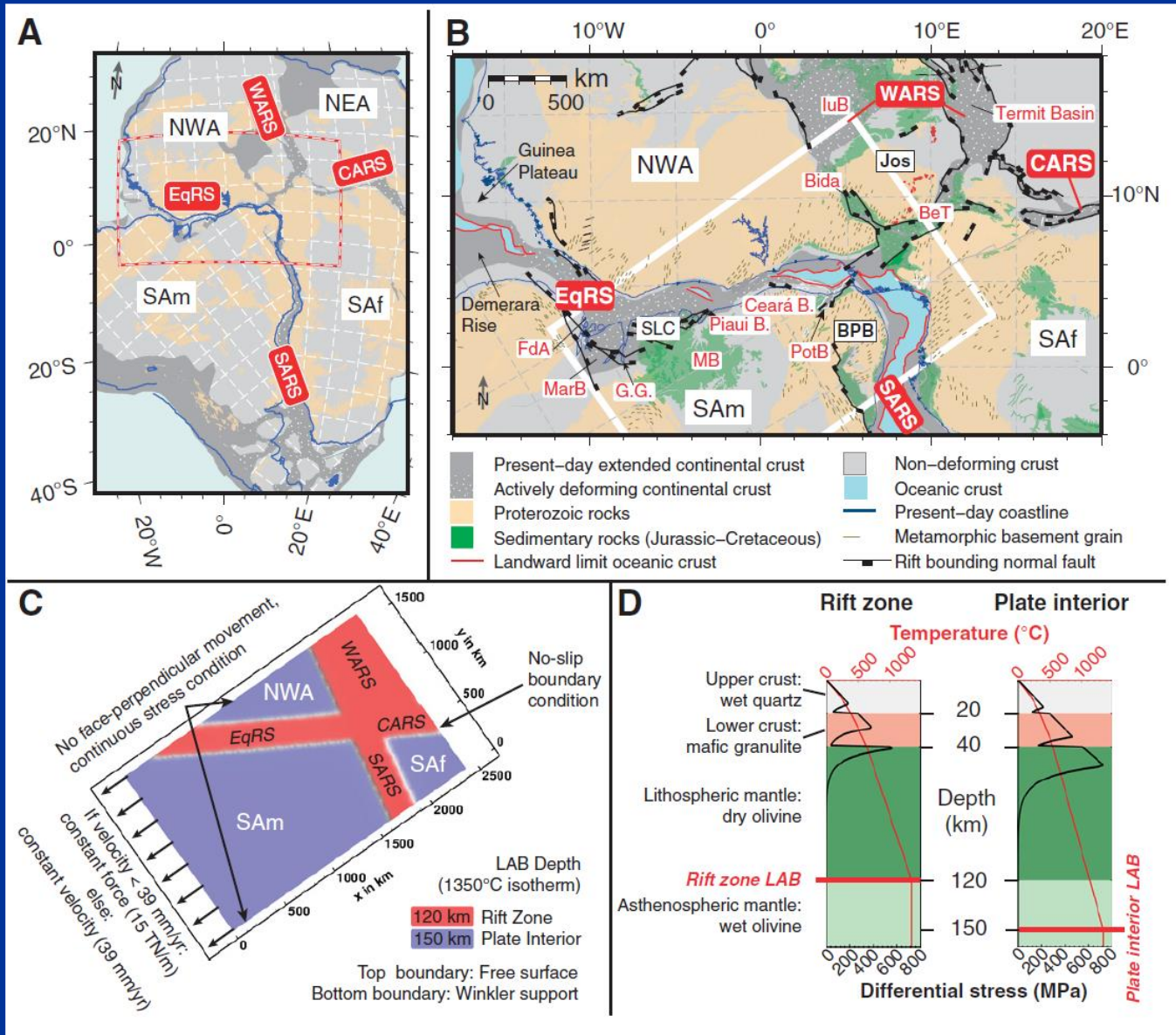
$$F_{extension} = 2\tau_{yield} L_z.$$

Effect of oblique rifting (example)

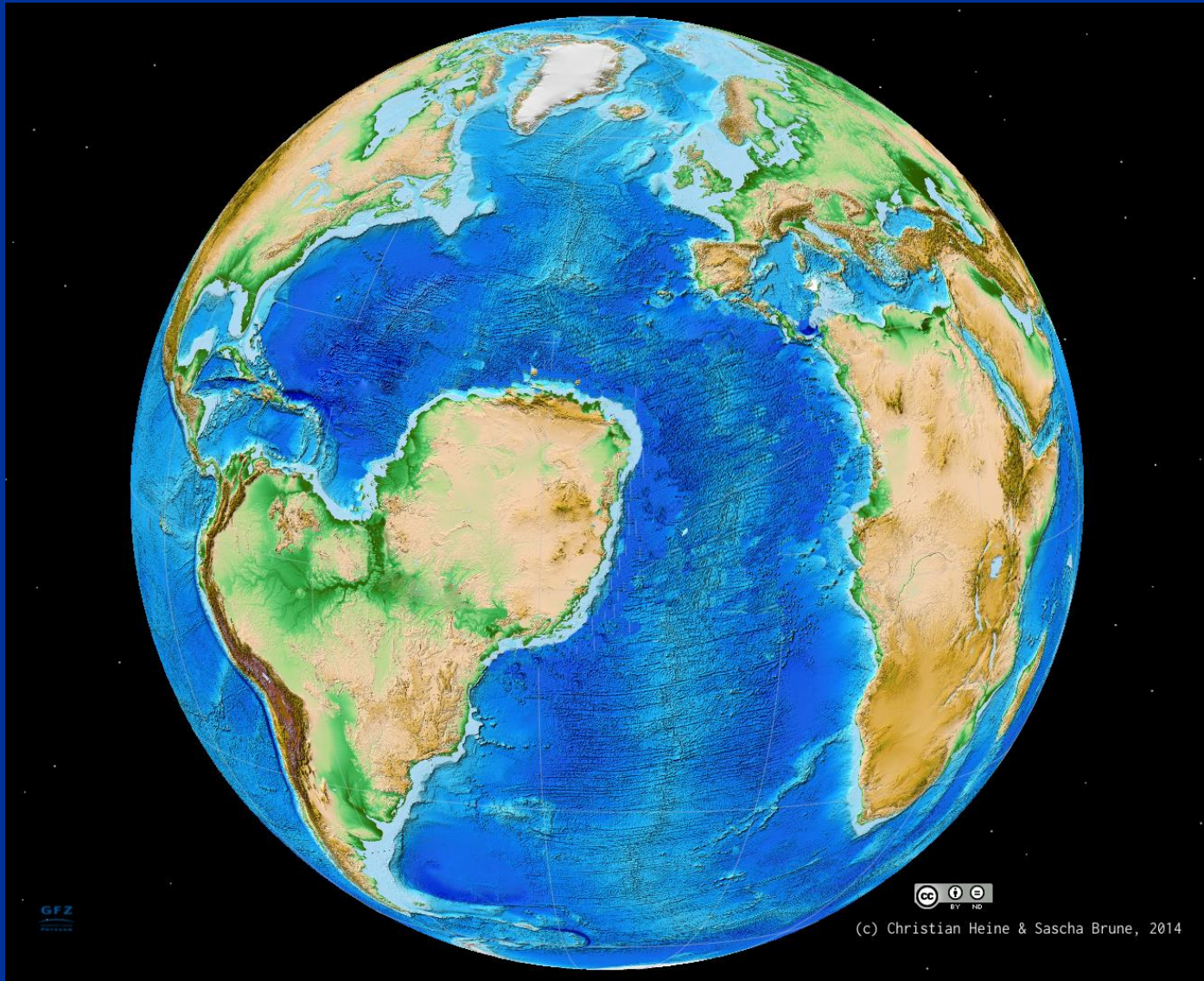


Effect of oblique rifting (example)

Heine and Brune, *Geology*, 2014

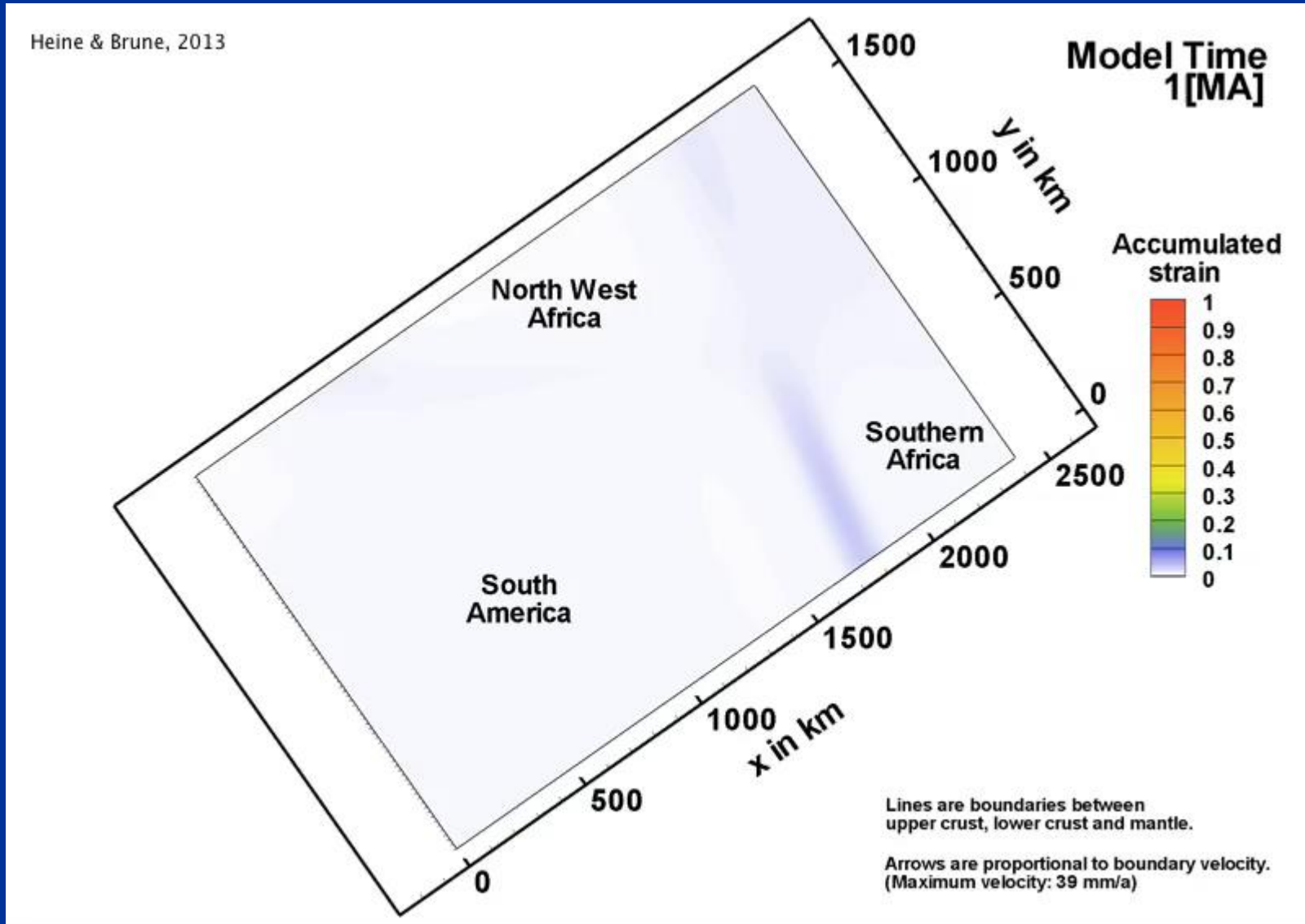


Effect of oblique rifting (example)



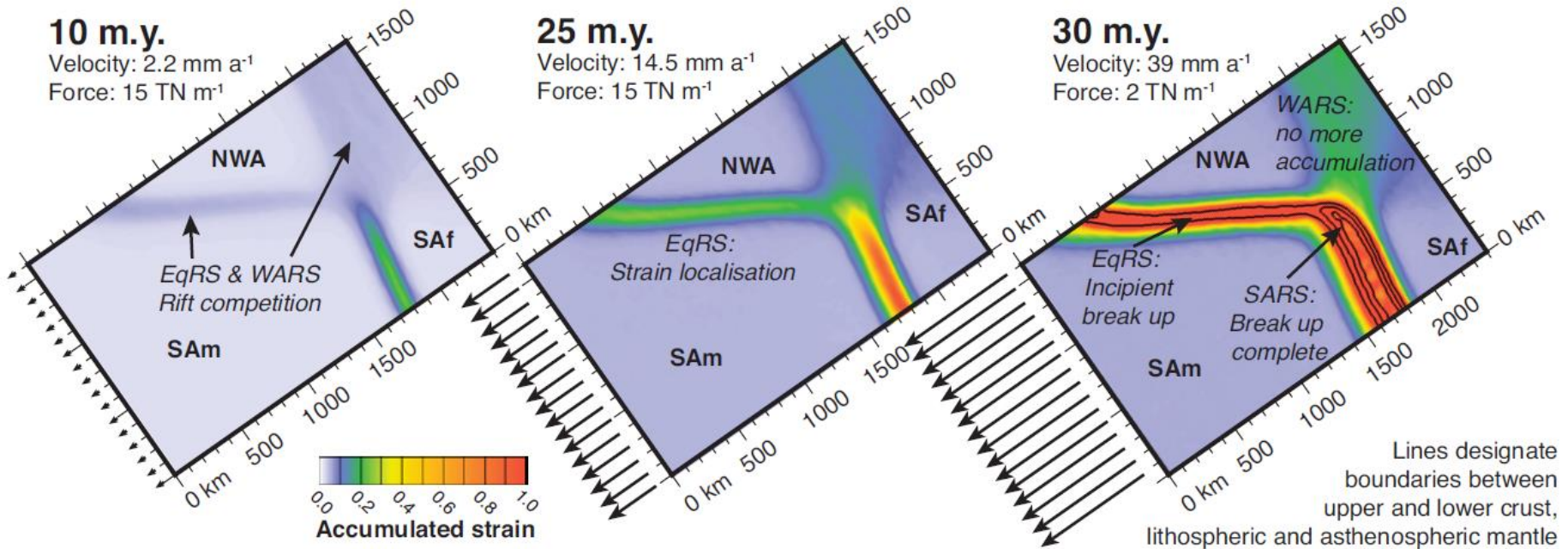
Effect of oblique rifting (example)

Heine and Brune, Geology, 2014



Effect of oblique rifting (example)

Heine and Brune, Geology, 2014



Conclusion

To break a continent are required:

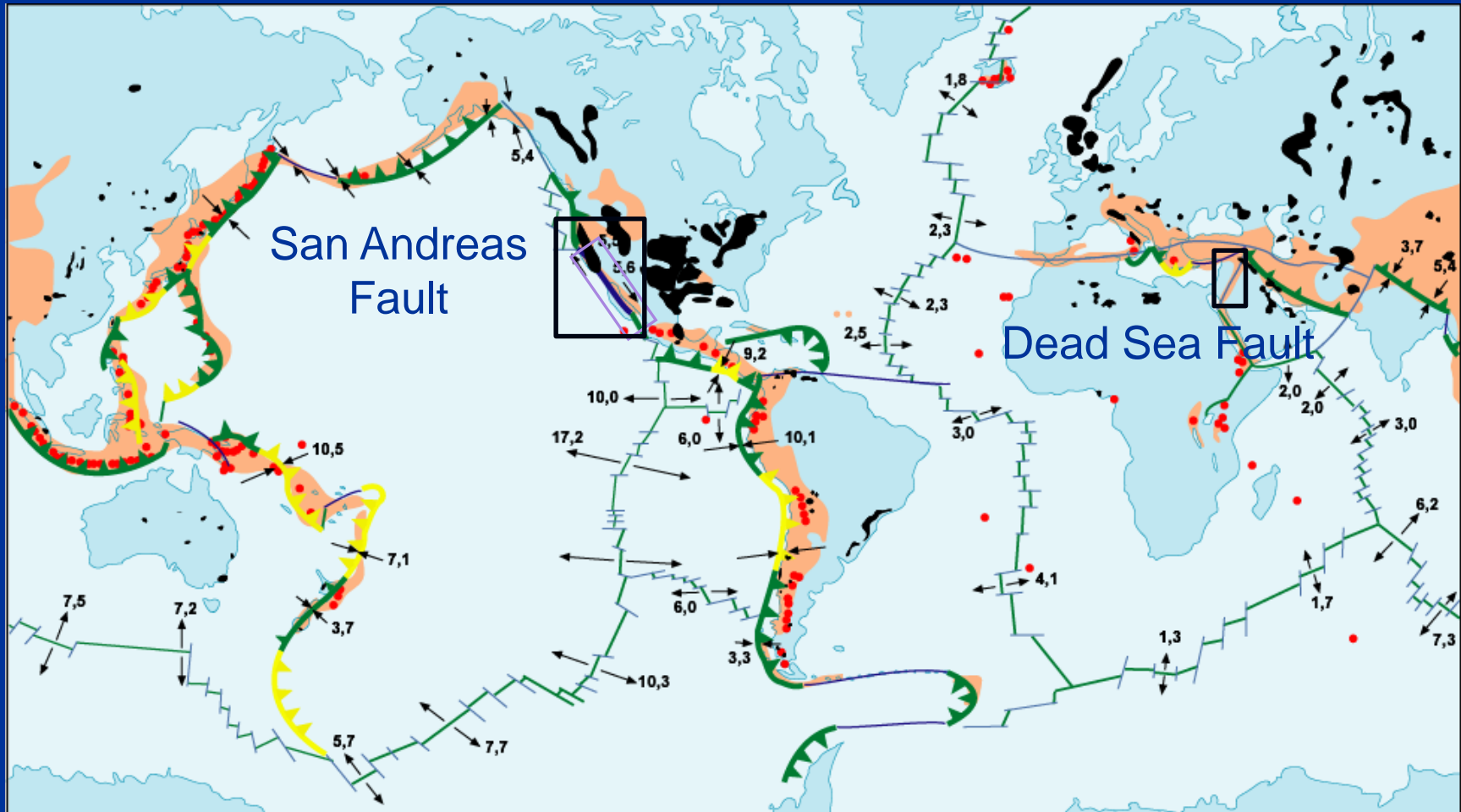
(1) extensional deviatoric stresses (internal, from ridge push or subduction zones roll-back) and (2) lithospheric weakening

Large Igneous Provinces are optimal for lithospheric weakening, as they may both thin lithosphere and generate magma-filled dikes.

Intensive strike-slip deformation is also helpful

Continental transform faults (case Dead Sea Transform)

Continental Transform Faults



convergent plate margins

▶ erosive
 ▶ accretionary
 ▶ transform

divergent plate margins

← → 7,7 spreading rate (cm/a)

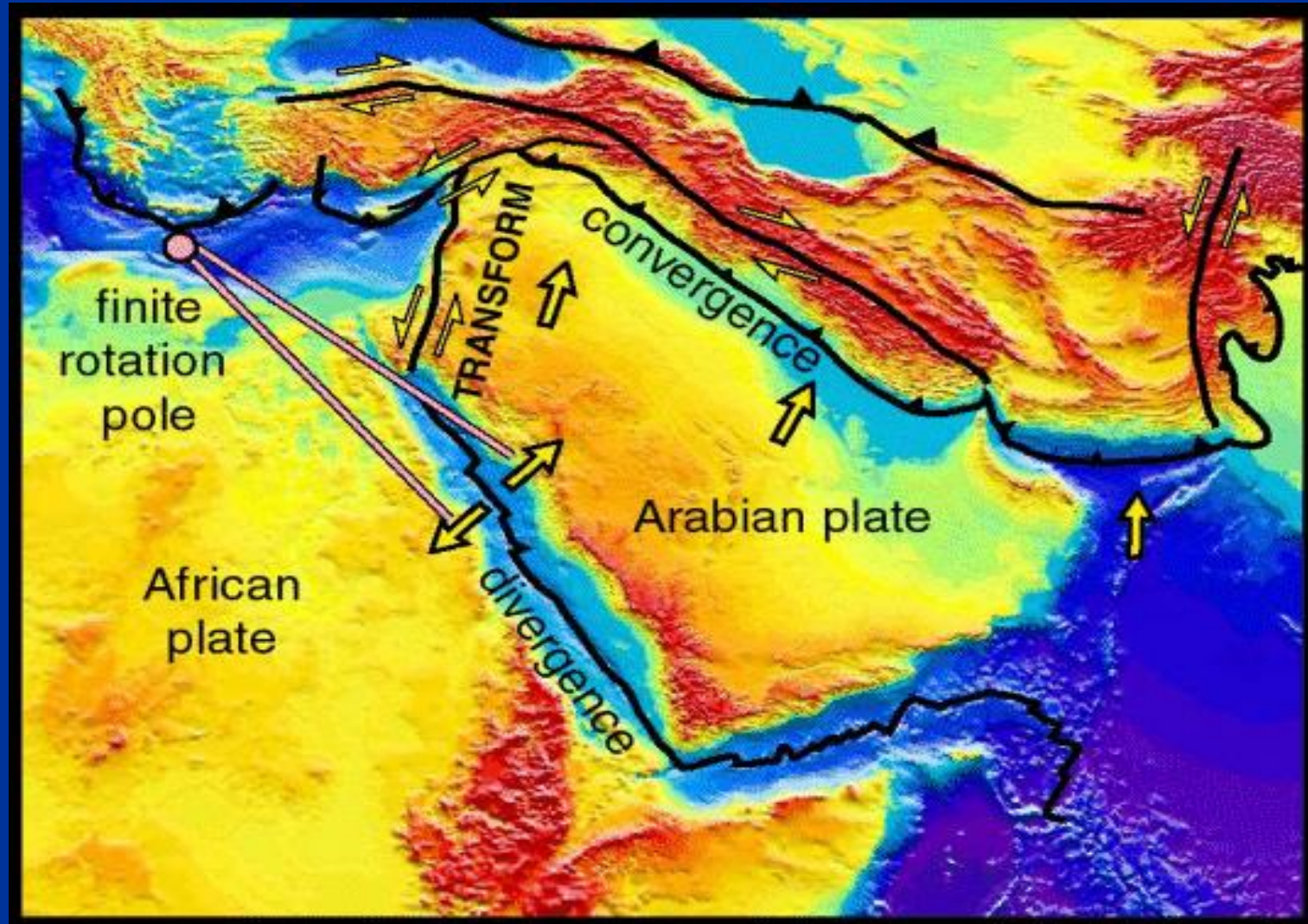
◡ earthquake zones

● active volcanoes

■ hydrocarbons

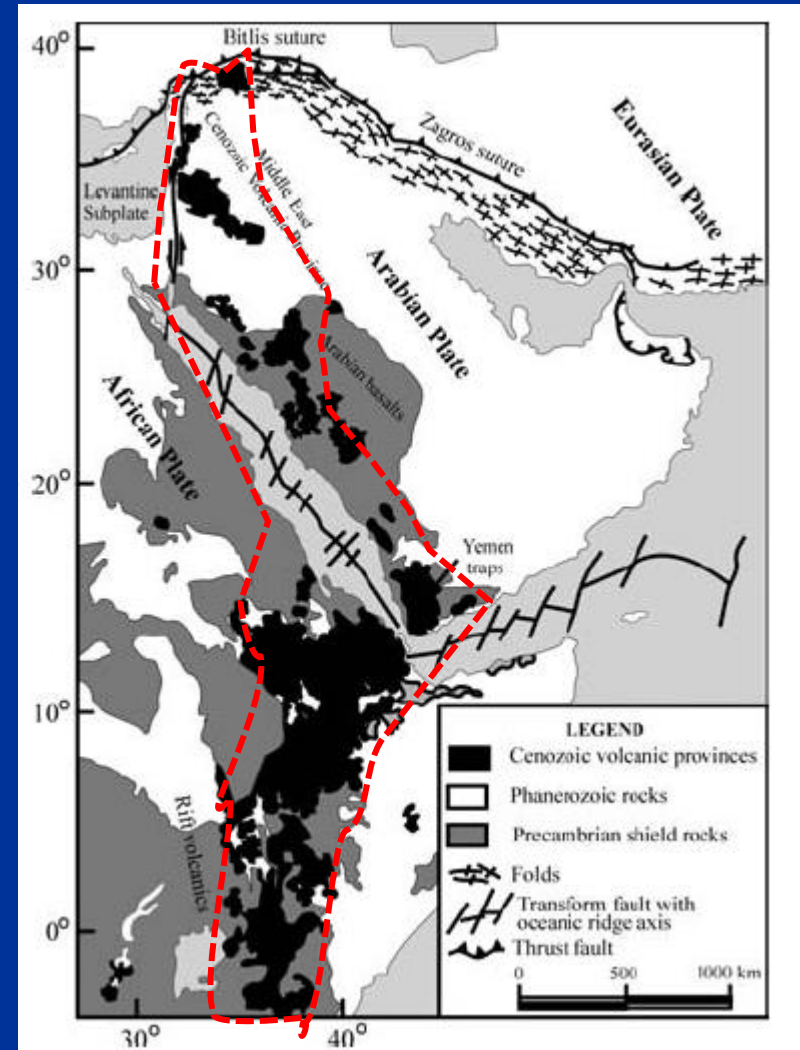
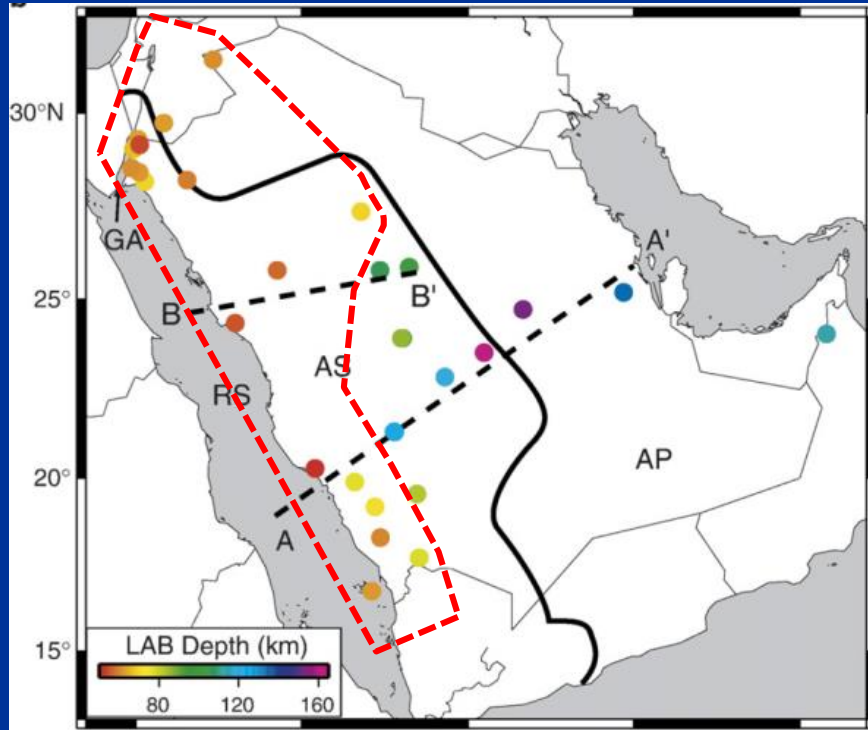
Regional setting

With the surface heat flow of 50-60 mW/m², the DST is the coldest continental transform boundary



Lithospheric thickness and magmatism

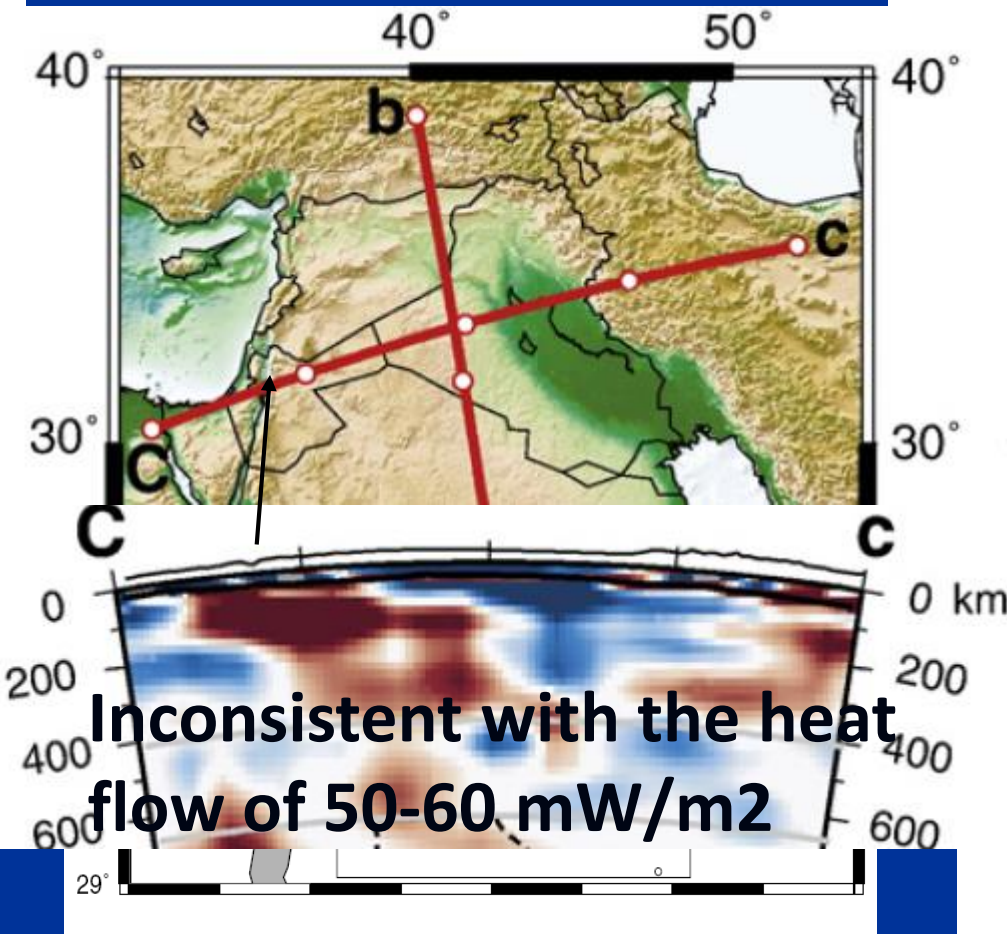
Magmatism at 30-0 Ma



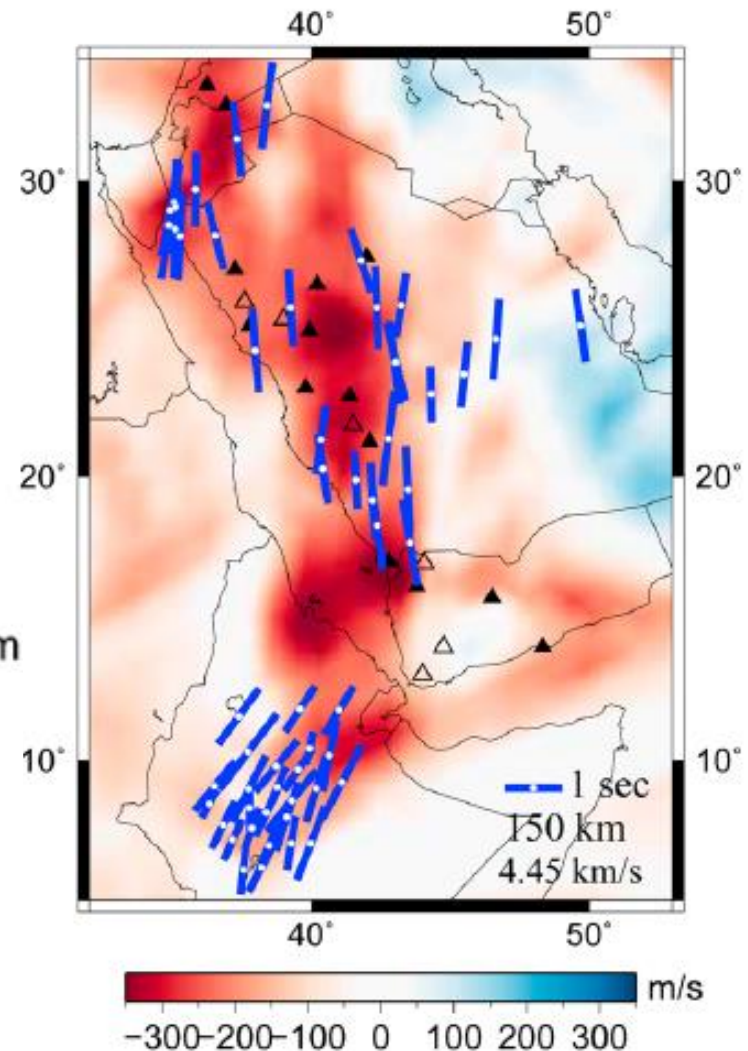
Lithosphere-asthenosphere boundary (LAB) from seismic data

Chang and Van der Lee, EPSL, 2011

Chang et al, GRL, 2011



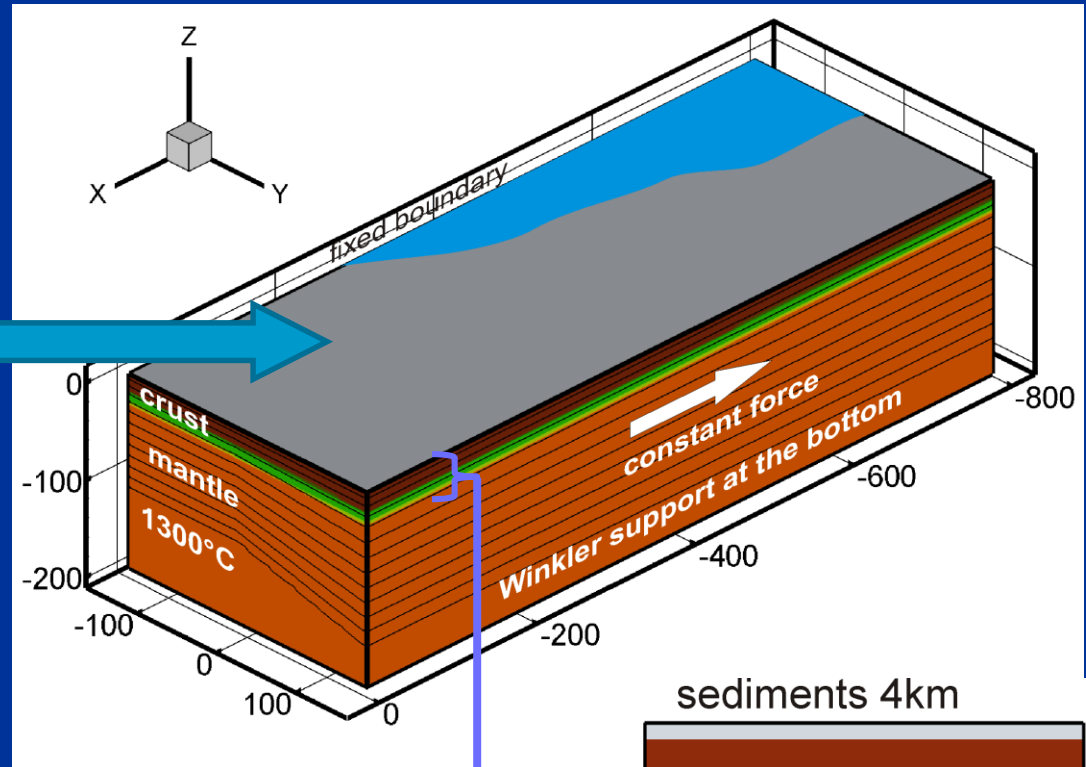
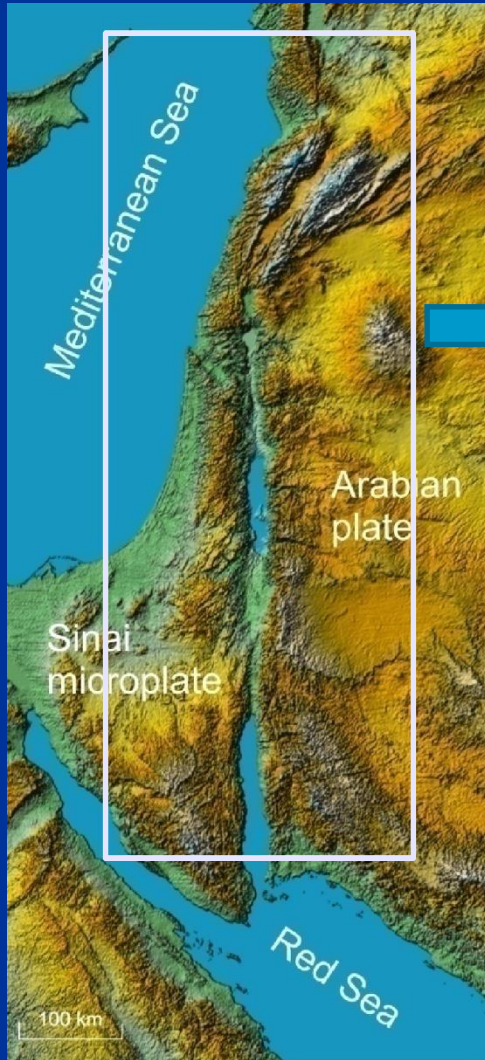
Inconsistent with the heat flow of 50-60 mW/m²



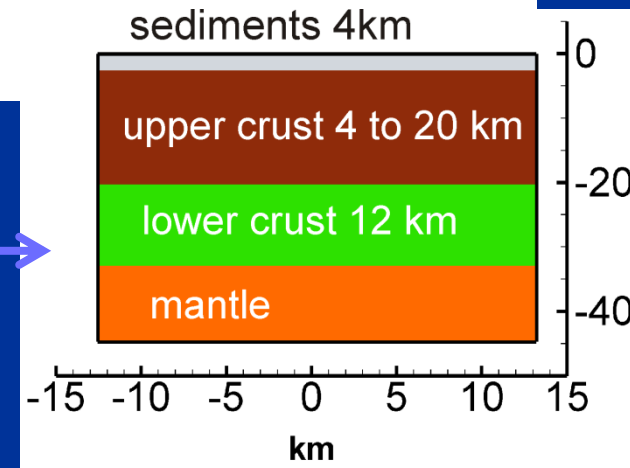
Conclusion

Lithosphere around DST was thinned in the past and related high heat flow had not enough time to reach the surface

Model setup

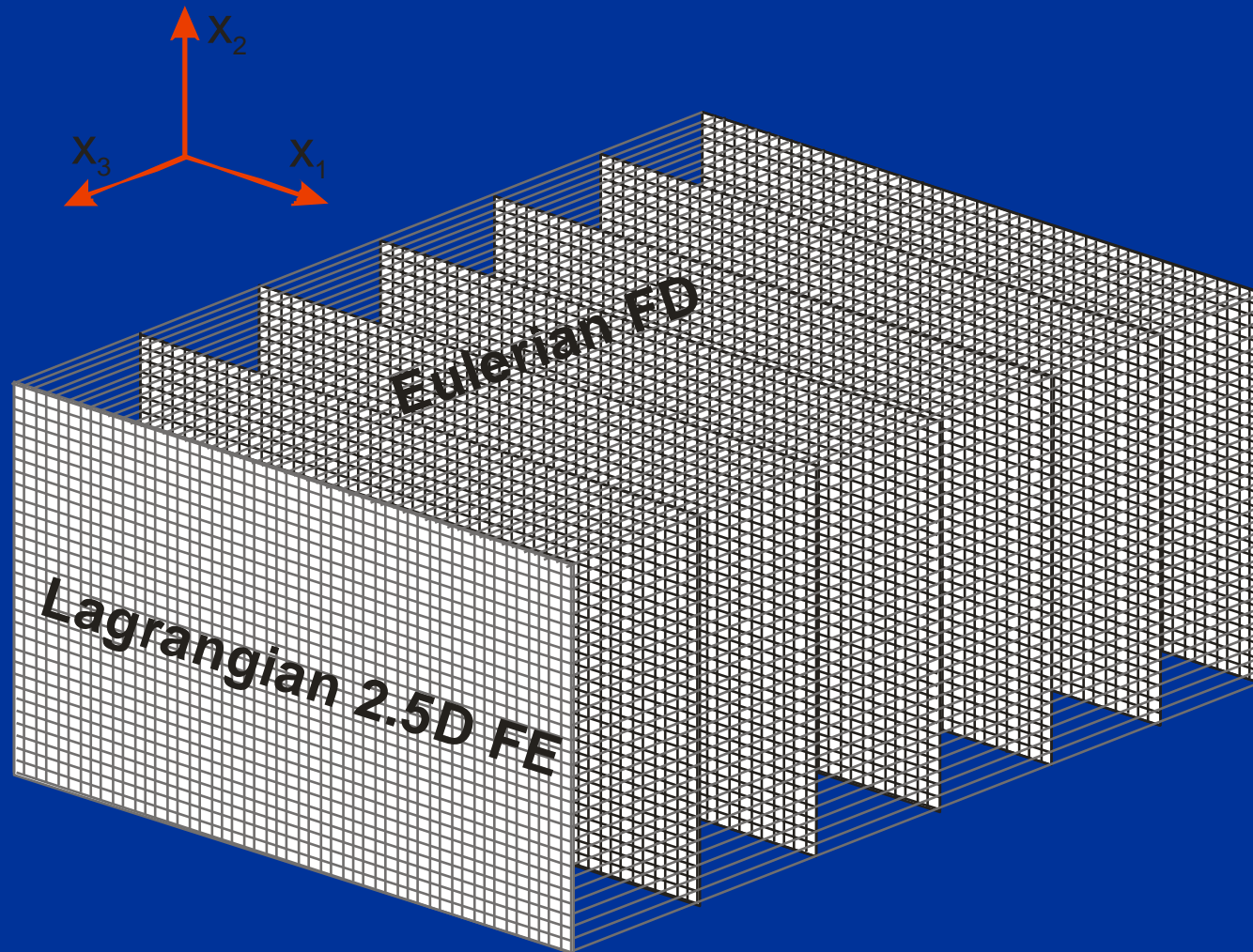


Flat Earth
approximation



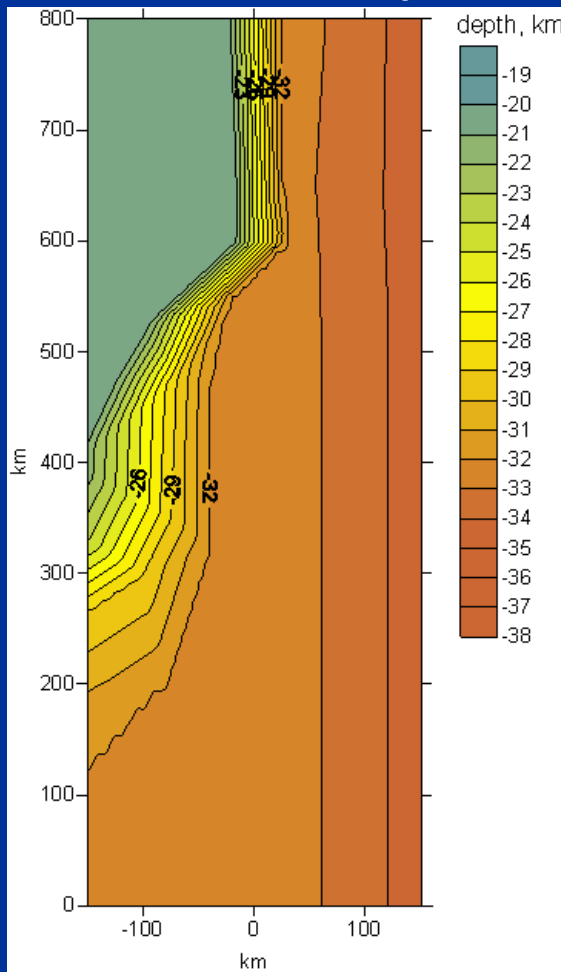
Modeling technique LAPEX 3D combining FE and FD

(Petrinin and Sobolev, Geology, 2006, PEPI, 2008)

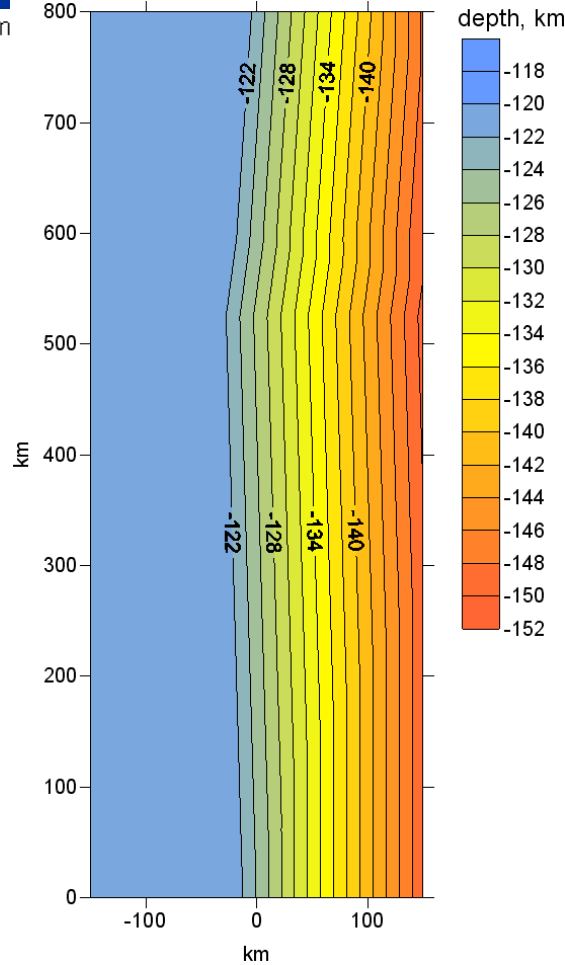


Initial lithospheric structure:

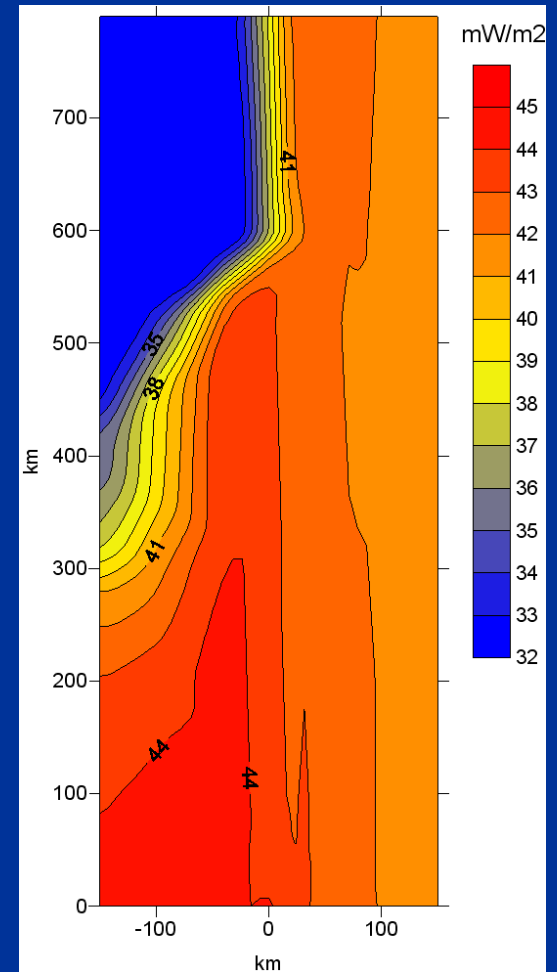
Moho map



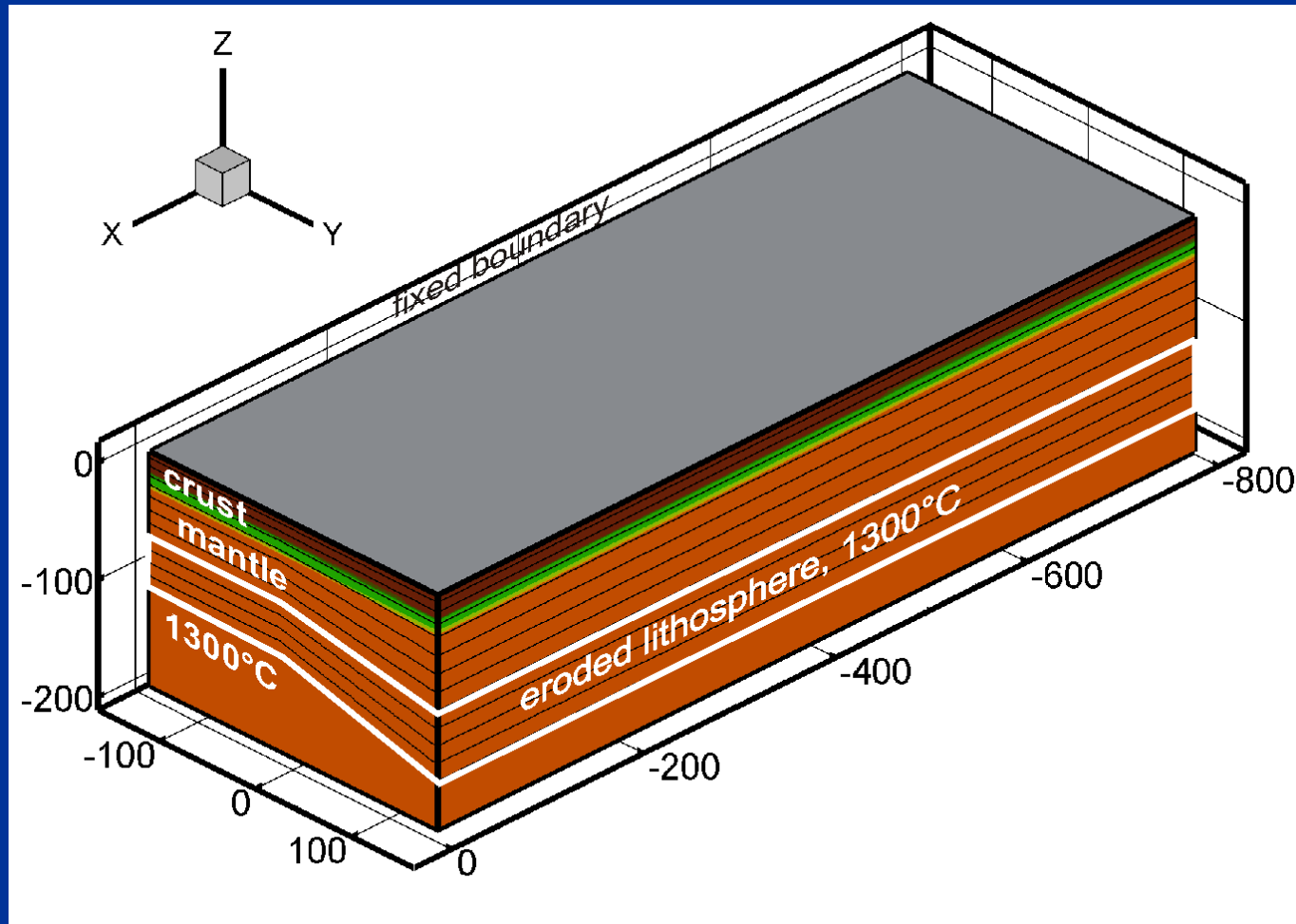
LAB map



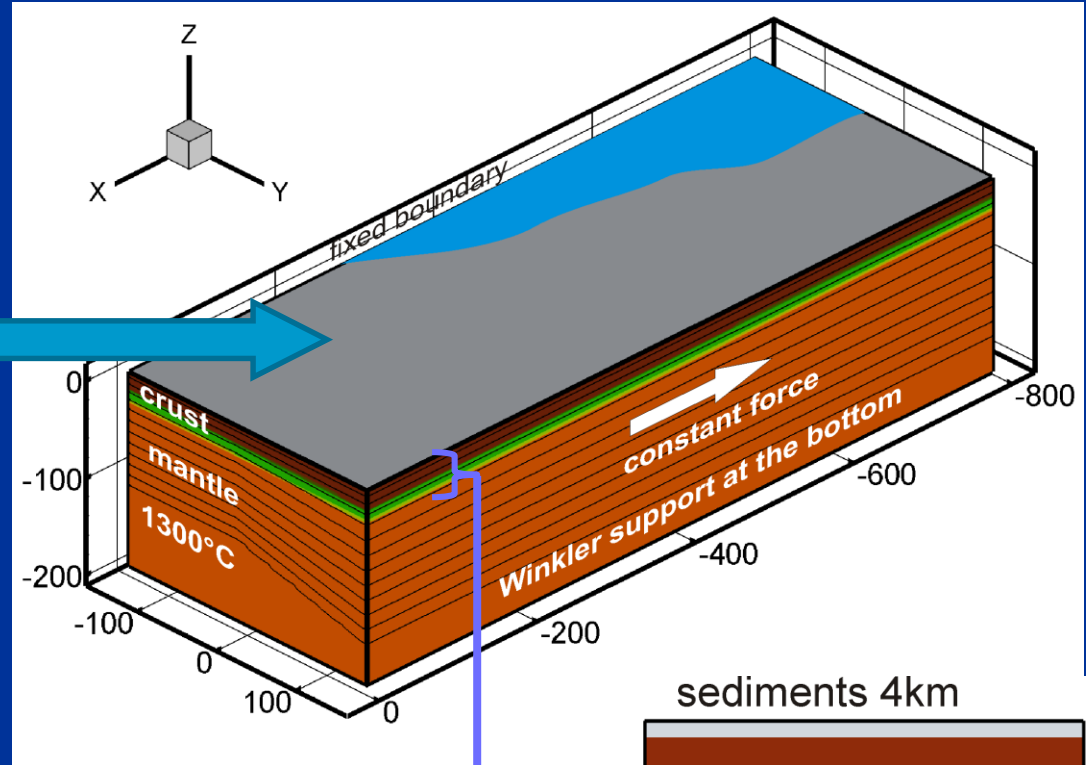
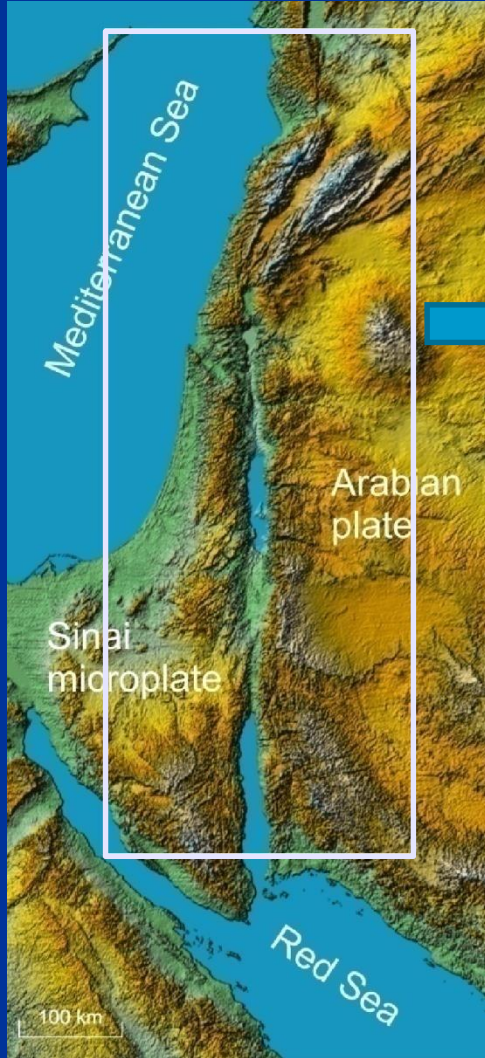
Heat flow



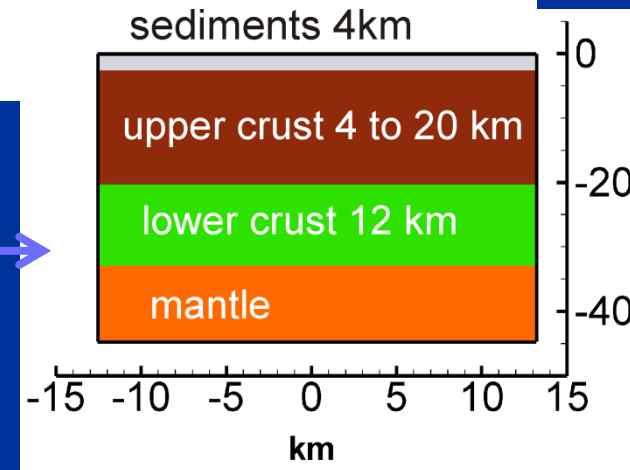
Modeling results: role of the thermal erosion of the lithosphere



Model setup

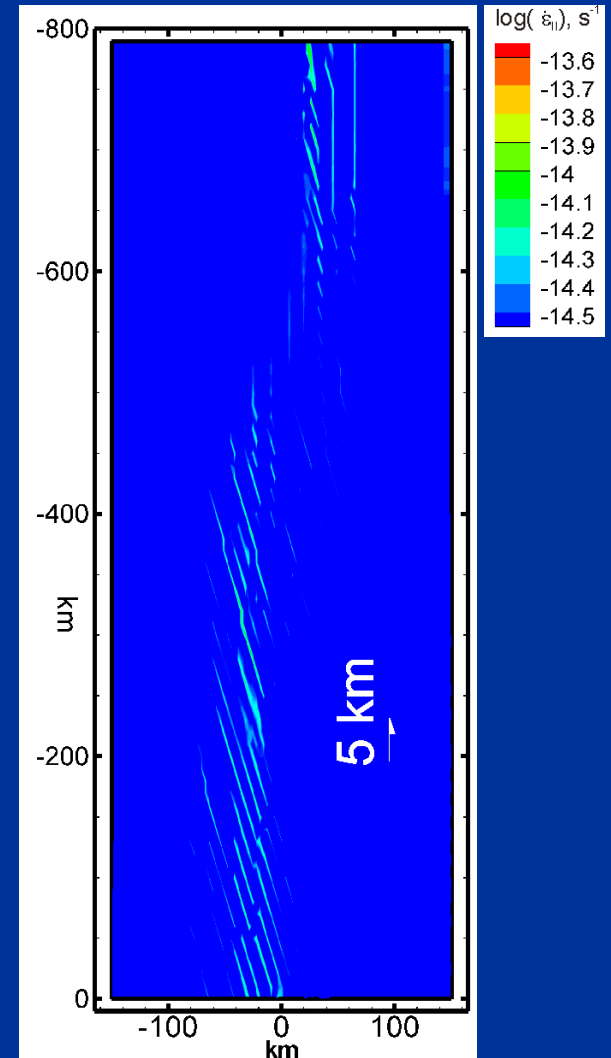
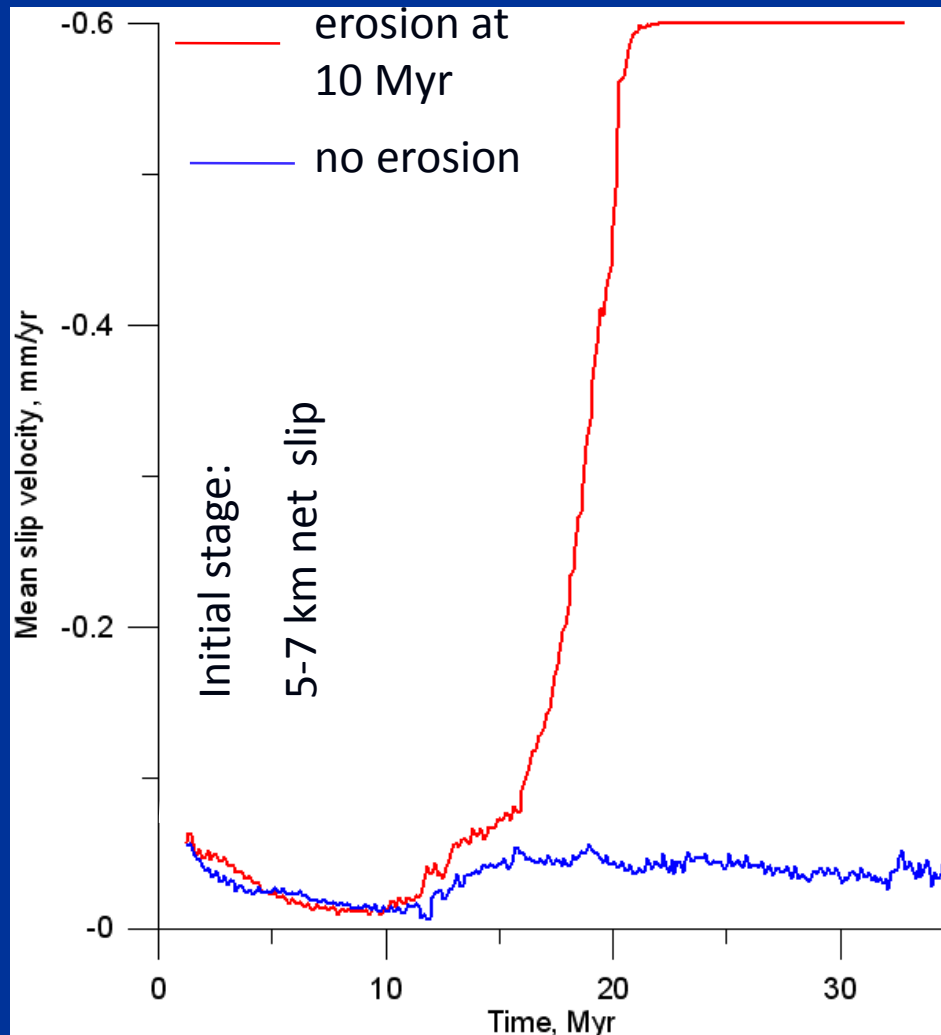


Flat Earth
approximation

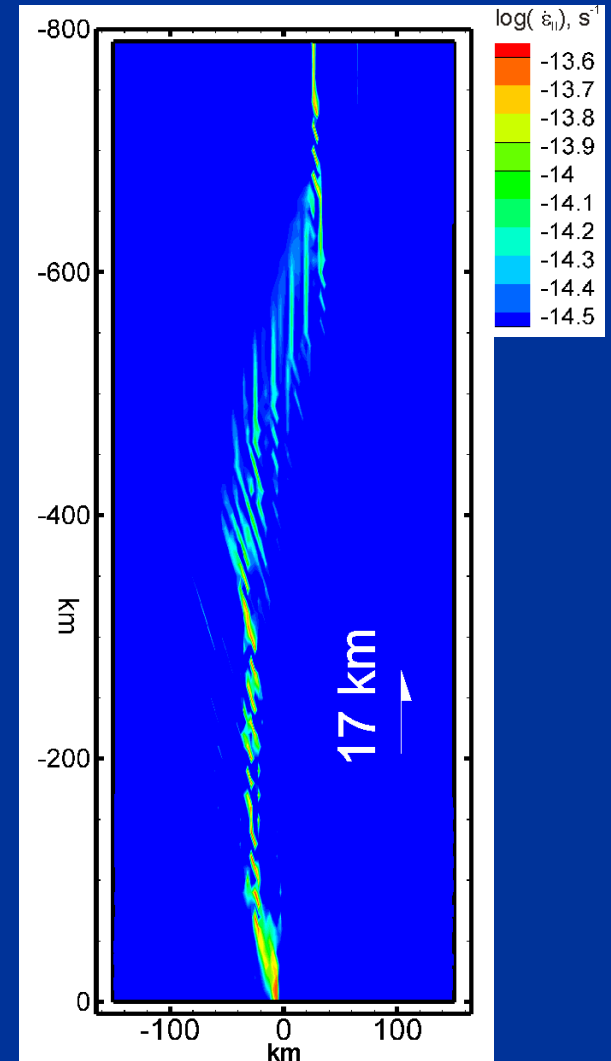
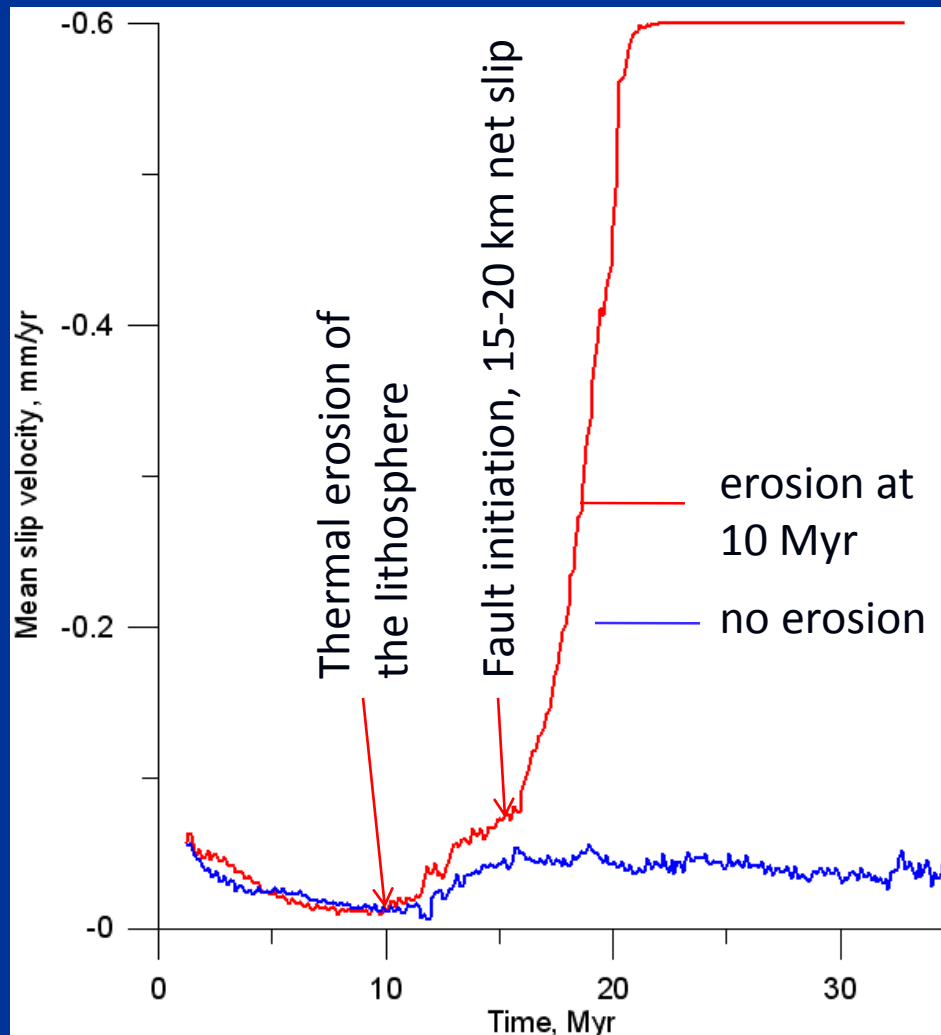


Modeling results: role of the thermal erosion of the lithosphere

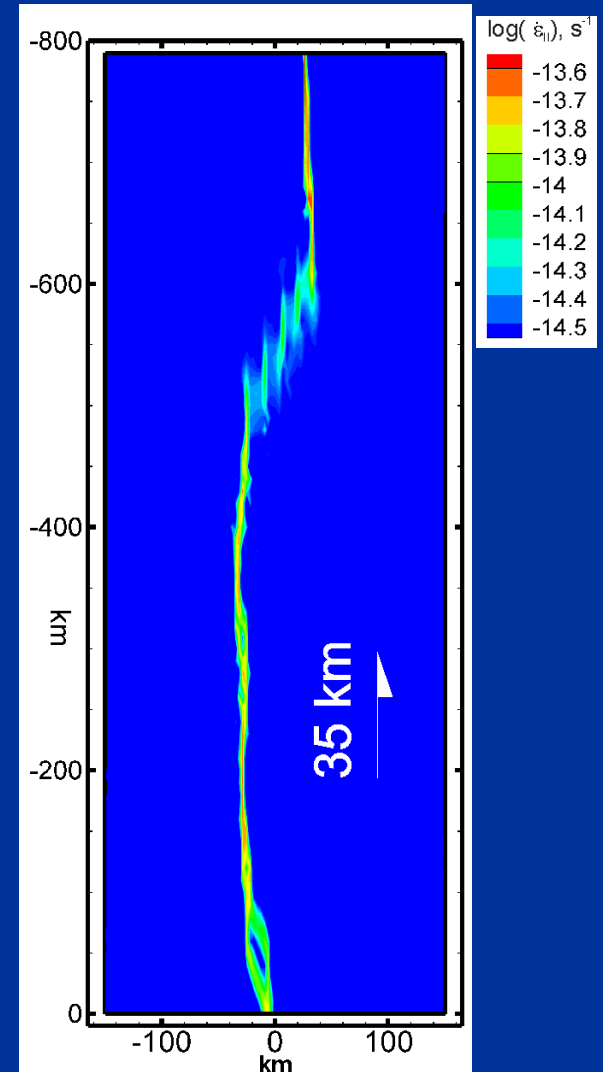
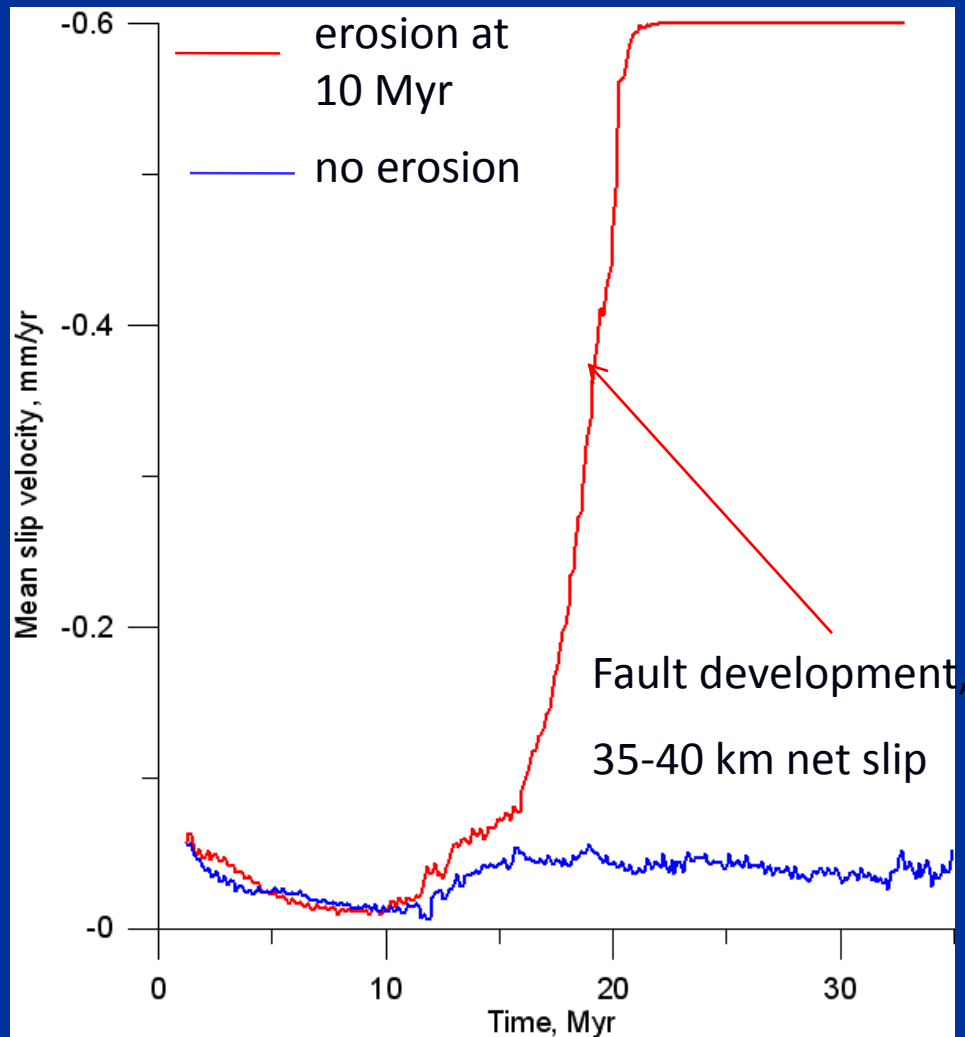
Applied force is $1.6 \times 10^{13} \text{ N/m}$



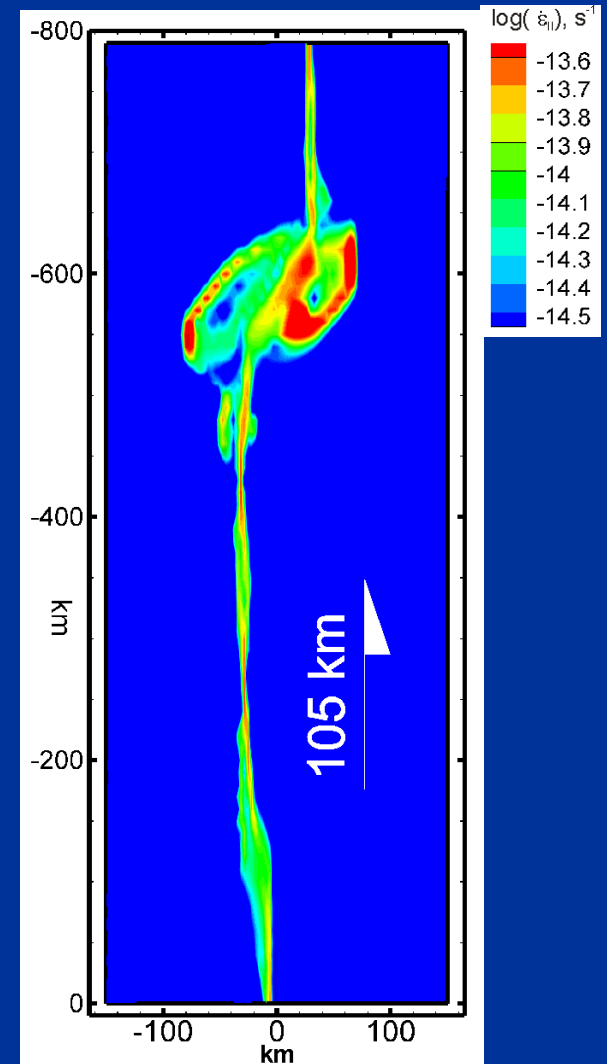
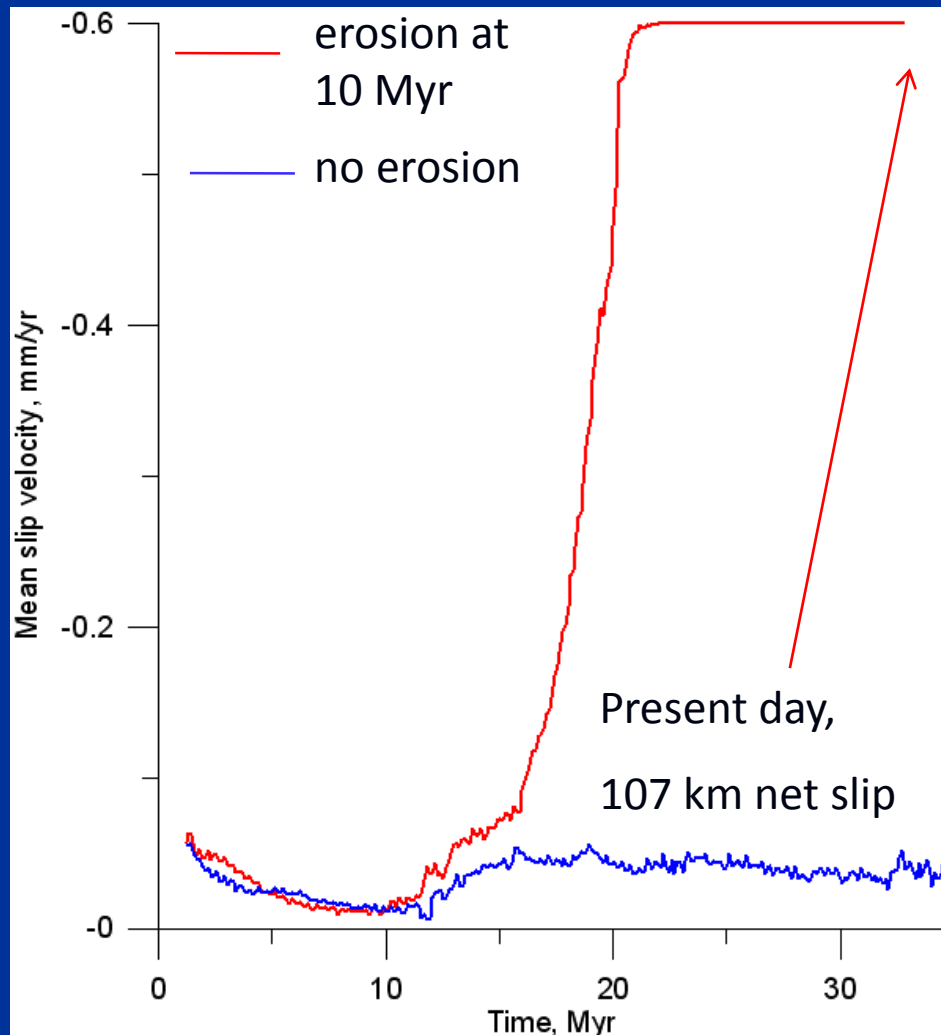
Modeling results: role of the thermal erosion of the lithosphere



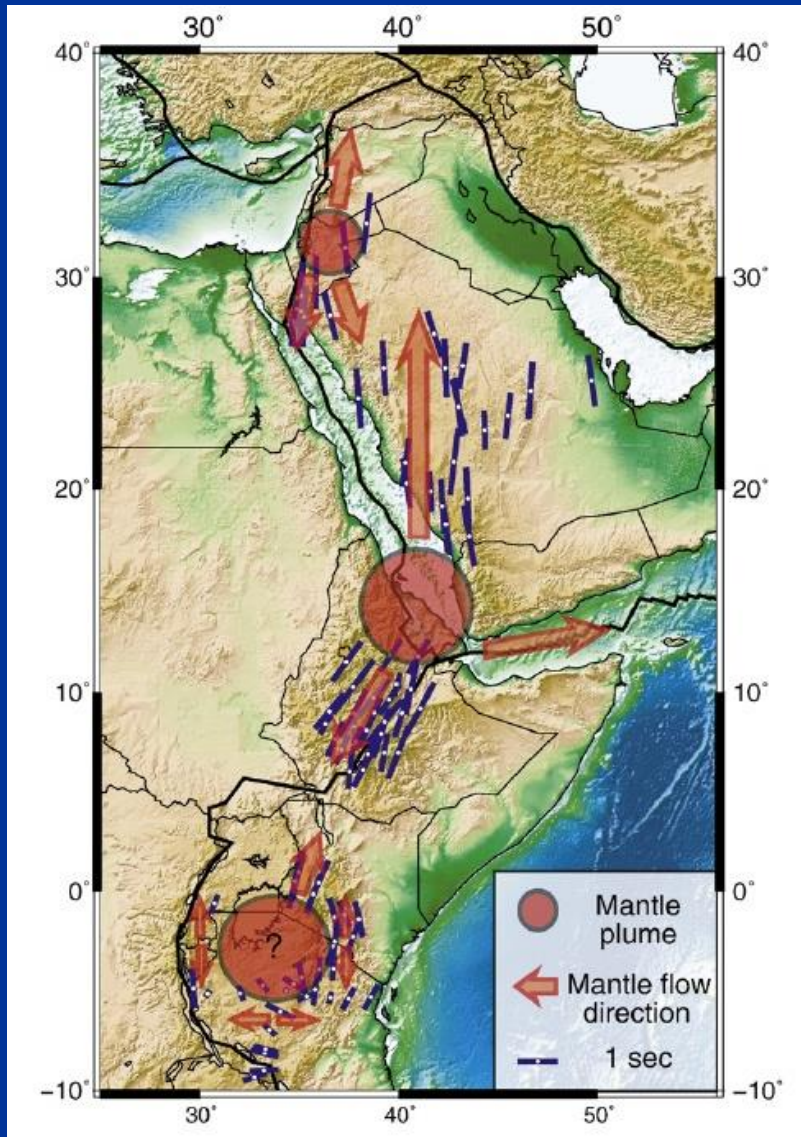
Modeling results: role of the thermal erosion of the lithosphere



Modeling results: role of the thermal erosion of the lithosphere



Possible scenario



Plumes at 25-35 Ma

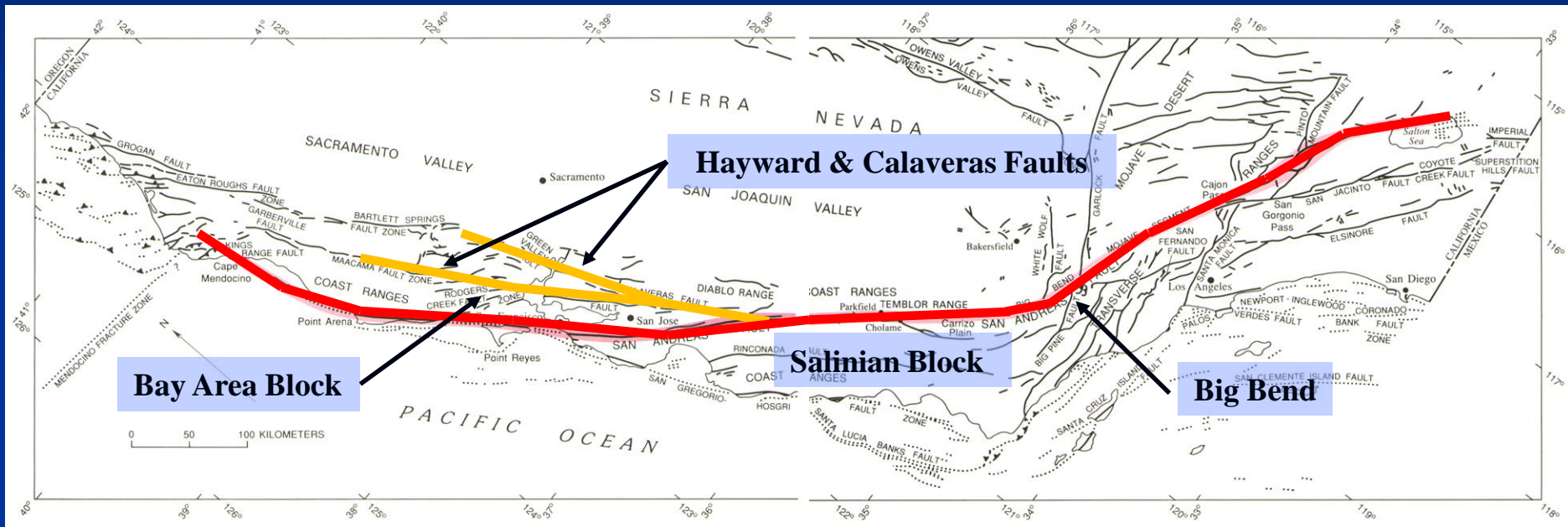
Lithospheric erosion 20-30 Ma

Localization of the DST
15-17 Ma

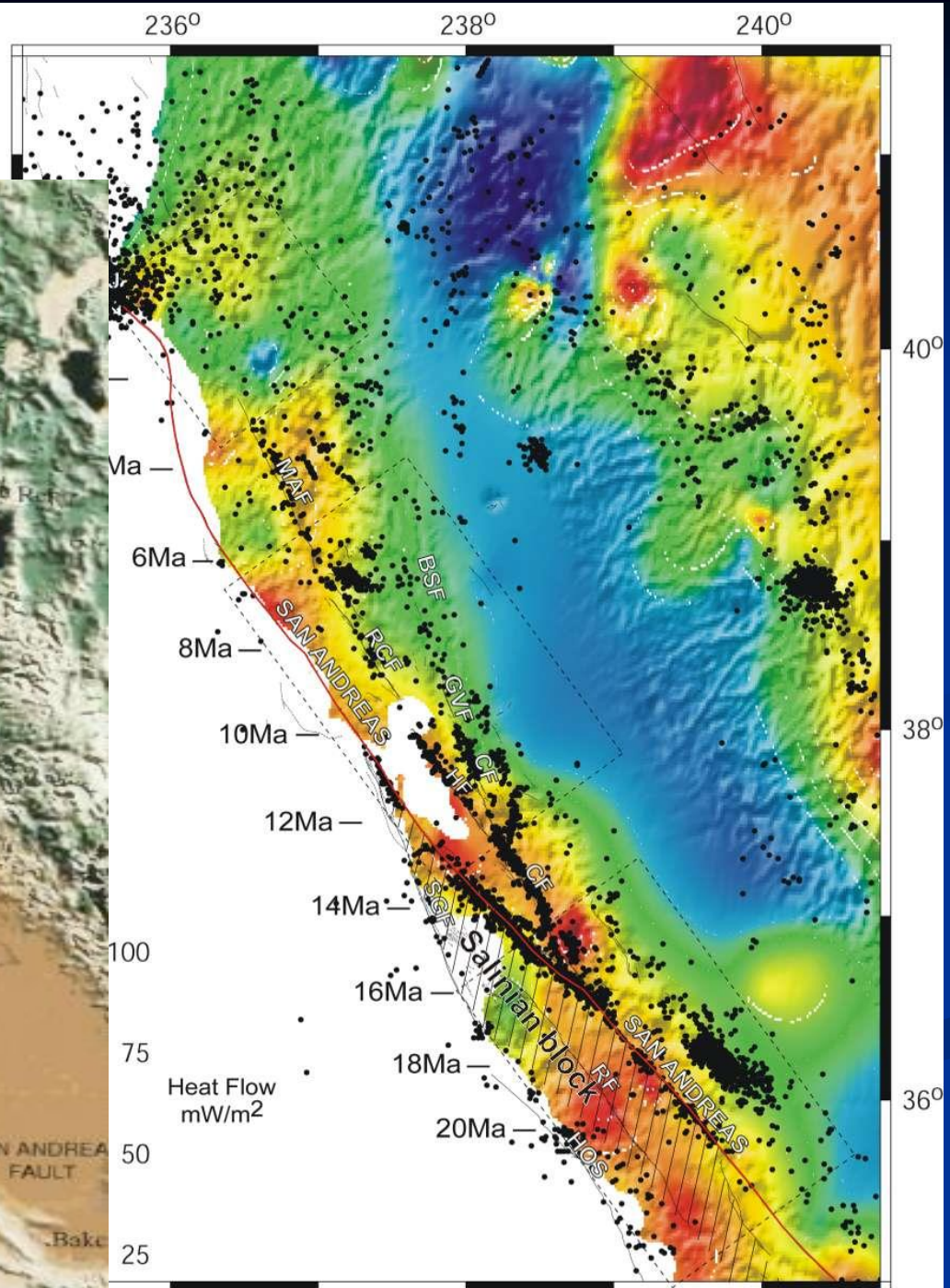
**Lithospheric erosion
has triggered the DST**

San Andreas Fault System

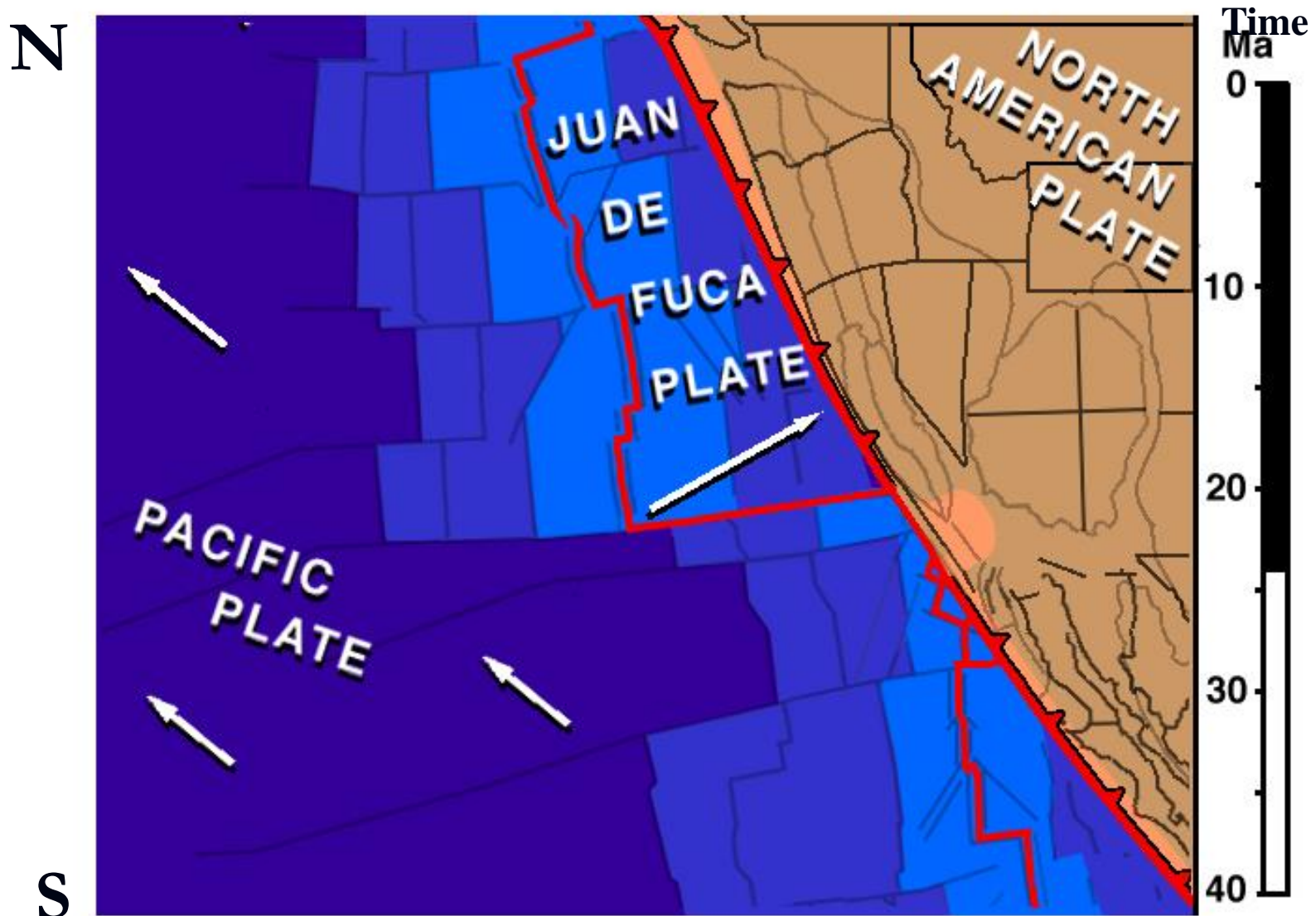
San Andreas Fault System



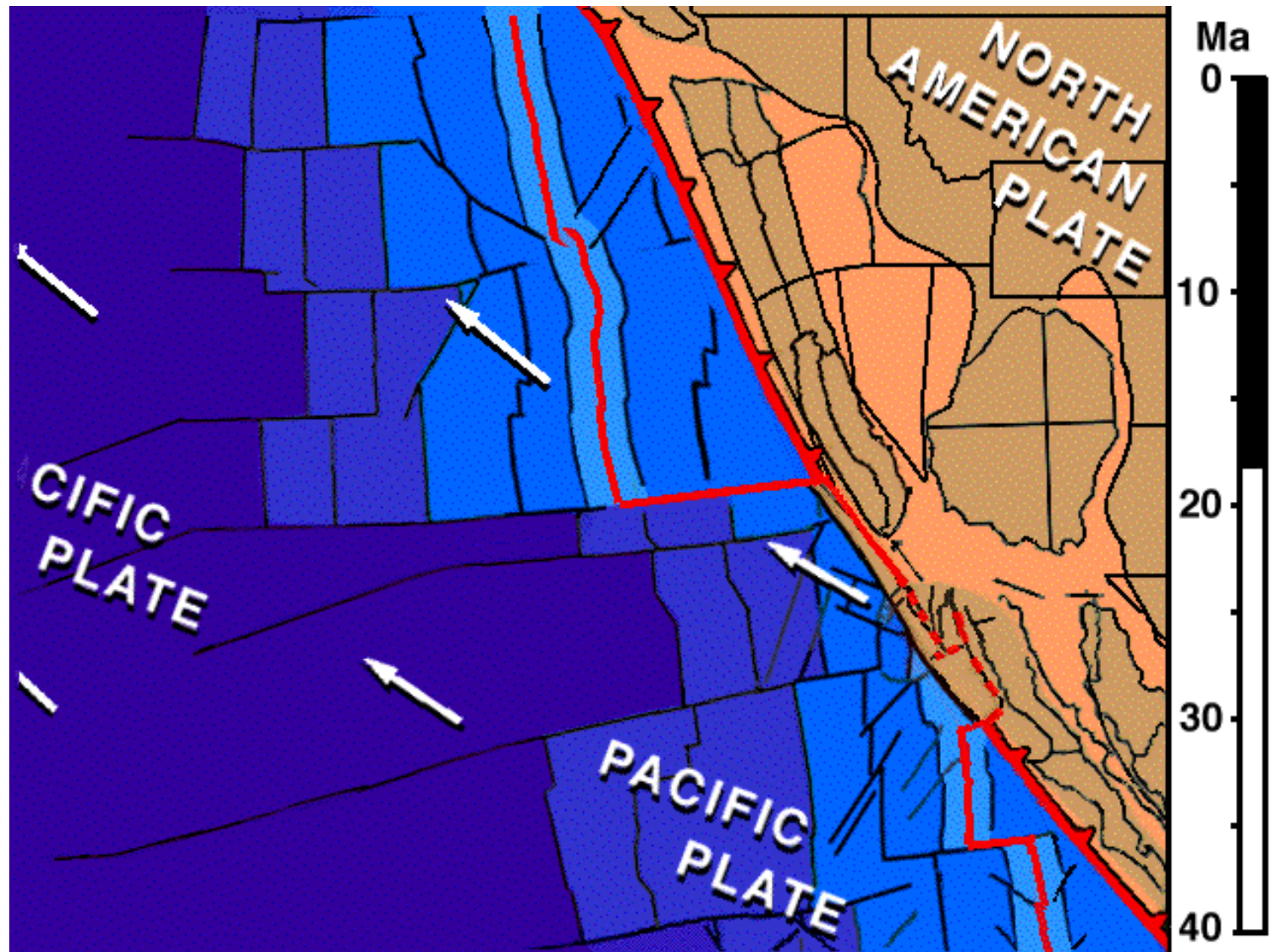
USGS Professional Paper 1515

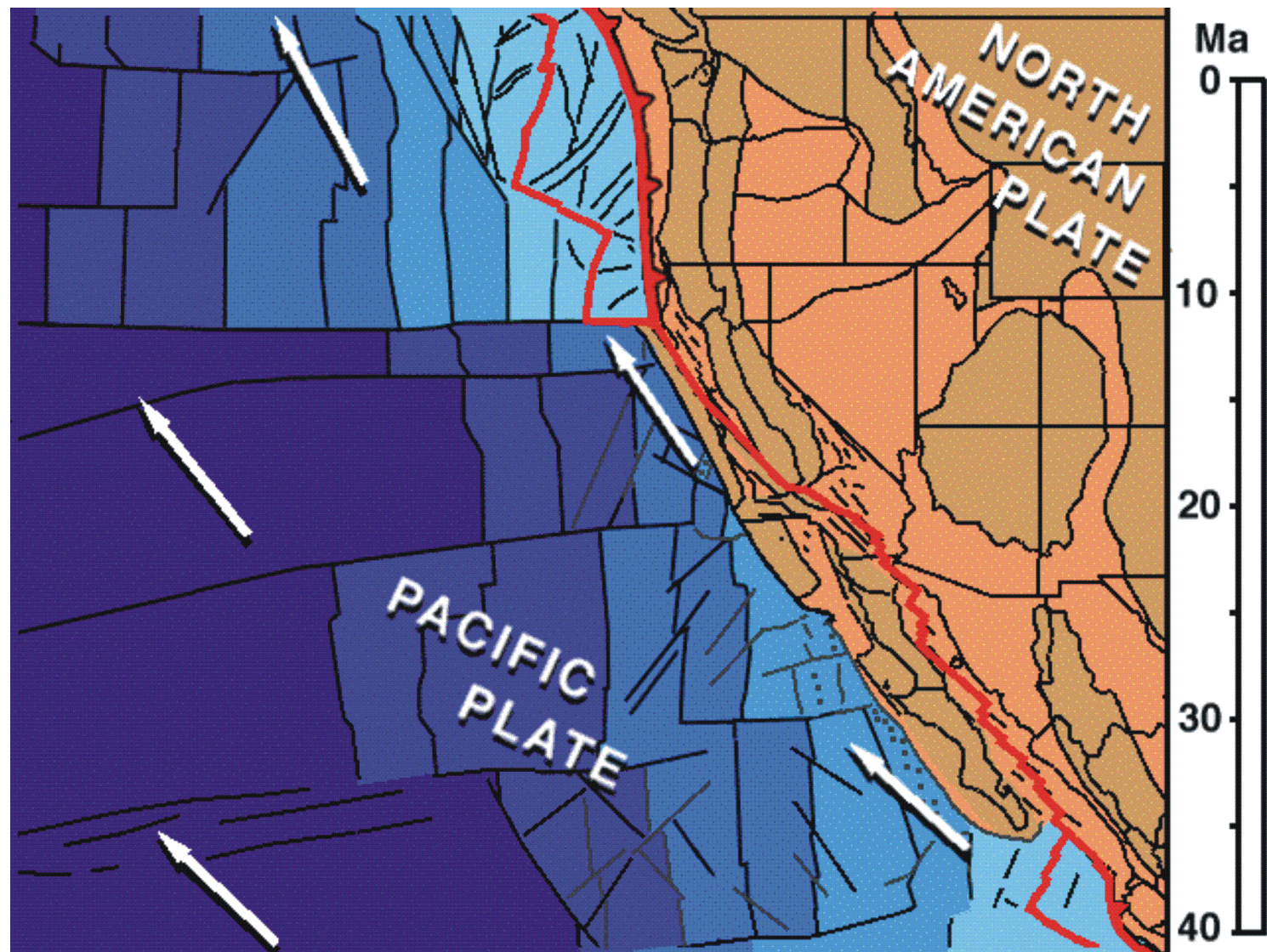


24 Ma: Shortly after Initiation of Strike-Slip



(animation by T. Atwater)





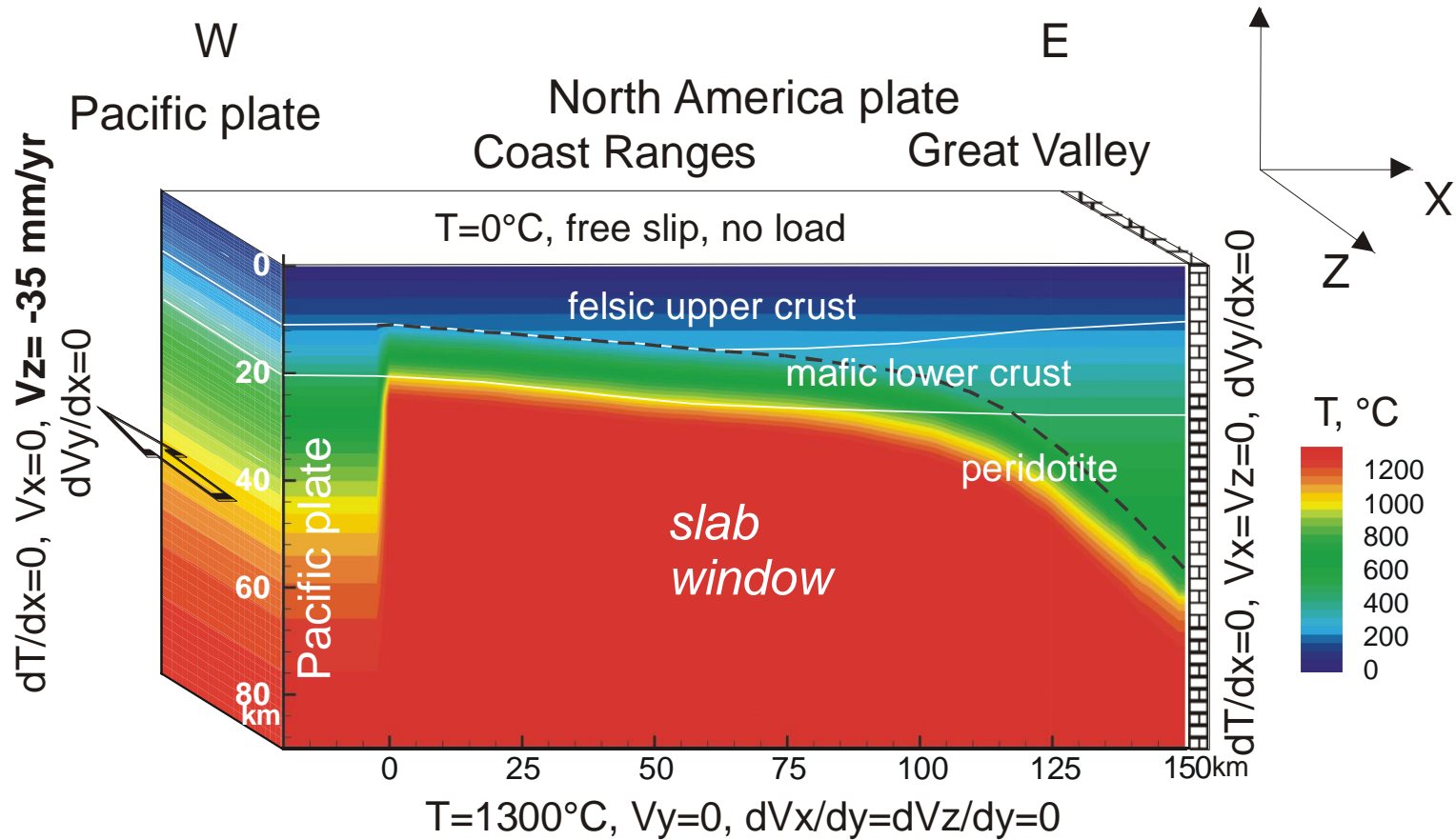
Questions addressed

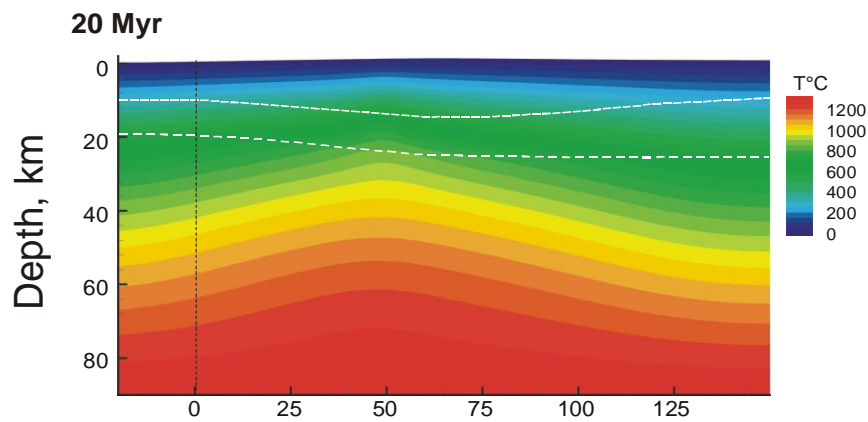
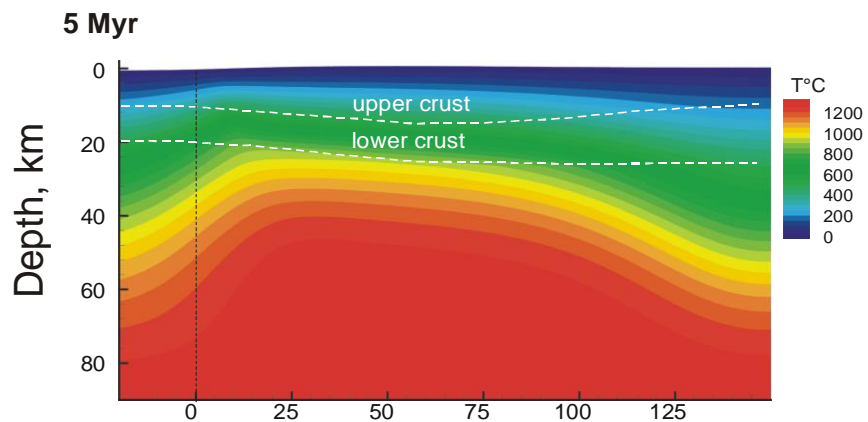
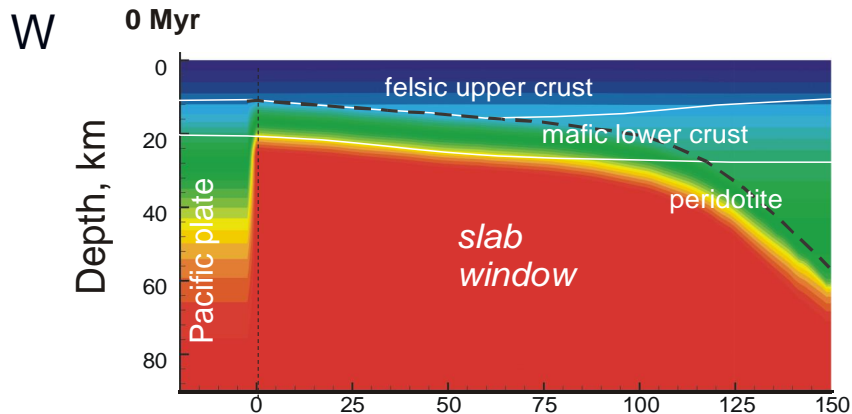
Why the locus of deformation in SAFS migrates landwards with time?

How differently would evolve SAFS with “strong” and “weak” major faults?

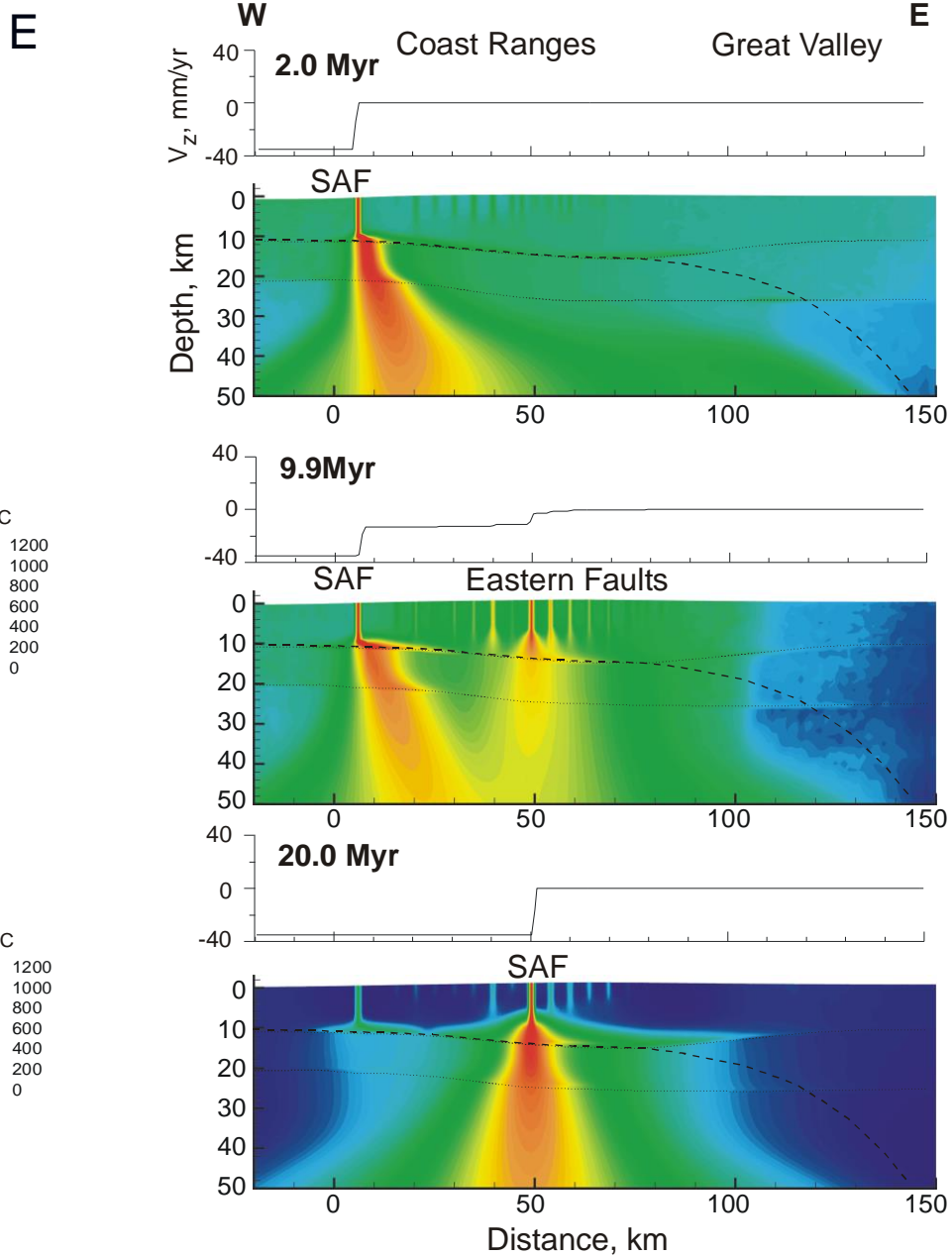
**Why the locus of deformation in SAFS migrates
landwards with time?**

Extended 2D Model Setup (South view)

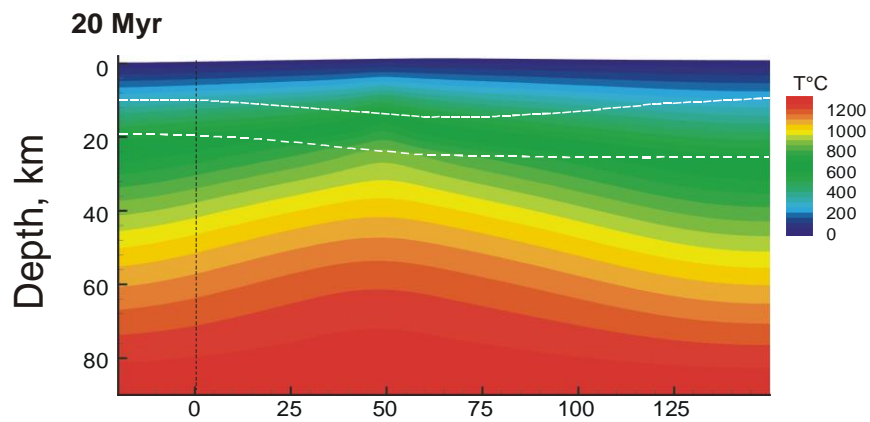
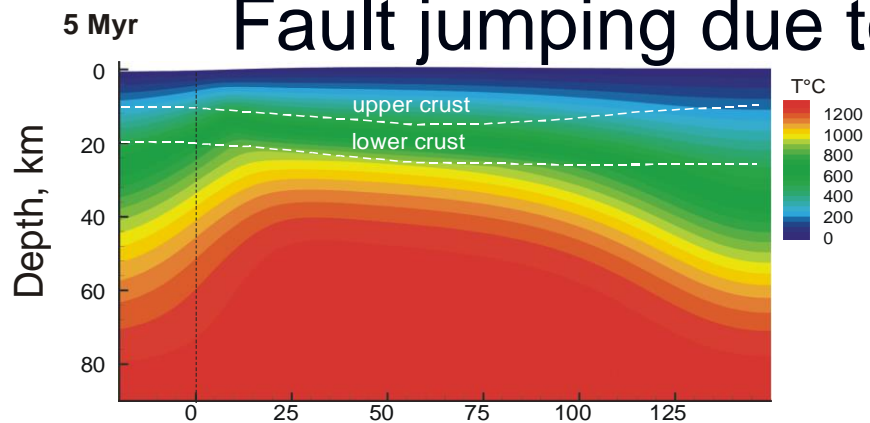
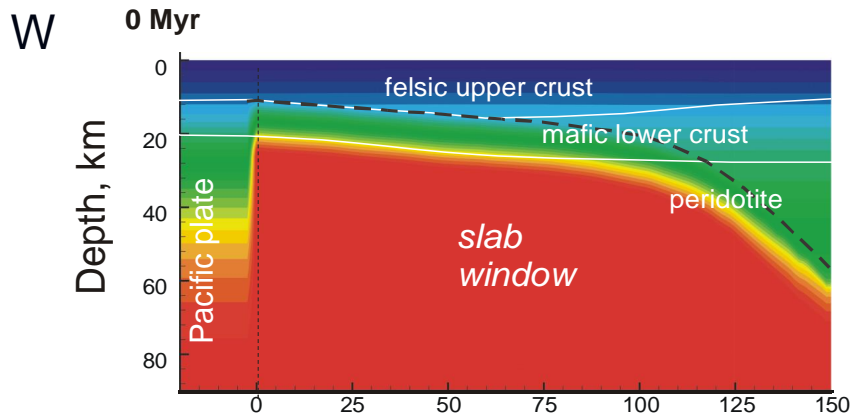




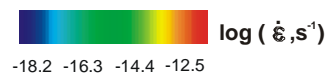
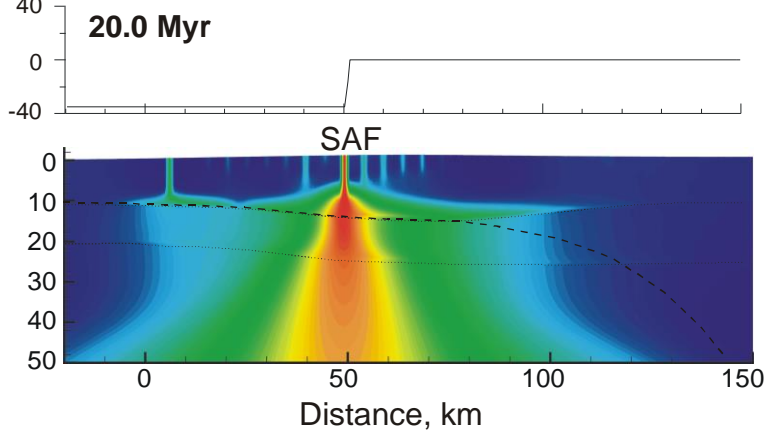
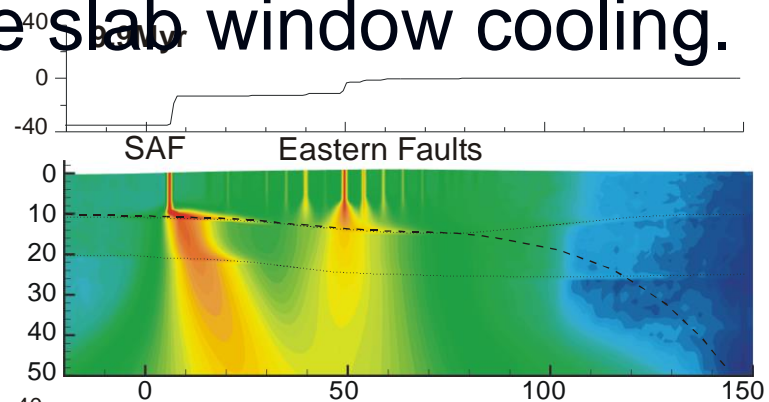
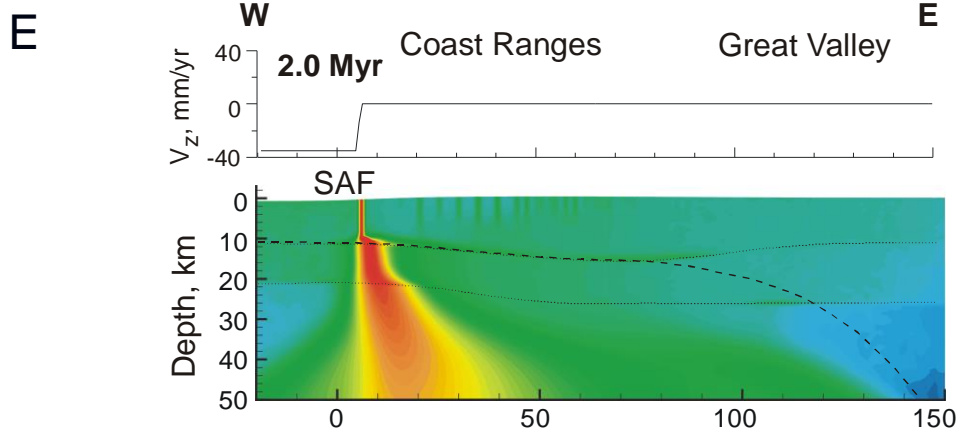
Distance, km



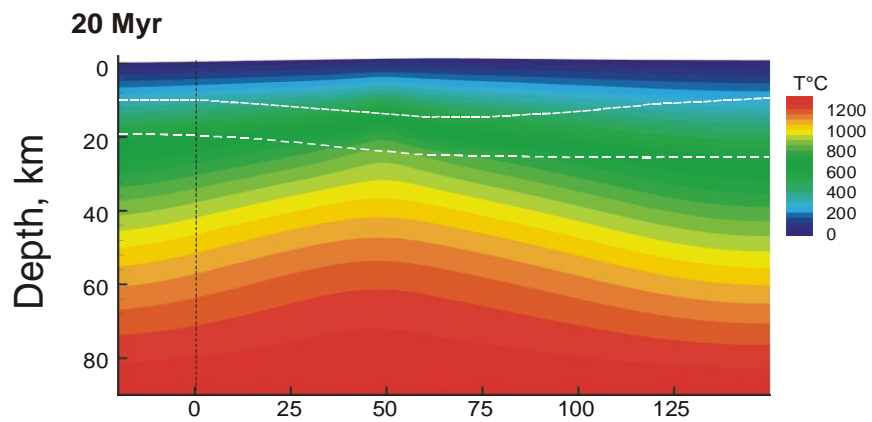
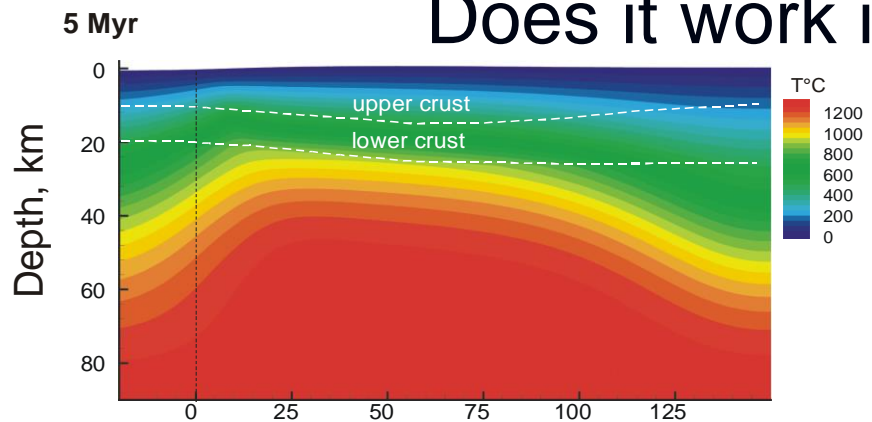
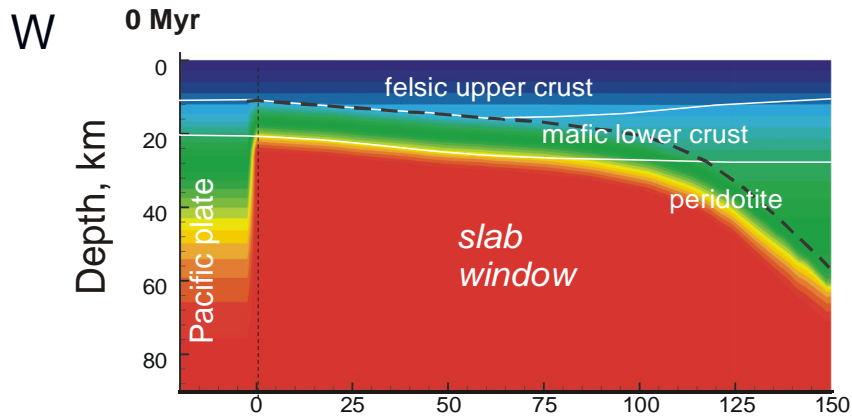
$\log(\dot{\epsilon}, s^{-1})$
-18.2 -16.3 -14.4 -12.5



Distance, km

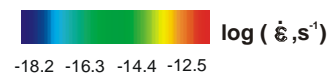
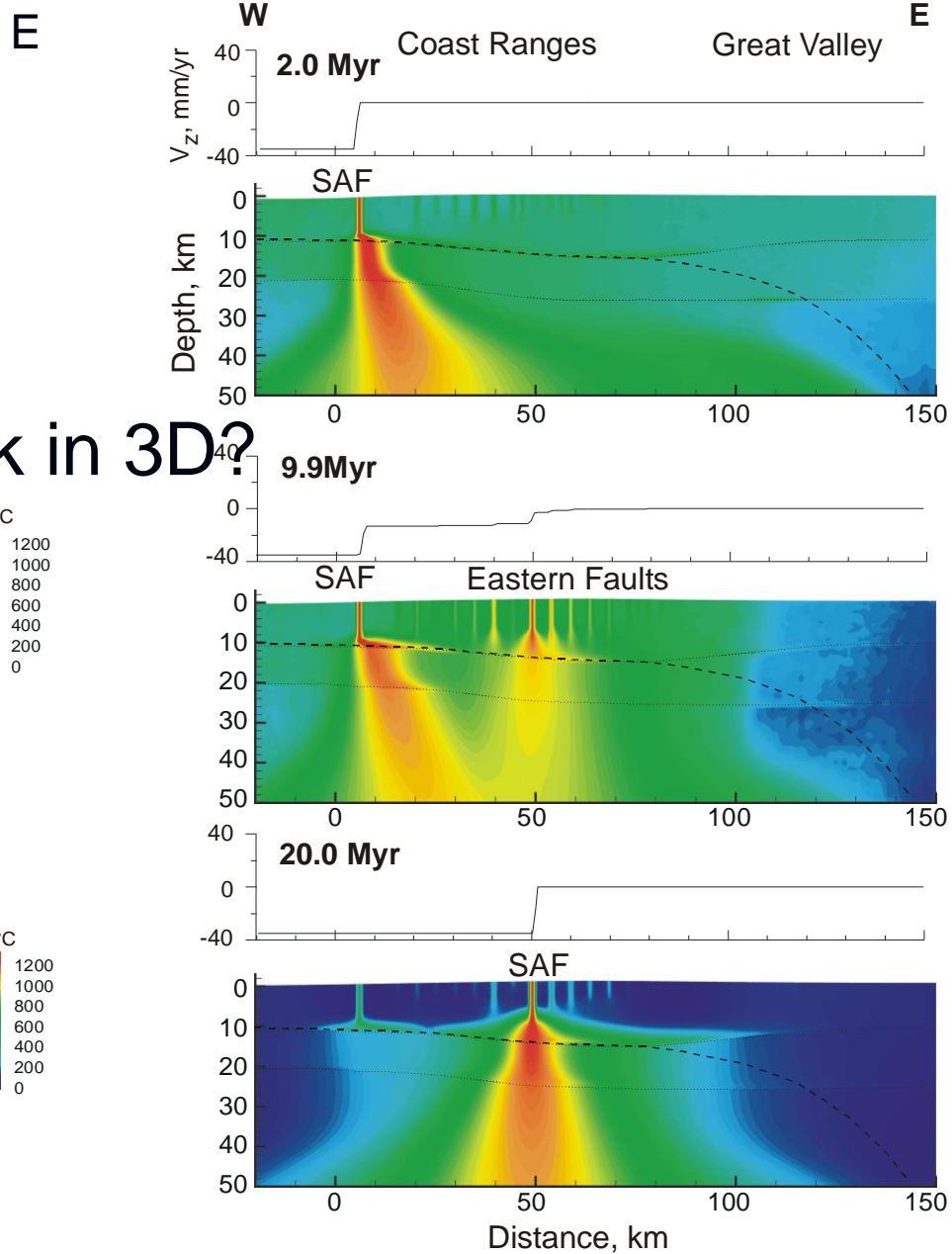


Fault jumping due to the slab window cooling.

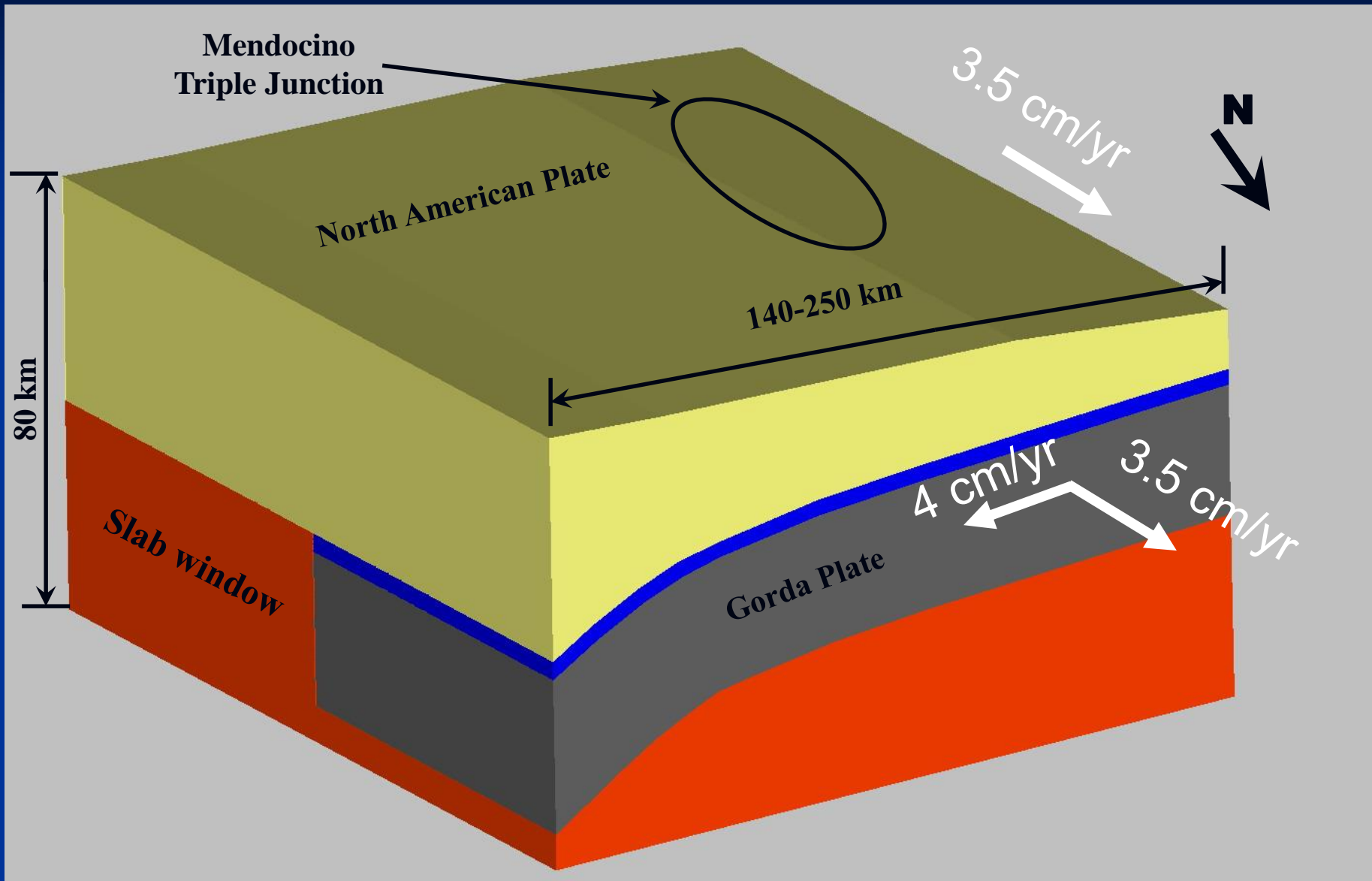


Distance, km

Does it work in 3D?



3D Model Setup (view from the North)

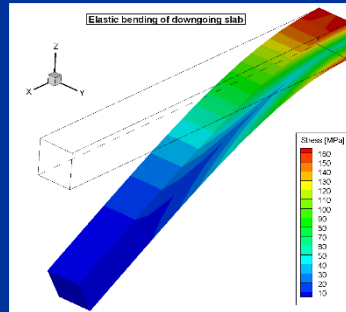


Physical background

Balance equations

Momentum:
$$\frac{\partial \sigma_{ij}}{\partial x_j} + \Delta \rho g z_i = 0$$

Energy:
$$\frac{DU}{Dt} = -\frac{\partial q_i}{\partial x_i} + r$$



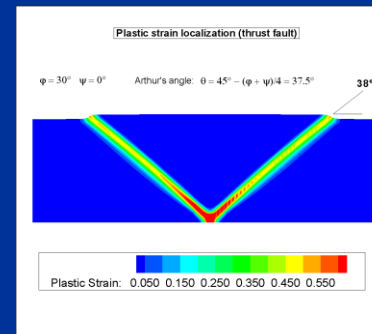
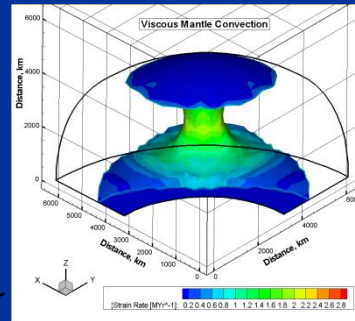
Deformation mechanisms

$$\dot{\epsilon}_{ij} = \dot{\epsilon}_{ij}^{el} + \dot{\epsilon}_{ij}^{vs} + \dot{\epsilon}_{ij}^{pl}$$

Elastic strain:
$$\dot{\epsilon}_{ij}^{el} = \frac{1}{2G} \hat{\tau}_{ij}$$

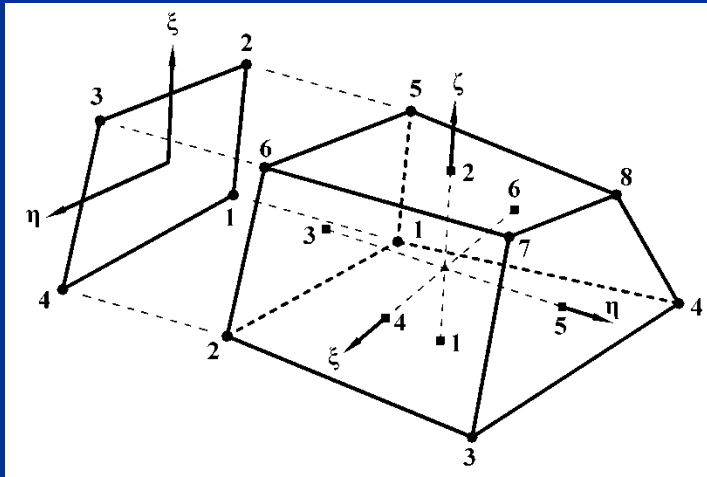
Viscous strain:
$$\dot{\epsilon}_{ij}^{vs} = \frac{1}{2\eta_{eff}} \tau_{ij}$$

Plastic strain:
Mohr-Coulomb
$$\dot{\epsilon}_{ij}^{pl} = \dot{\gamma} \frac{\partial Q}{\partial \tau_{ij}}$$



Numerical background

Discretization by Finite Element Method



Fast implicit time stepping
+ Newton-Raphson solver

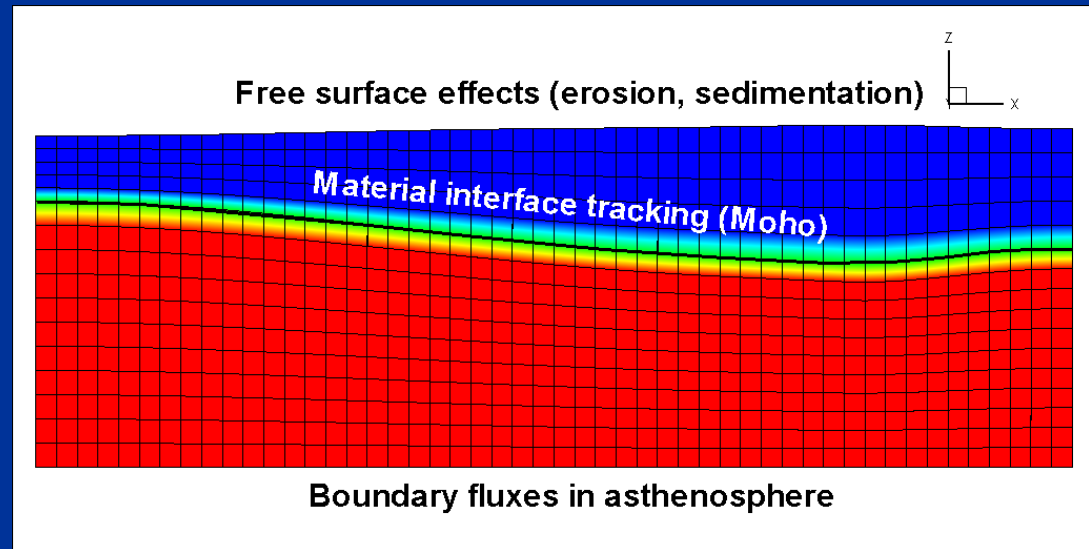
$$\mathbf{u}_{k+1} = \mathbf{u}_k - \mathbf{K}_k^{-1} \mathbf{r}_k$$

\mathbf{r} – Residual Vector

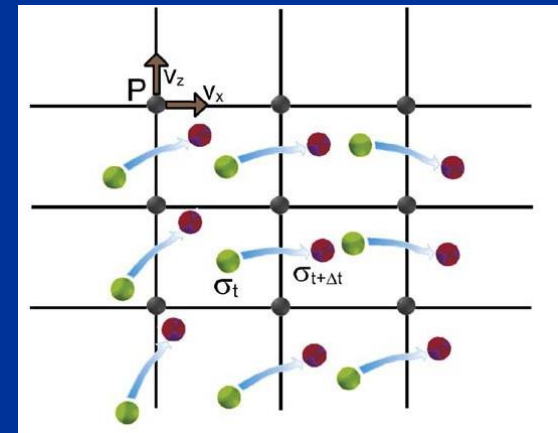
$$\mathbf{K} = \frac{\partial \mathbf{r}}{\partial \Delta \mathbf{u}} \quad \text{– Tangent Matrix}$$

Popov and Sobolev (2008)

Arbitrary Lagrangian-Eulerian kinematical formulation



Remapping of
entire fields by
Particle-In-Cell
technique

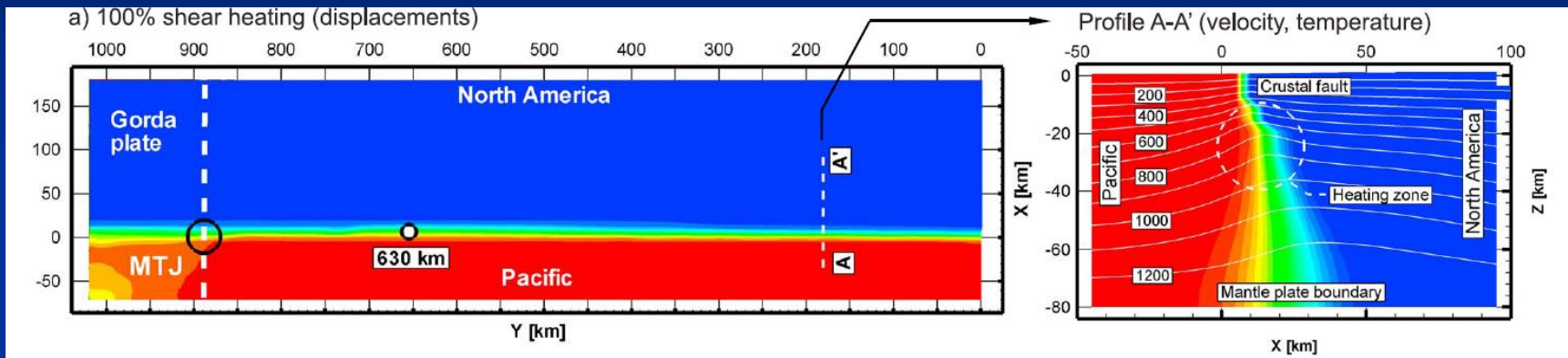


“Strong” and “weak” faults models

“Strong faults” model: the friction coefficient decreases only slightly (from 0.6 to 0.3) with increasing plastic strain

“Weak faults” model: the friction coefficient decreases drastically (from 0.6 to 0.07) with increasing plastic strain

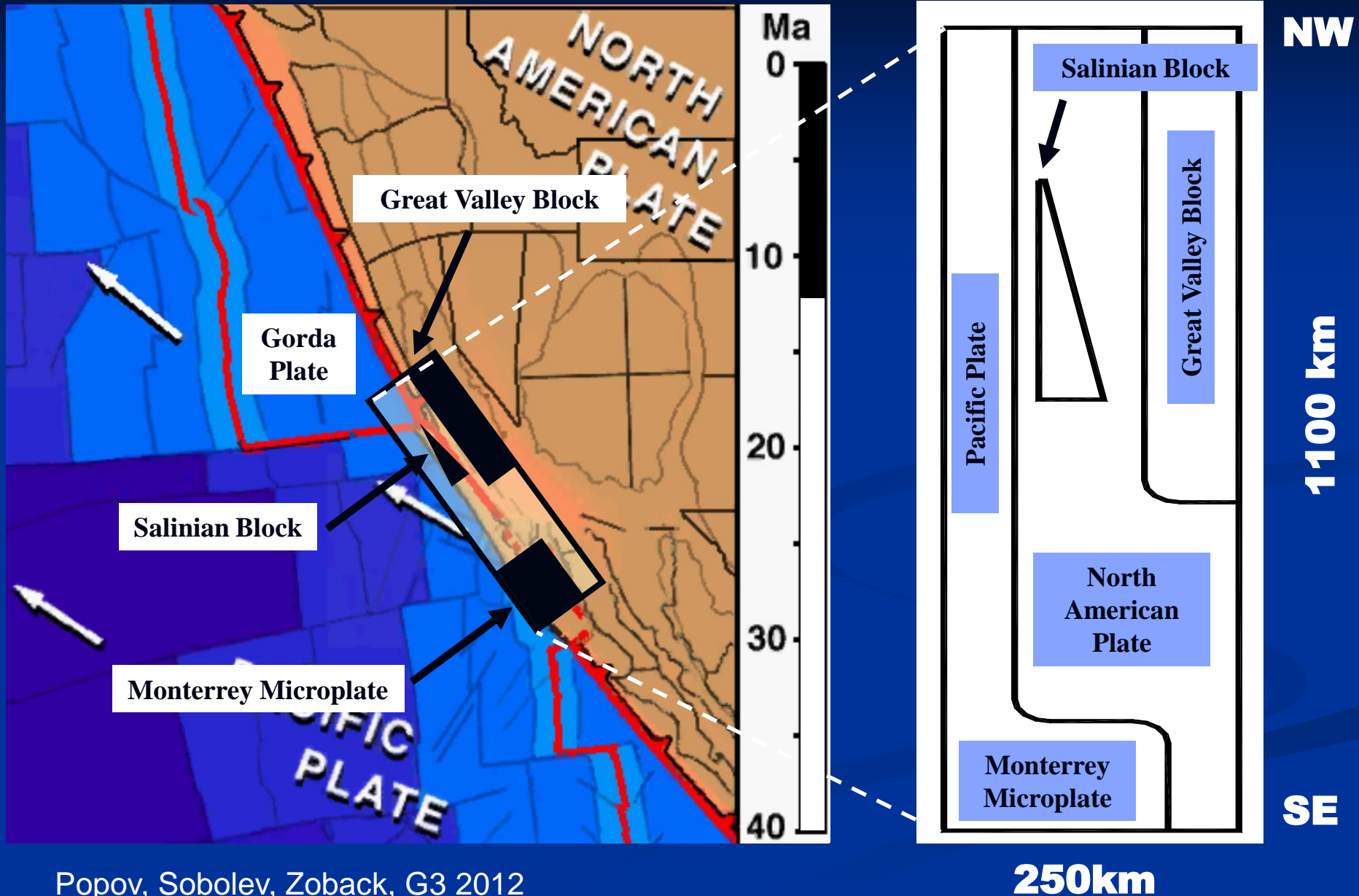
3D Model (slab window cooling)



In 3D fault doesn't jump due to the slab window cooling. The reason is along-strike mechanical interaction (transpression) inhibiting fault jumping.

Therefore new explanation is required

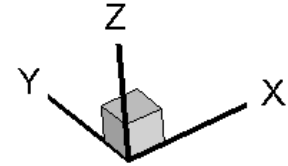
3D Model Setup with heterogeneity (at 12-15 MA)



NW

Effective strain rate [s-1]

Time: 12.00[MA]

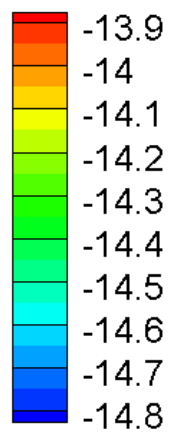


North American Plate

Pacific Plate

100km scale

Strain Rate



Higher strain rate ↑

Model depth 80km

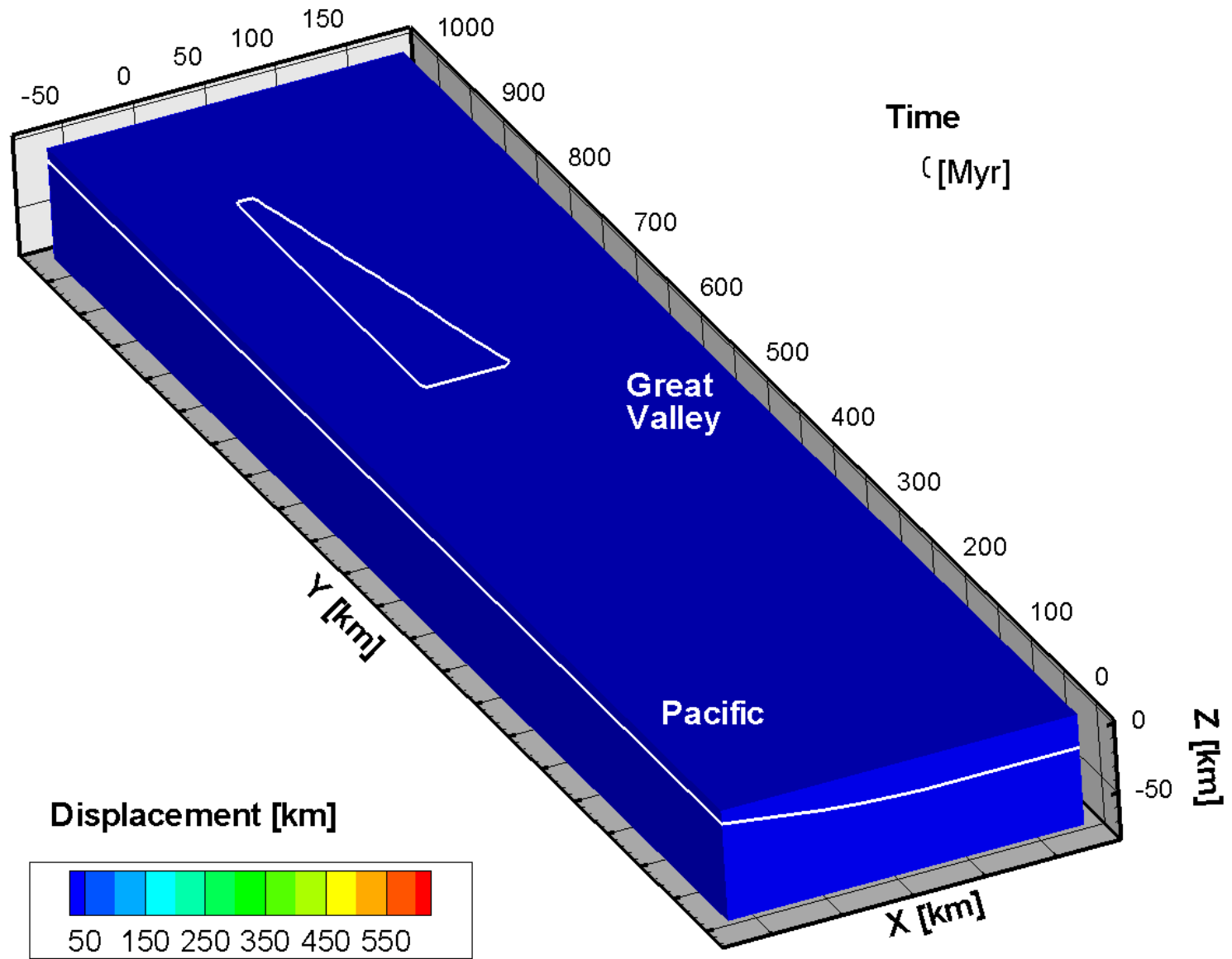
White line shows Salinian block

Subducting Gorda Slab on the NW part

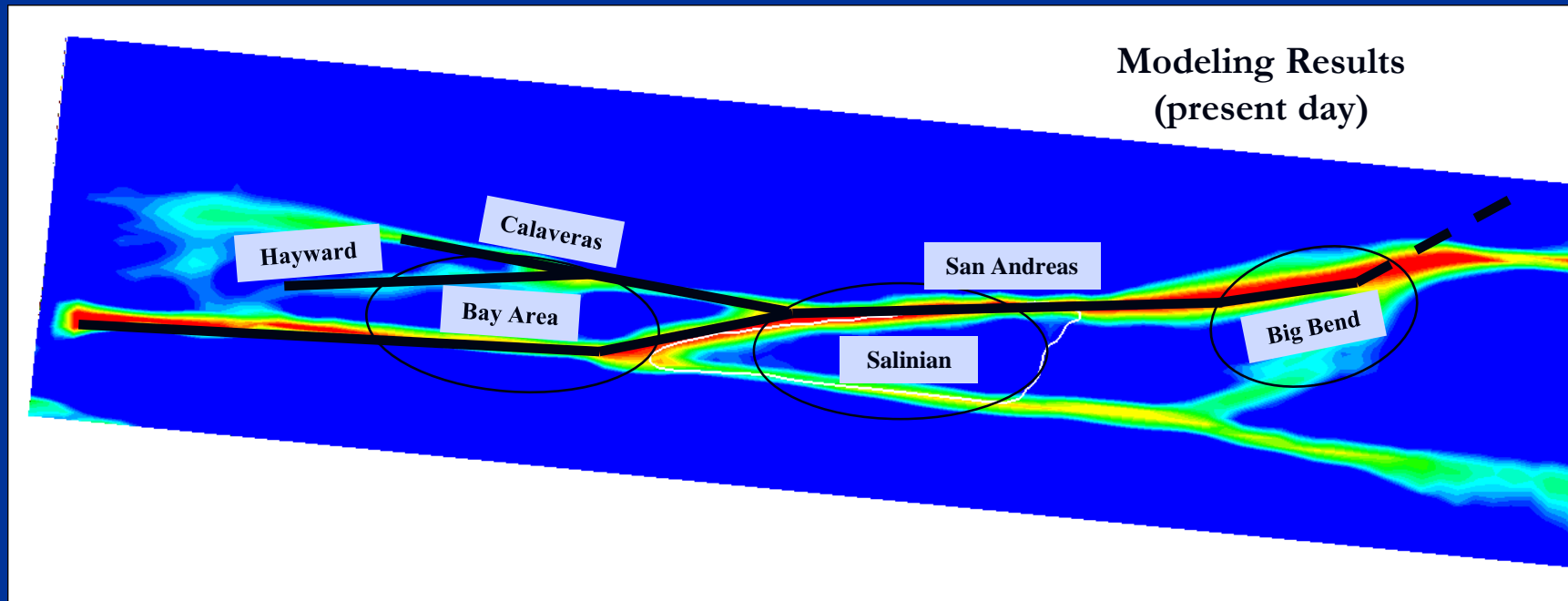
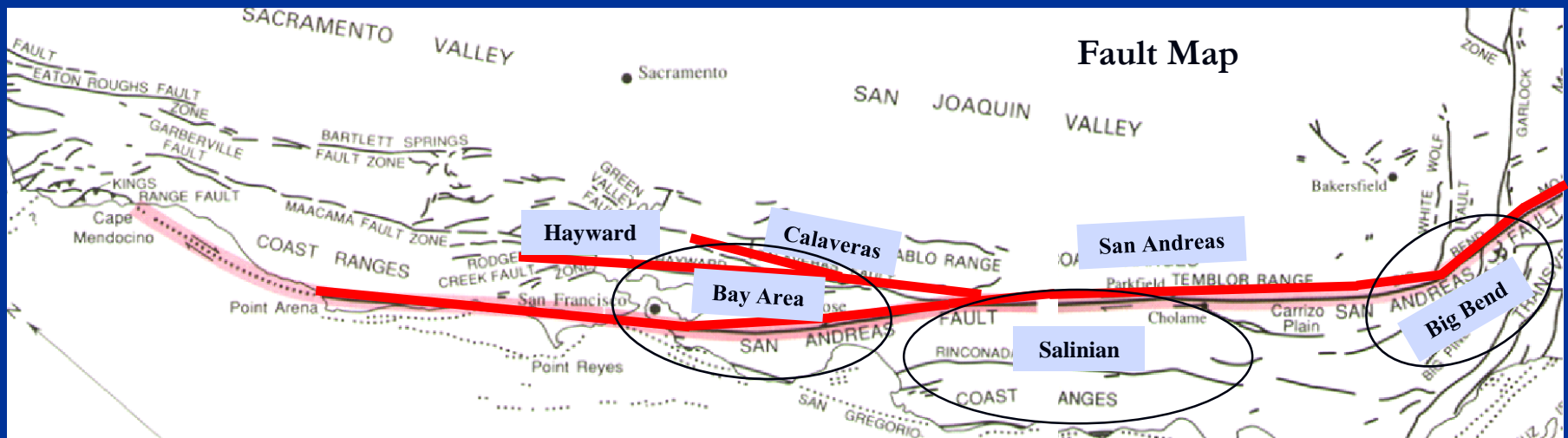
Strike-slip velocity 3.5 cm/yr is applied to Pacific Plate

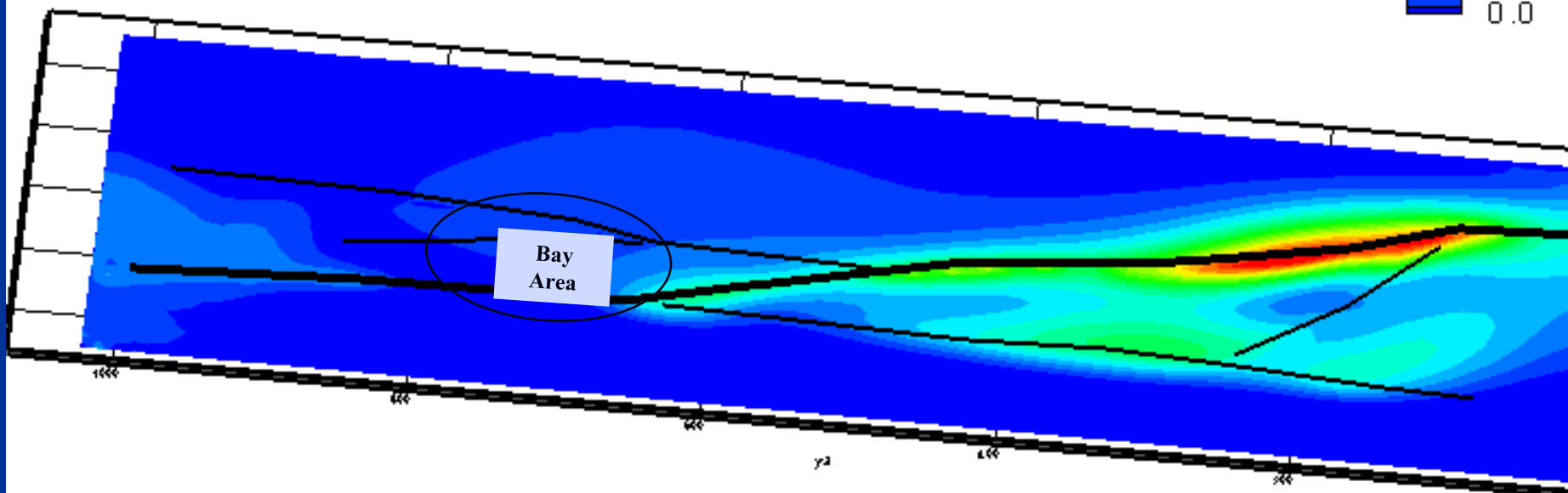
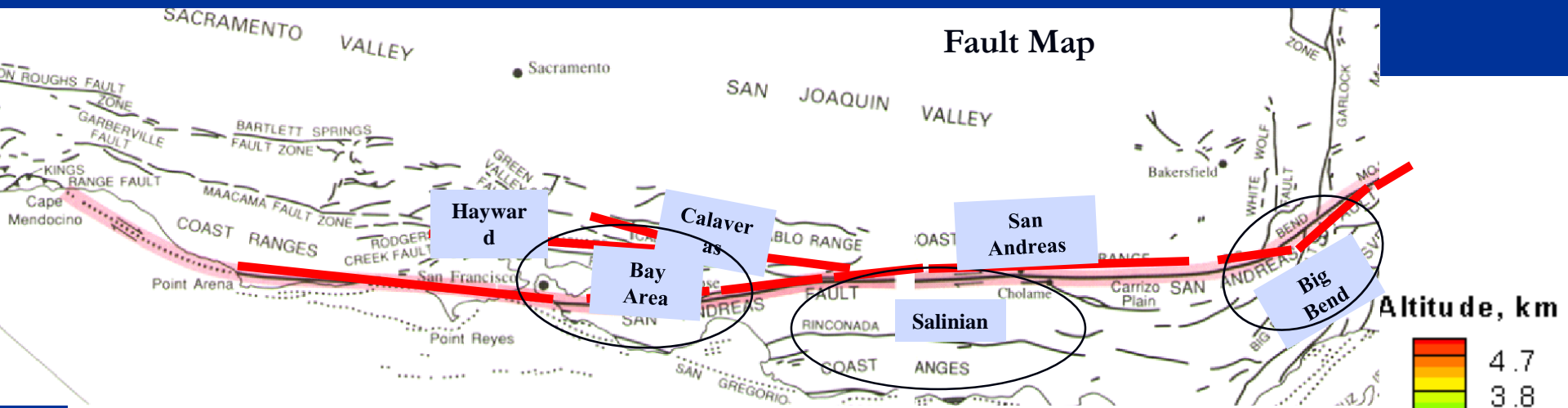
Output in fixed Pacific Plate reference frame

SE

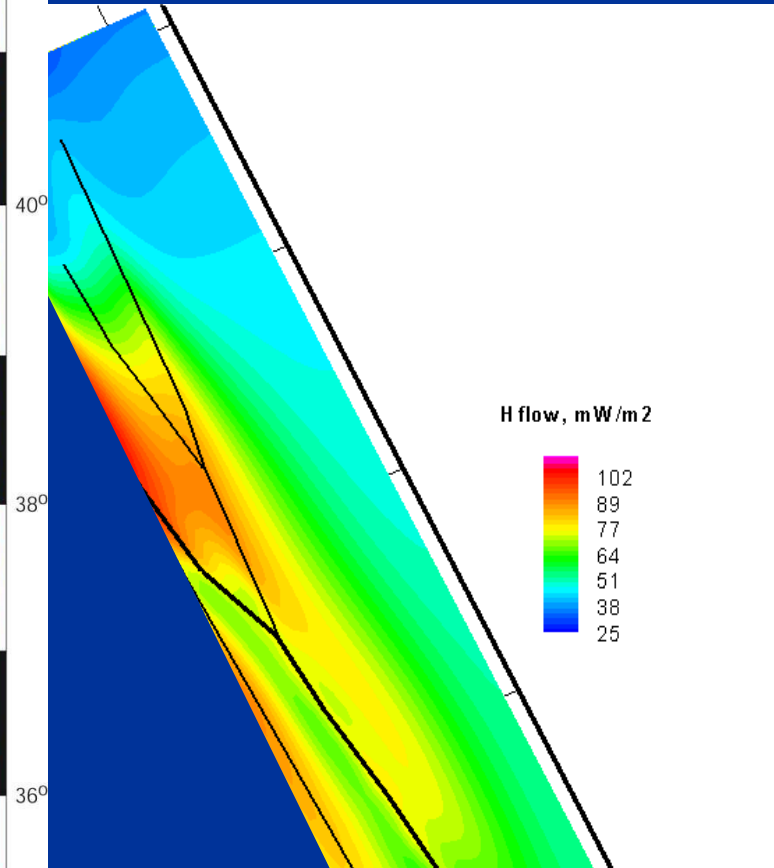
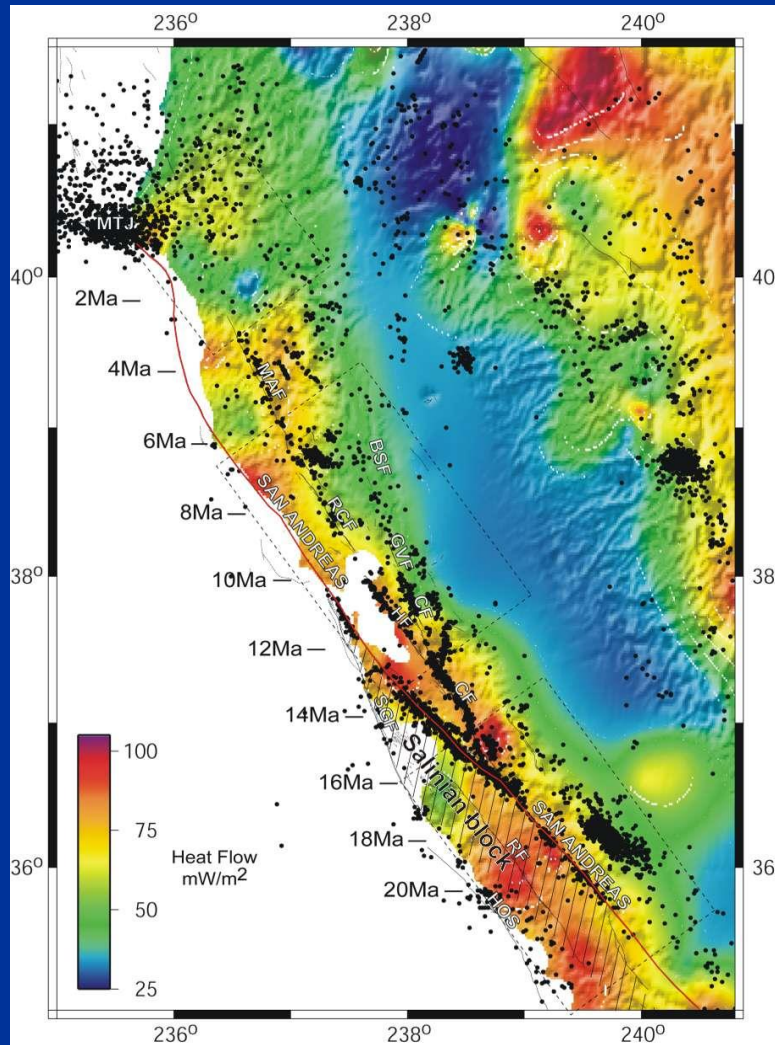


Qualitative comparison of basic fault features



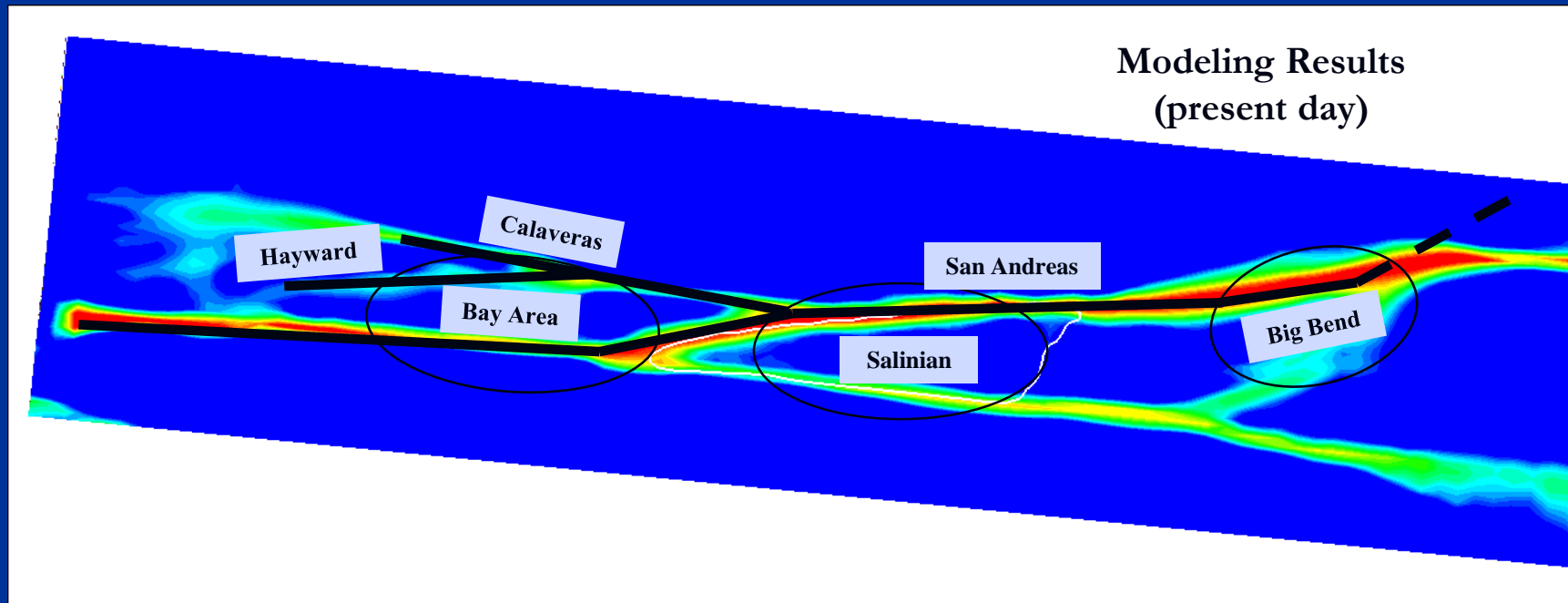


Modeled surface heat flow

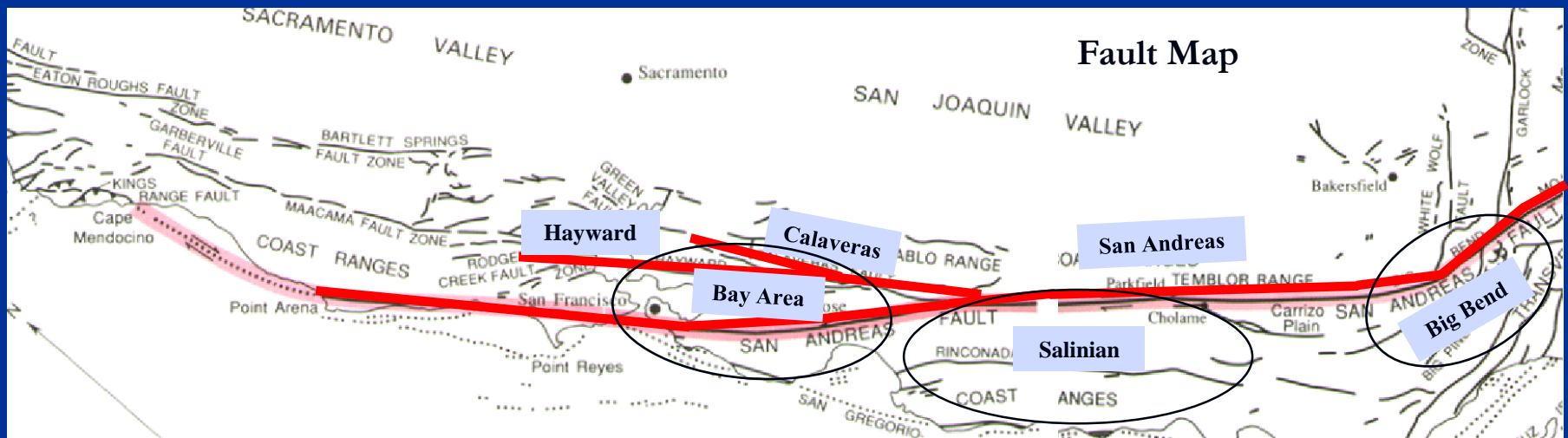


“Weak faults” versus “strong faults” model

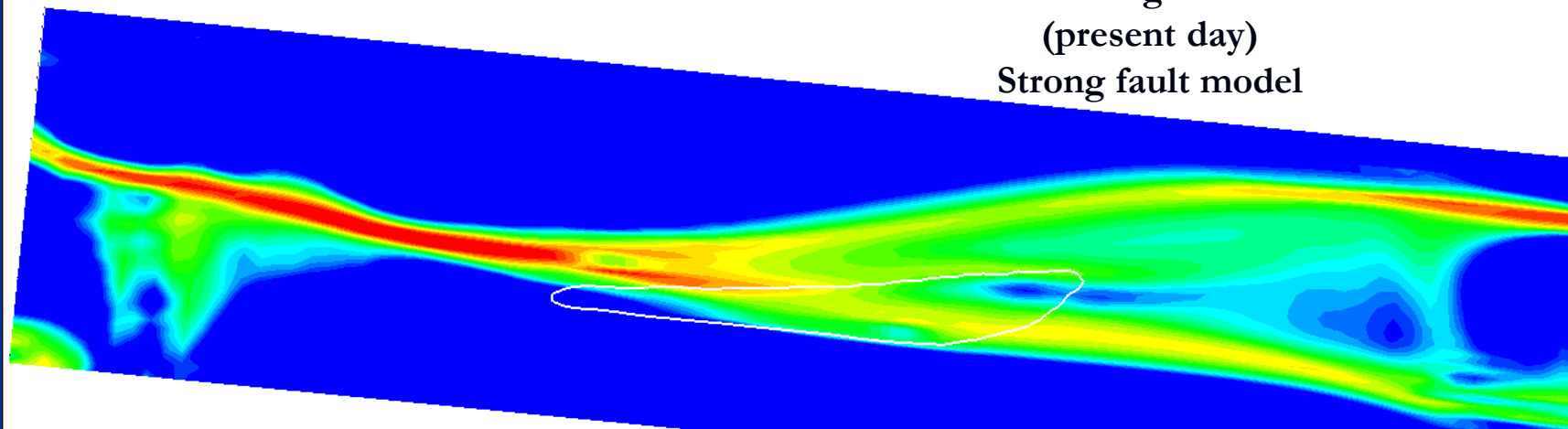
Weak-faults model



Strong-faults model



Modeling Results
(present day)
Strong fault model



Major faults in SAFS must be weak!

Conclusions for SAFS

Present day structure and landward motion of SAFS is controlled by kinematic boundary conditions and lithospheric heterogeneity, including captured Monterray microplate

Major faults at SAFS must be “weak”